



CENTRE NATIONAL D'ÉTUDES SPATIALES

IASI-NG Status of development and preparatory activities by the mission team

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1- IASI-NG Mission rationale and Objectives

2- IASI-NG Mission requirements

3- The french mission group : MENINGE

4- IASI-NG feasibility studies

5- Consequences of the successful IASI-NG feasibility studies

6- IASI-NG Schedule

Expressed in the Position papers collected by Eumetsat

In France proposed by scientists to CNES AO for new missions in 2009

■ **Continuity of IASI.**

- **Justified by Operational use of NWP centers, and other emerging services (MACC, VAAC), and long term trends monitoring as well.**

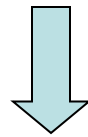
⇒ **Same performances as IASI used as thresholds for follow-on**

⇒ **Same viewing geometry required**

■ **Improved performances to tackle issues unsolved with IASI.**

- **More precise humidity profiles**
- **More documented atmospheric layers near surface (BL) , or around tropopause (UTLS).**
- **Better use of information in cloudy conditions**
- **More detailed profiles of minor components**
- **Detect other species**

EUM : operational meteorology
F : AC and climate



Balance between these applications

Priority has been given to

- **Continuity (full spectrum, + imager)**
- **Improved performances in radiometry and spectral resolution**

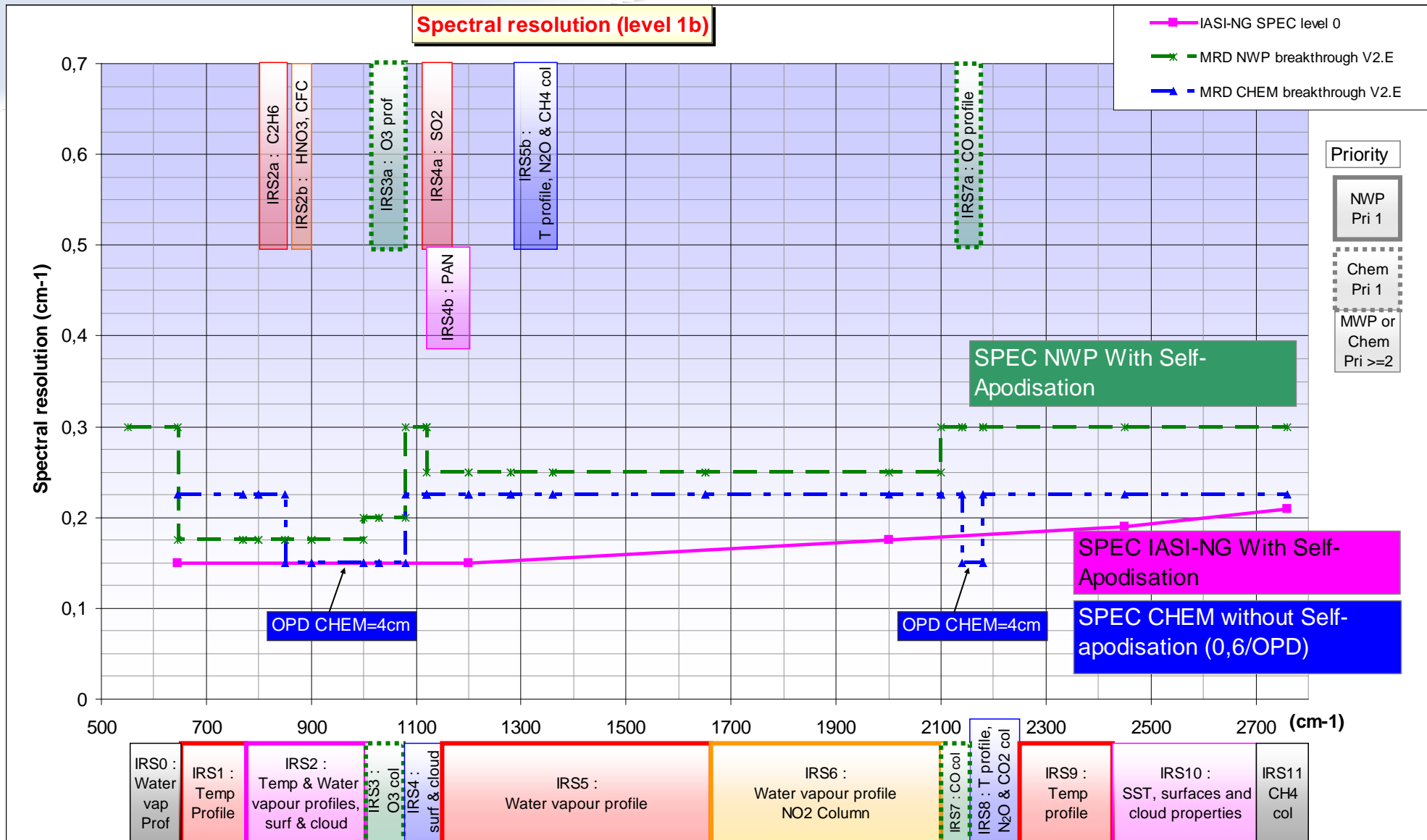
- All spectrum very useful
 - Other atmospheric data mandatory (cloud, temperature, humidity)
 - Large swath => Very good coverage.
 - Stability is very important.
- No convincing products remaining to be improved
 - CO₂ limited to upper troposphere
 - Most of products good in upper troposphere but of coarser quality in lower troposphere or BL (e.g. O₃, CO) hampering strong **development** of AQ applications
- Improve NeDT and Spectral resolution would help to detect and quantify more species and also obtain profiles.

