

IASI on MetOp-B Radiometric Calibration

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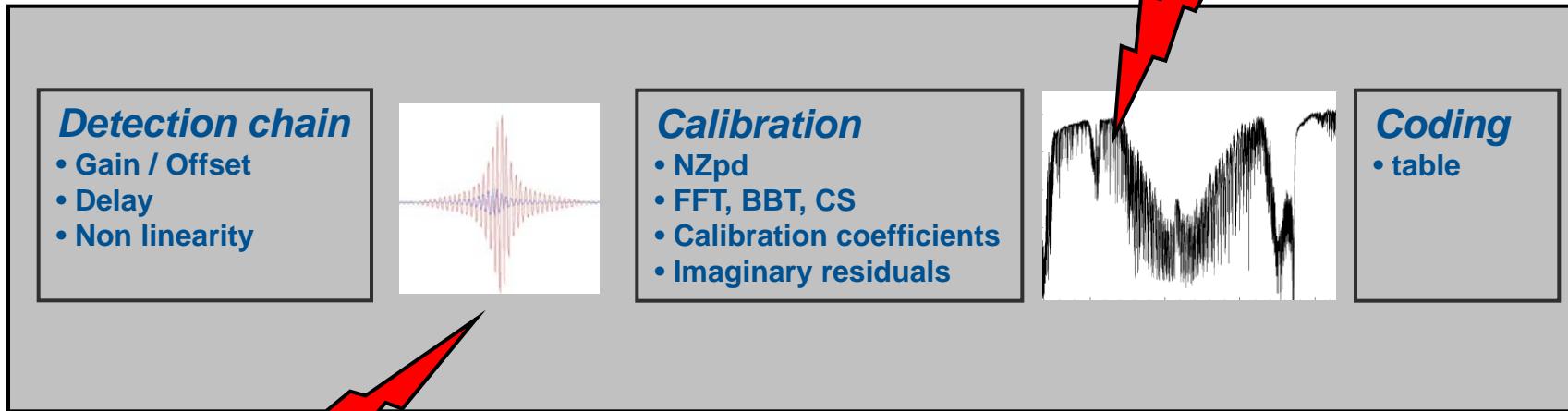
1 Centre National d'Etudes Spatiales (CNES)

2 Noveltis (Toulouse)

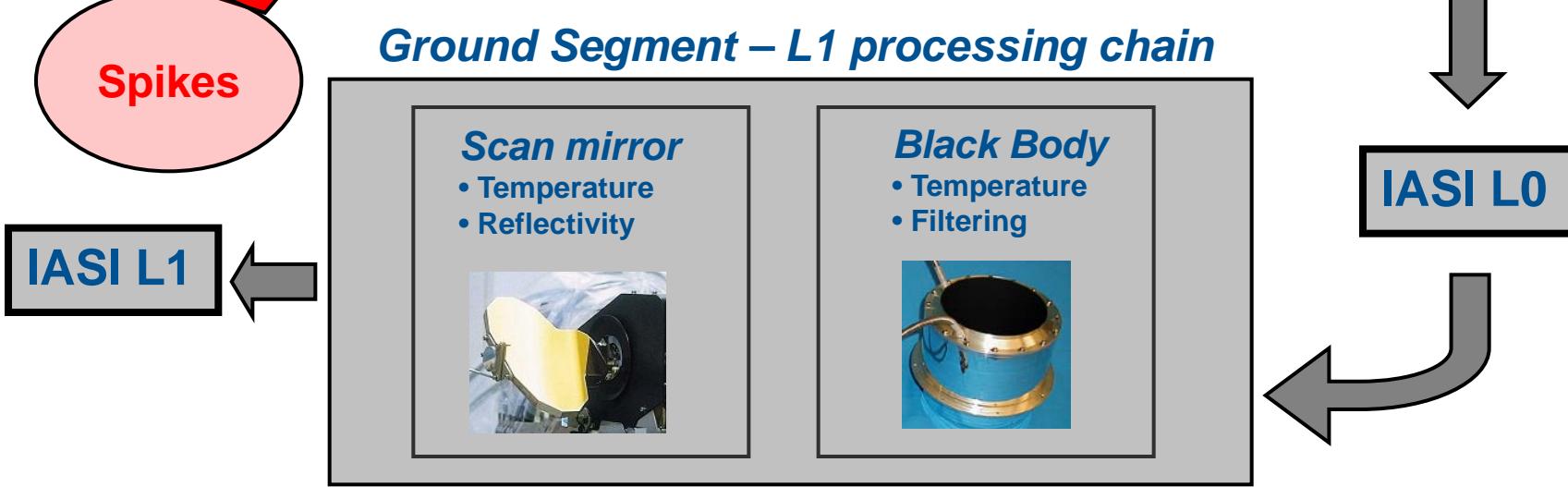
3 AKKA (Toulouse)