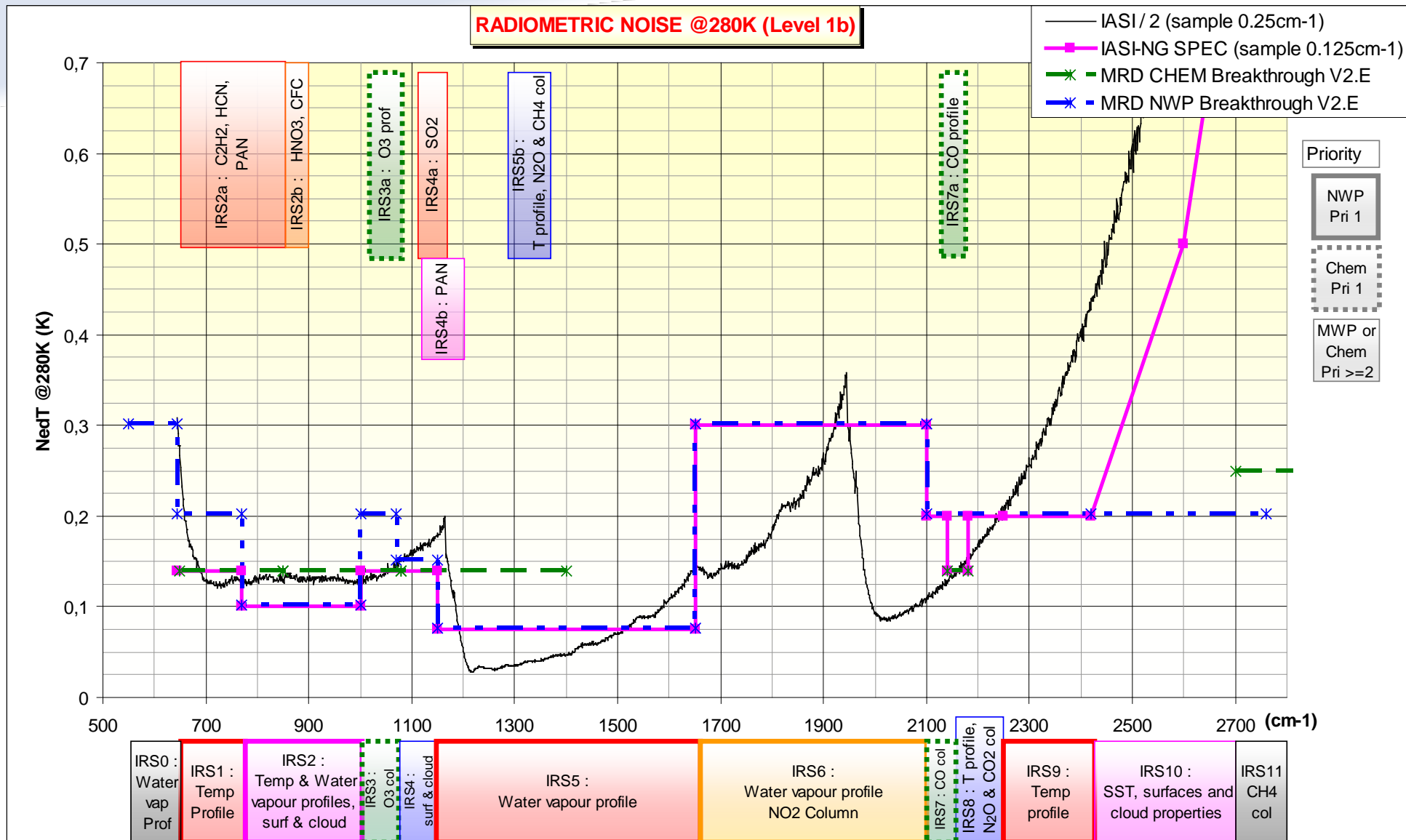
■ Two rounds :

- ◆ In phase 0 by CNES and french scientists,
- ◆ In phase A to consolidate the requirements (see further studies by MENINGE group)

■ Main results : Resolution/2 and radiometric noise /2 >

- ◆ Better accuracy of $T(z)$ in near surface layers
- ◆ More information on CO and CH₄
- ◆ Slight improvement on H₂O (?)



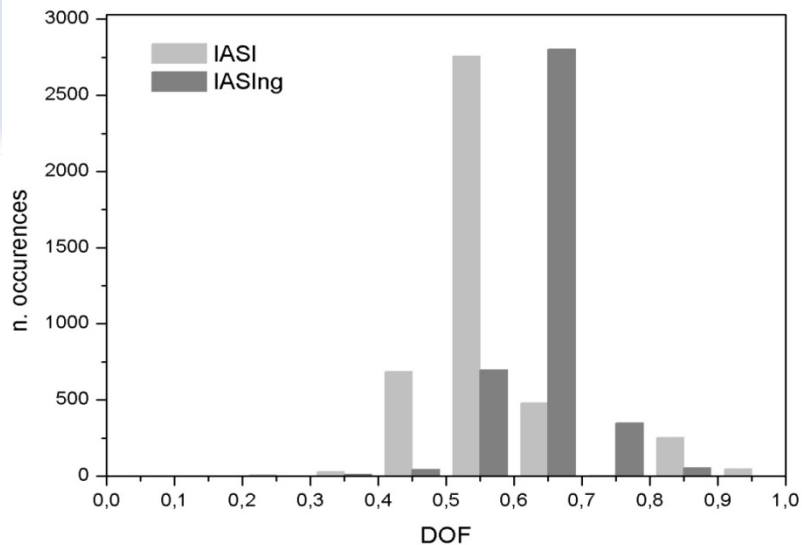


- **MENINGE : A french mission team**
 - But with links with Eumetsat, with ISSWG, with Sentinel 5 etc...

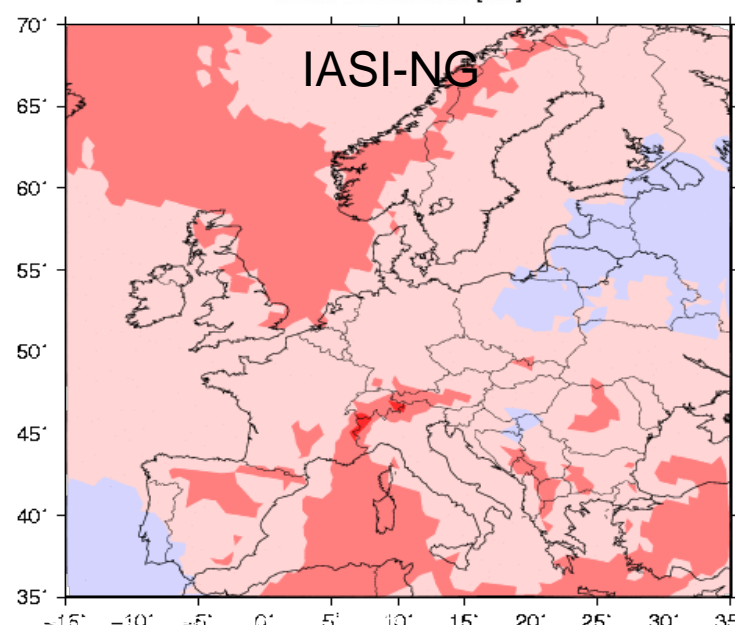
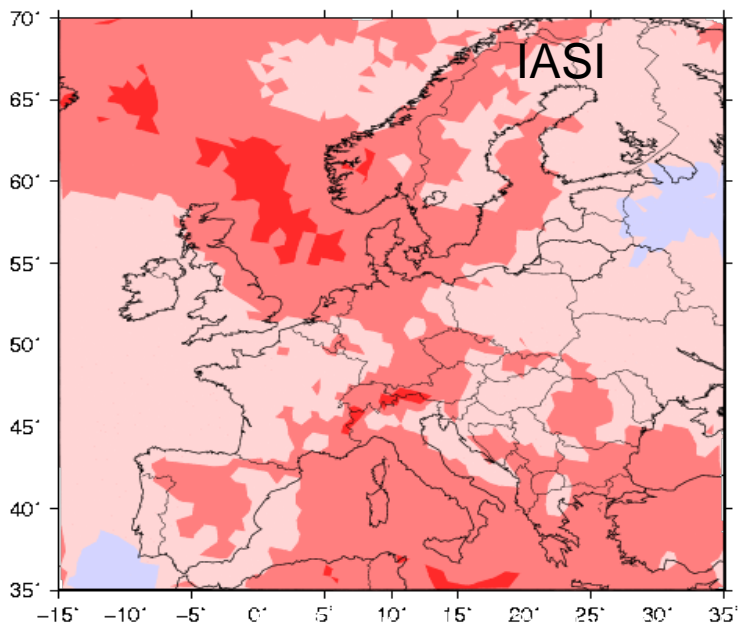
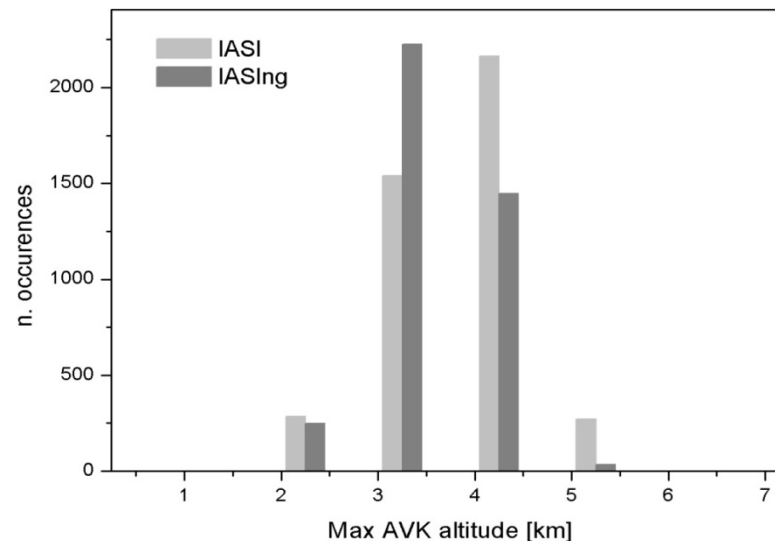
- **10 meetings during phase A**
 - ◆ **Well informed of the progress**
 - ◆ **It performed consolidation studies**
 - About needs and requirements
 - About expected performances of Level 2 or + products
 - For trade-off in the requirements for Phase B
 - ◆ **It contributed to decision making**
 - Presentation at User consultation meeting (Darmstadt)
 - Document to TOSCA

Example 1 : O₃

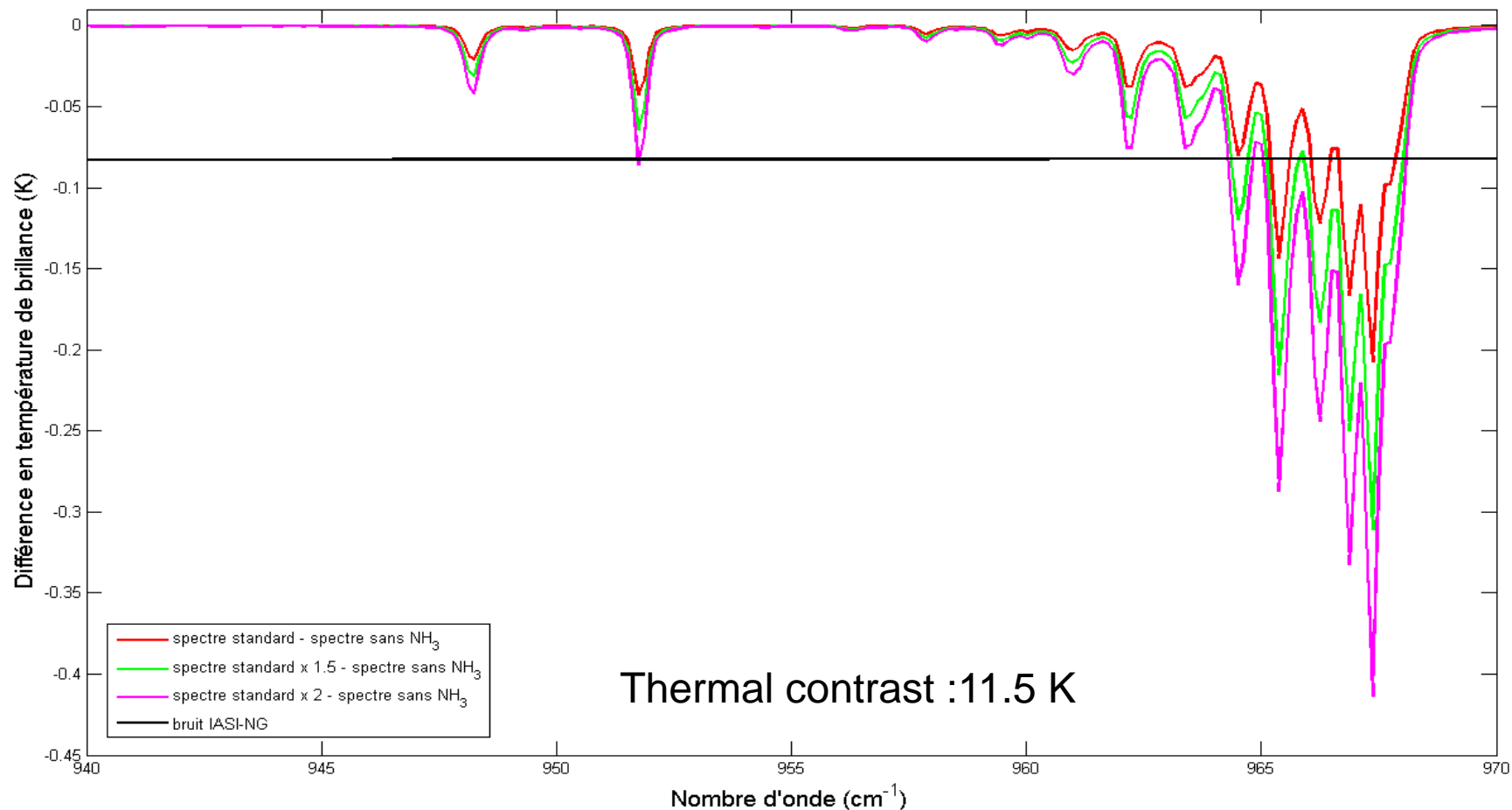
DOF 0-6 km



Max AVK 0-6 km



Différence en température de brillance entre un spectre IASI-NG simulé sans NH₃ et un spectre IASI-NG simulé à partir de l'atmosphère standard en multipliant NH₃ par 1, 1



Courtesy : J. Hadji-Lazaro

■ **MASS**

- ◆ IASI = 235 kg
- ◆ **IASI-NG spec : 350 kg (including 20% margin)**

■ **VOLUME**

- ◆ IASI : one sensor module + deported IMS/DPS inside the satellite ==> total = 1.7 m³
- ◆ **IASI-NG spec : 1500 mm x 1500 mm x 1200 mm**

■ **POWER (EOL):**

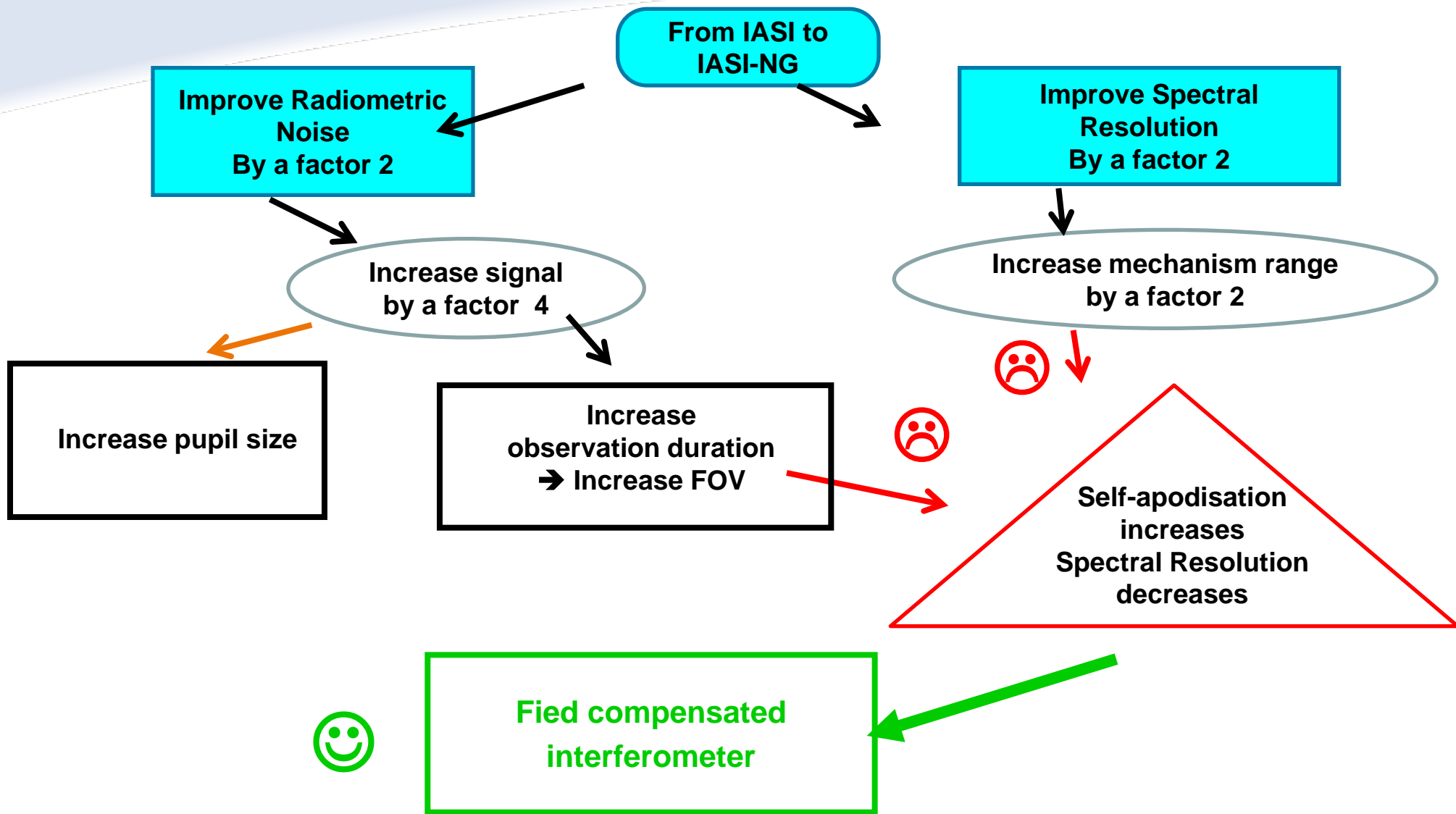
- ◆ IASI = 240 W
- ◆ **IASI-NG spec : 500 W (including 20% margin)**

■ **TM DATA RATE:**

- ◆ IASI : 1.5 Mbit/s
- ◆ **IASI-NG spec : 6 Mbit/s**

These requirements were given to 2 industrial companies for the phase A feasibility studies (competitive studies).

- **Based on MRD breakthrough + additional requirements e.g. about pseudo noise.**
- **Pseudo noise allocation derived from IASI requirements**
 - ◆ **Main noise provided by detection chain (NeDT)**
 - ◆ **Instrument defaults (ISRF shape, centroid, PSF, ...) specified as pseudo noise**
 - **Each default shall not represent more than 25% of the NeDT (globally each default contribute to few percent of the global performance) for homogeneous scenes**
 - **Default impact only heterogeneous scene has been sized to 50% of the NeDT**
- **Definition of typical heterogeneous scenes**
 - ◆ **At sounder pixel level : two half field with 10 K difference**
 - **Direct impact on heterogeneous scenes (sized to 50%)**
 - ◆ **At instrument field of view : one sounder pixel clear, the other part of the field is cloudy**
 - **Potential impact on homogeneous scenes (sized to 25%)**

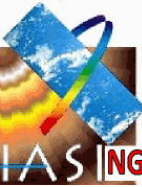


- 2 competitive industrial studies started early 2010 and finished early 2012 :
 - 2 x 3 instrument types first studied (all Fourier Transformed spectrometer with Field Compensation interferometers but with various types of field compensation)
 - Then each company focused on its best solution (choice based on the following criteria: performances, resources, schedule, risks, costs).
 - Feasibility and design optimization studies were then consolidated during 2 years
 - These deeper studies allowed to consolidate resources budgets and optimize cost evaluation
 - These studies were completed by risk reduction activities on the most sensitive/innovative points (pre-development, mock-up, test, simulation...).

Both industrial concepts have been considered feasible at the end of the phase A studies, by the CNES project team and an independent review board.

- A few potential non-compliances to phase A IASI-NG Instrument requirements at the end of phase A studies have been analyzed by MENINGE group:
 - ◆ Local NeDT non compliance
 - ◆ Absolute radiometric calibration above 2400 cm^{-1}
 - ◆ Spectral/spatial co-registration requirements
 - ◆ Absolute spectral calibration

These few non-compliances at instrument level are fully acceptable at mission level.



Résultats Etudes MENINGE
(profils verticaux T, humidité et optimisation bandes B2-B3...)

■ Known lines in this spectral range?

H₂O and isotopologues, NO, NO₂, NH₃, PH₃, OH, OCS, C₂H₂, C₂H₄, HCN, COF₂

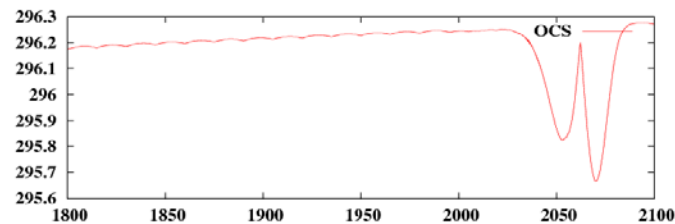
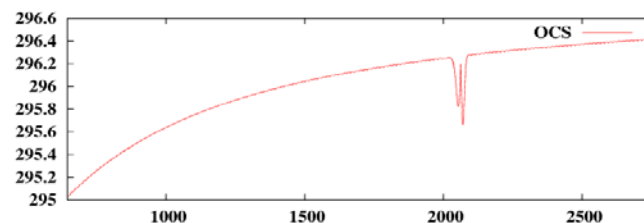
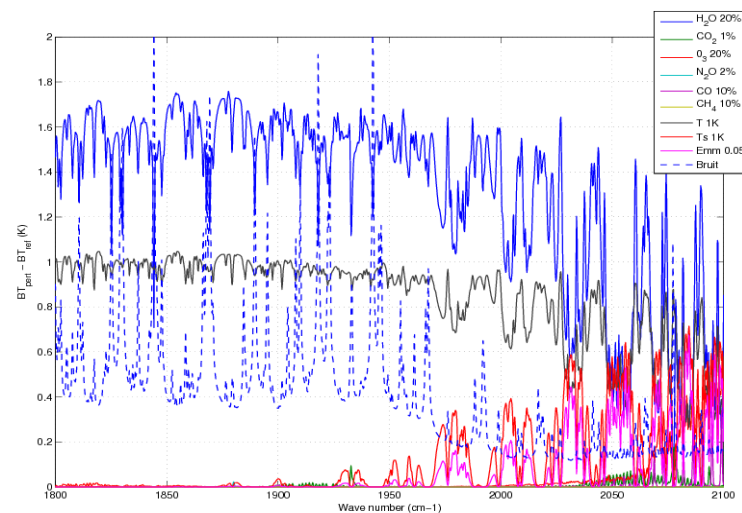
■ Who are the users?

Sometimes used to inverse H₂O

■ What is the impact of radiometric noise?

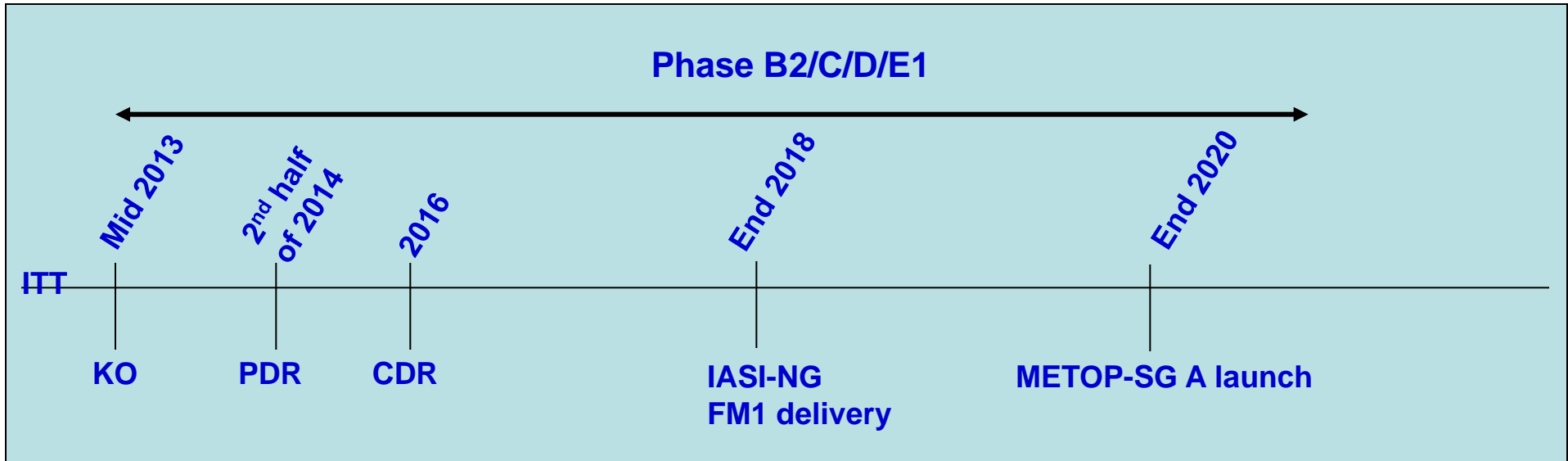
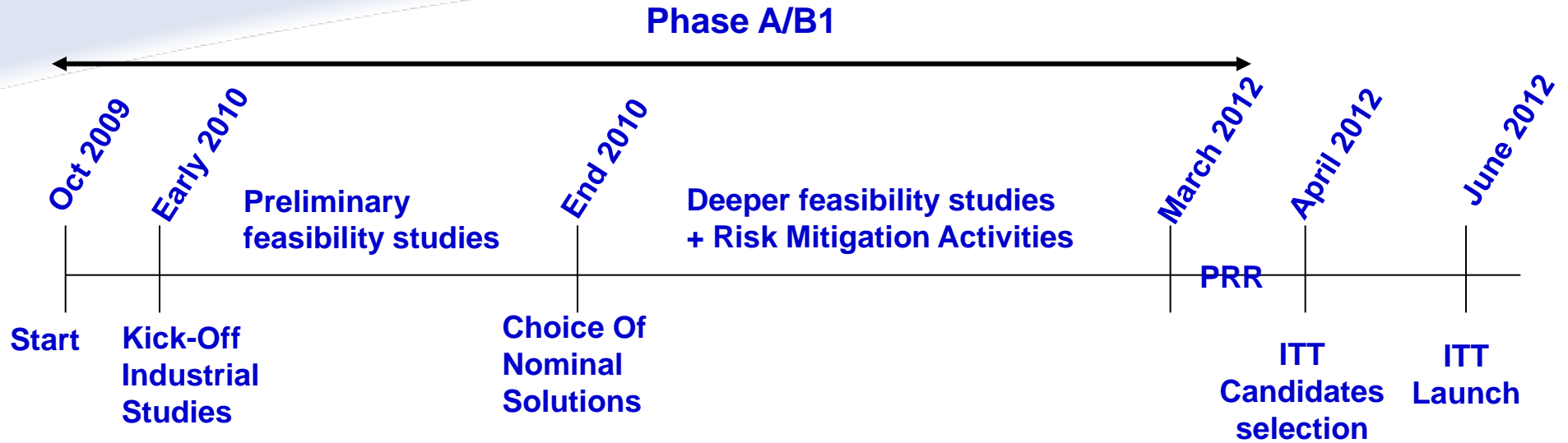
For NO (1950 cm⁻¹) signal too weak to be detectable except after averaging (=> noise lowered down). Noise can potentially impact OCS retrieval (2035-2040 cm⁻¹).

- Conclusion : Noise slightly above the current requirement is acceptable. It could allow a small shift of the band limit (not higher than 2030 cm⁻¹)



- Two feasible industrial solutions fully compliant with the IASI-NG mission requirements have been studied and are proposed for phase B-C-D-E.
- MENINGE group and EUMETSAT EPS-SG project confirmed the acceptability of the IASI-NG expected performances
- CNES commitment in April 2012 to provide IASI-NG instruments to EUMETSAT for the EPS-SG program
- IASI-NG ITT for B-C-D-E phases launched in June 2012 (after an update to new interface requirements evolutions coming from ESA following the MetOp-SG PRR).
Only 2 candidates (the 2 industrial companies who worked on the IASI-NG Phase A).
- Presentation to ISSWG which will become the Mission group for IASI-NG
- Cooperation agreement between CNES and EUMETSAT prepared for signature

Selection of the best industrial proposal for B-C-D-E phases on going



- **Choice of Manufacturer and concept shortly**
- **Start of phase B in fall 2013.**

- **Meeting of MENINGE group**
- **Science plan for IASI-NG to be established by ISSWG**
- **Start studies of ISSWG**

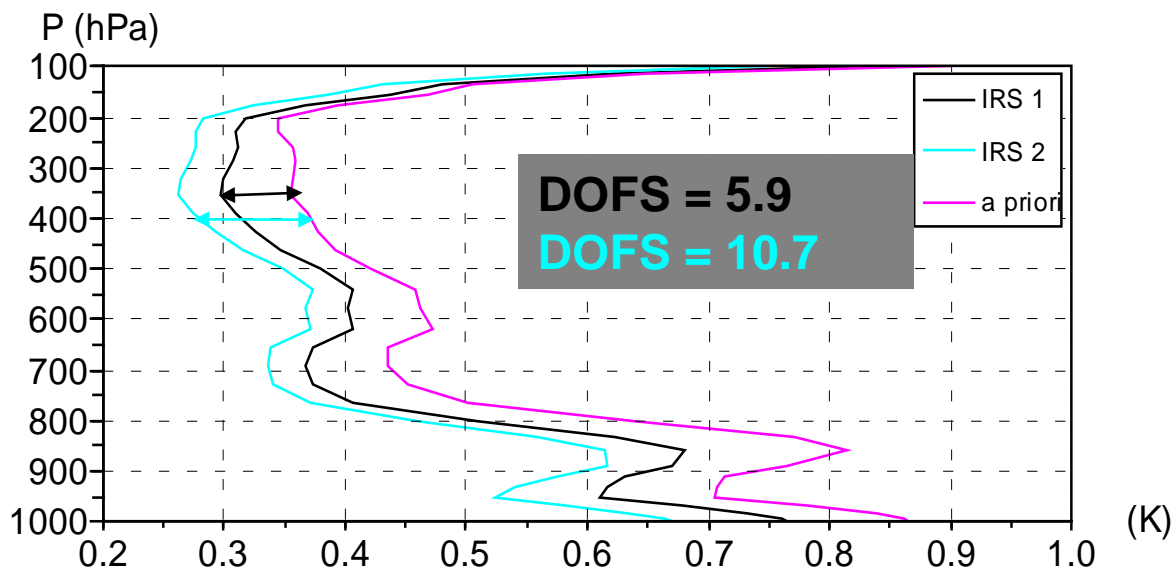
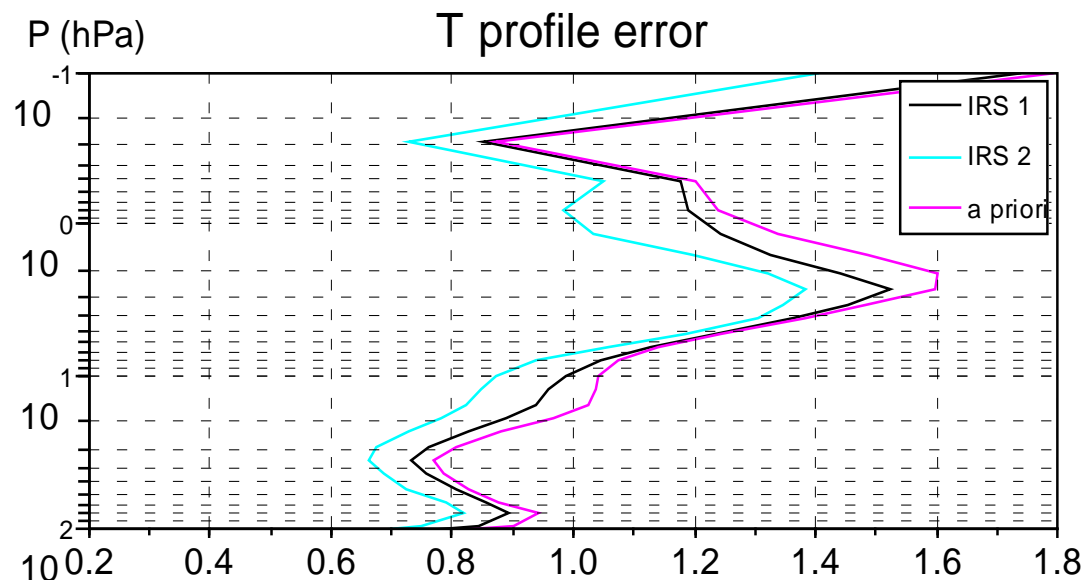
- **Final agreement btw CNES and EUMETSAT**

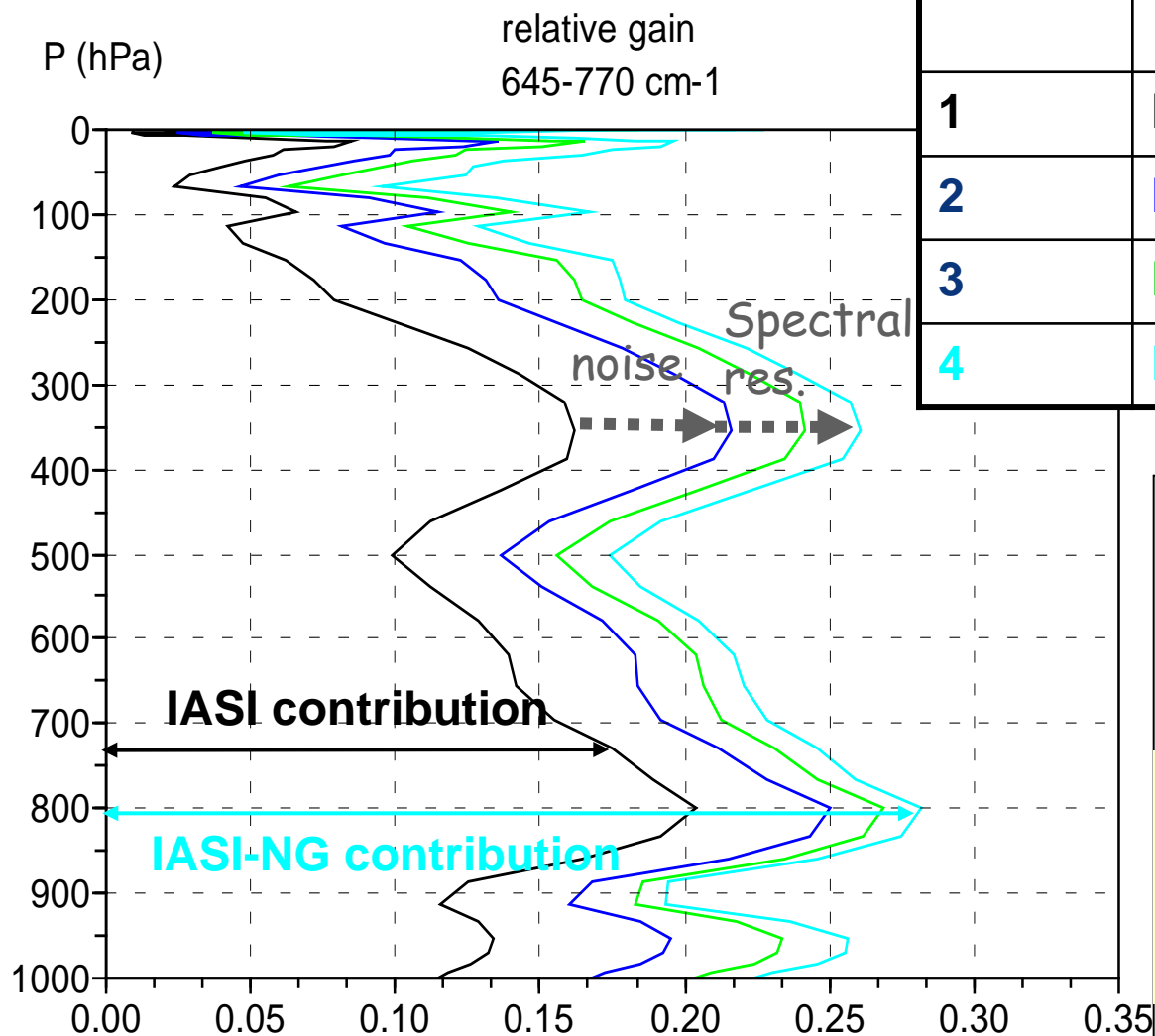
- **Studies on synergy with Sentinel 5 and other instruments**

- **Approval of EPS-SG programme by EUMETSAT Council mid 2014**

Back-up slides

- Profile retrieval using the 15 μm and 4 μm CO_2 bands
 - Tropical atmosphere
 - Noise contribution from uncertainties on surface temperature and emissivity, humidity profile.
 - A *a priori* covariance from ECMWF
- With respect to a *a priori* uncertainty, IASI-NG contribution is about twice IASI contribution





	Noise	Spectral resolution
1	IASI	IASI
2	IASI/2	IASI
3	IASI/(2√2)	IASI
4	IASI/2	IASI/2

- The relative gain (or error reduction) is defined as $(a \text{ posteriori} - a \text{ priori}) / (a \text{ priori})$
- It is in the range 5 - 25%.

- Spectral resolution improves the instrument contribution, beyond noise reduction by increasing the number of channels (\sqrt{n})

Atmospheric chemistry : CO profile (2140-2180 cm^{-1})

	MRD threshold	MRD breakthrough	IASI	IASI-NG reference	IASI-NG with cold bench
<i>Spectral resolution (cm^{-1})</i>	0.15	0.1	0.5	0.25	0.25
<i>Spectral sampling (cm^{-1})</i>	0.124	0.083	0.25	0.125	0.125
<i>NeDT (K)</i>	0.15	0.1	0.28	0.16	0.12
<i>DOFS</i>	2.84	3.78	1.88	2.51	2.69
<i>Tropospheric column (0-12 km) (%)</i>	2.14	1.95	3.39	2.24	2.15
<i>Boundary layer (0-3 km) (%)</i>	10.2	7.8	16.05	10.7	10.1

IASI instrument is not sufficient to provide a vertical profile (less than 2 DOFS). IASI-NG, though not meeting the level 1 requirements, is close to the threshold level 2 performances (especially for the cold bench configuration).

Atmospheric chemistry : CH₄ column (1280-1360 cm⁻¹)

	MRD threshold	MRD breakthrough	IASI	IASI- NG
<i>Spectral resolution (cm⁻¹)</i>	0.3	0.15	0.4	0.2
<i>Spectral sampling (cm⁻¹)</i>	0.248	0.124	0.25	0.125
<i>NeDT (K)</i>	0.2	0.1	0.1	0.05
<i>Total column (%)</i>	13.1	6.9	9.9	4.6

- The CH₄ column can be retrieved with a precision in agreement with the MRD threshold (resp. breakthrough) from IASI (resp. IASI-NG) observations.

Different trade off's during Phase 0 done only for dynamic FTS

Radiometry

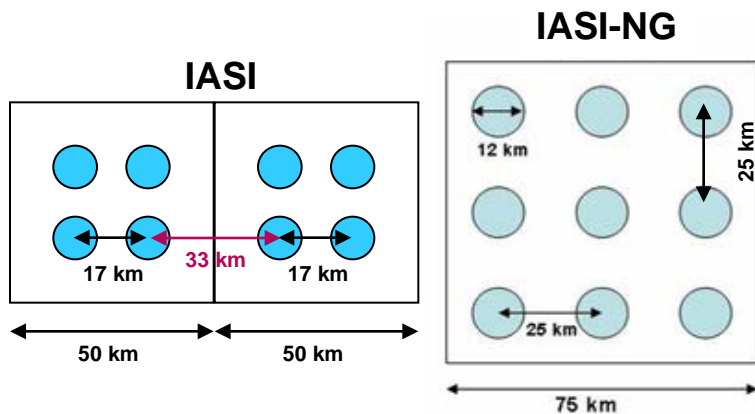
- Entrance pupil diameter Increase (→flux):
 - IASI-NG = 120 mm versus IASI = 80 mm
- Instrument Field of view Increase (→integration time):
 - IASI-NG = 75 km×75 km (9 pixels) versus IASI = 50 km×50 km (4 pixels) → acquisition duration for each interferogram = 450 ms versus 150 ms for IASI
- Detectors temperature reduction (active cooling)
 - IASI-NG T detector < 65 K versus IASI = 92 K

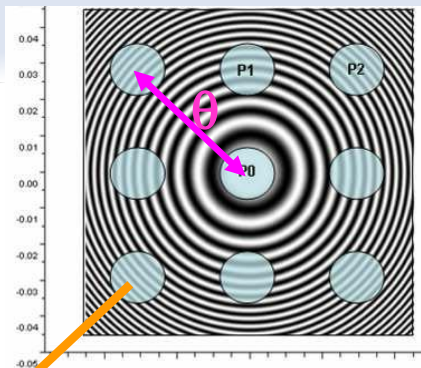
Spectral resolution

- Optical Path difference increase by factor 2:
 - single sided interferometer
 - one mobile cube having the same IASI stroke
 - TWO mobile cubes having the IASI stroke / 2
 - double sided interferometer
 - one mobile cubes having the IASI stroke x 2
 - TWO mobile cubes having the same IASI stroke

BUT: «self-apodisation»

*For the spectral resolution, both the Optical Path Difference **AND** the self-apodisation must be improved.*





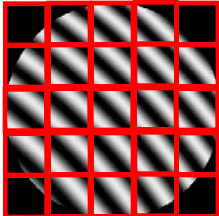
IASI-NG: Optical signal in the detectors plane for $\nu=2760 \text{ cm}^{-1}$ and $\text{OPD}_{\text{max}}=4 \text{ cm}$

The use of IASI like pixel acquisition concept is not possible. Two options has been studied.

Split the sounder pixel into many smaller pixels
→ Matrix detectors

Suppress/mitigate the self-apodisation effect
→ self-apodisation compensation

Sounder pixel



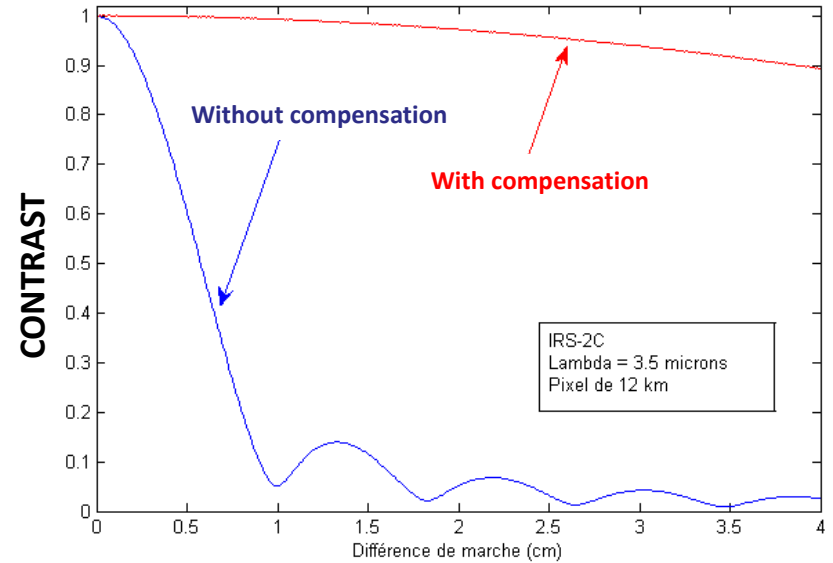
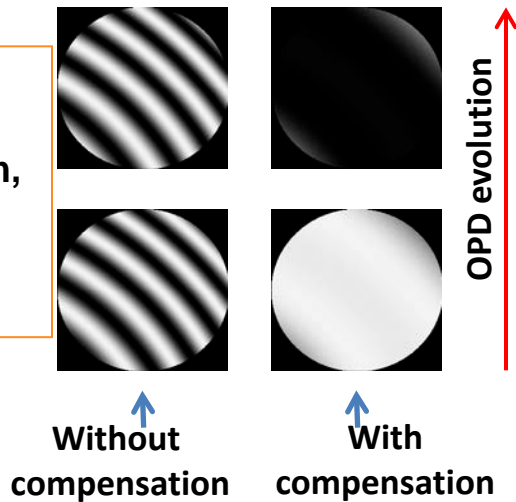
- For each of the 9 sounding pixel, acquisition of sub-pixels interferograms (typically 5x5 sub-pixels), then combination of the interferograms by resampling at “constant” OPD + filtering to generate one sounding pixel interferogram
- Final Self-apodisation for the corner pixels = IASI one

- Introduce, in the interferometer, a specific mechanism that works in synchronization with the Corner Cube Mechanism and that corrects, for each sounding pixel (but the center one), the Optical Path Difference by : $\delta_o (1 - \cos \theta)$, where δ_o is the OPD for a zero field → **ALL the pixels should have a similar behavior than the central pixel**

IASI-NG CNES Phase 0 Reference option

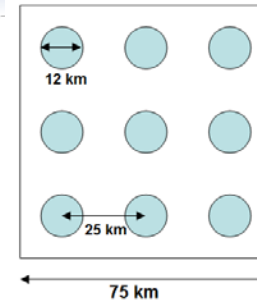
Optical simulation with and without active optical field compensation

Evolution of a corner pixel optical signal for monochromatic ray at $4\ \mu\text{m}$, for OPD variation of $1\ \mu\text{m}$, (between max and min signal ($\lambda/4$)) at OPD of 4cm





- ◆ Ground Pixel diameter of 12 km (=IASI)
- ◆ Ground sampling of 25 km (both axis)
- ◆ Number of sounder pixels per acquisition = 9 (IASI=4)
- ◆ Number of earth view per line = 20 (30 IASI)
- ◆ Interferogram acquisition duration = 450 ms (IASI=150 ms)



◆ Inlet PUPIL = 120 mm (IASI=80 mm)

- ◆ Focal plane :
 - 4 bands (IASI=3)
 - 9 sounder pixels per band (IASI=4)
 - PV detectors for all bands (IASI PC for B1; PV for B2/B3)
 - detectors cooled at 65 K with one active cooler (LPTC) (IASI=92 K passive)
 - IASI like cold optic concept