3rd IASI conference, 4-8 February 2013, Hyères, France

OUTLINE

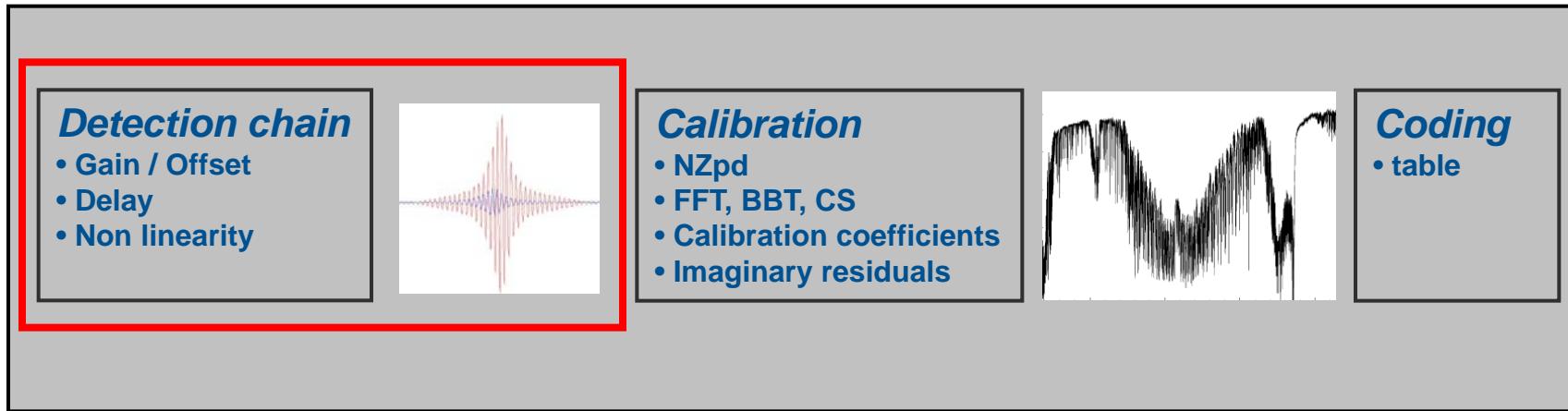


Ground Segment – L1 processing chain

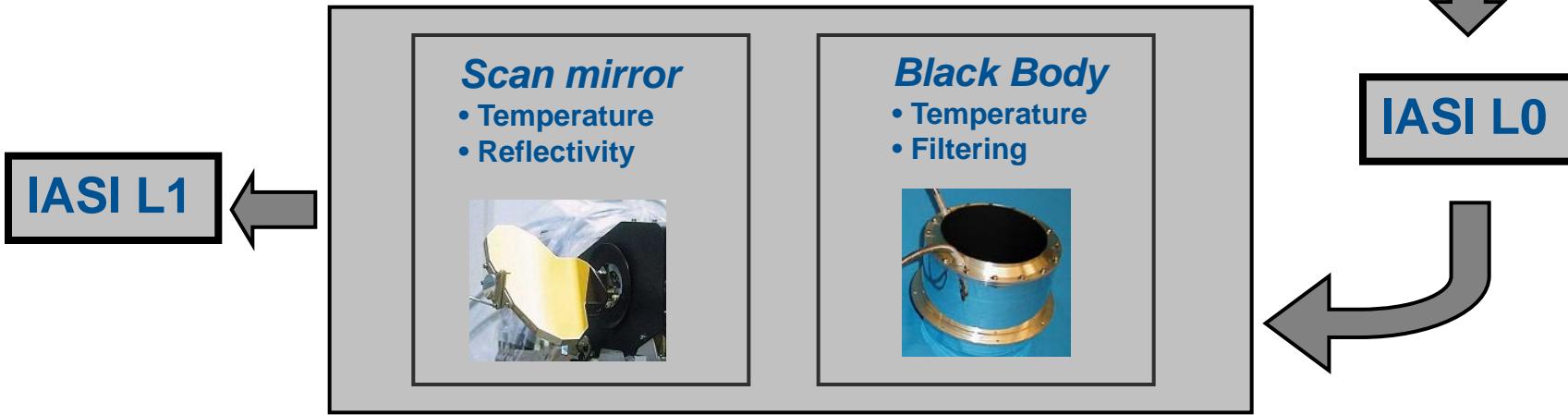


OUTLINE

Space Segment – IASI instrument



Ground Segment – L1 processing chain

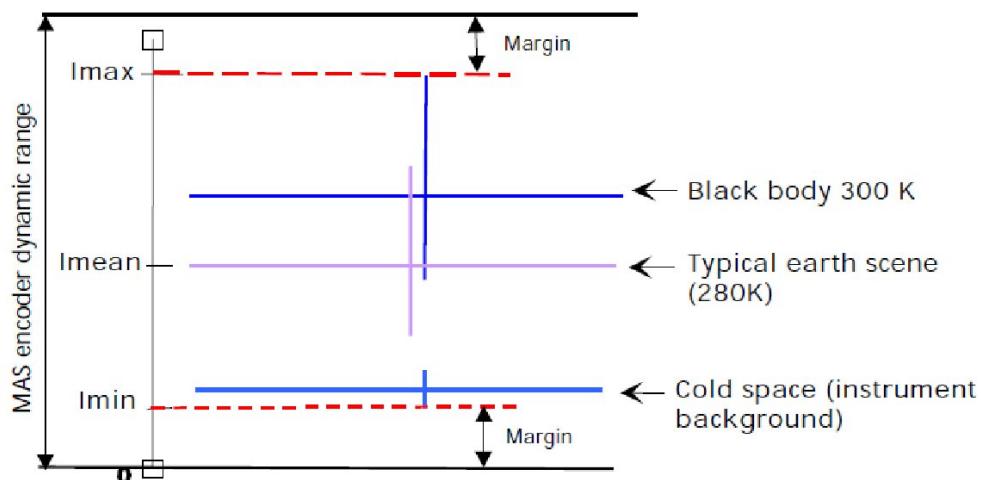


Detection chain



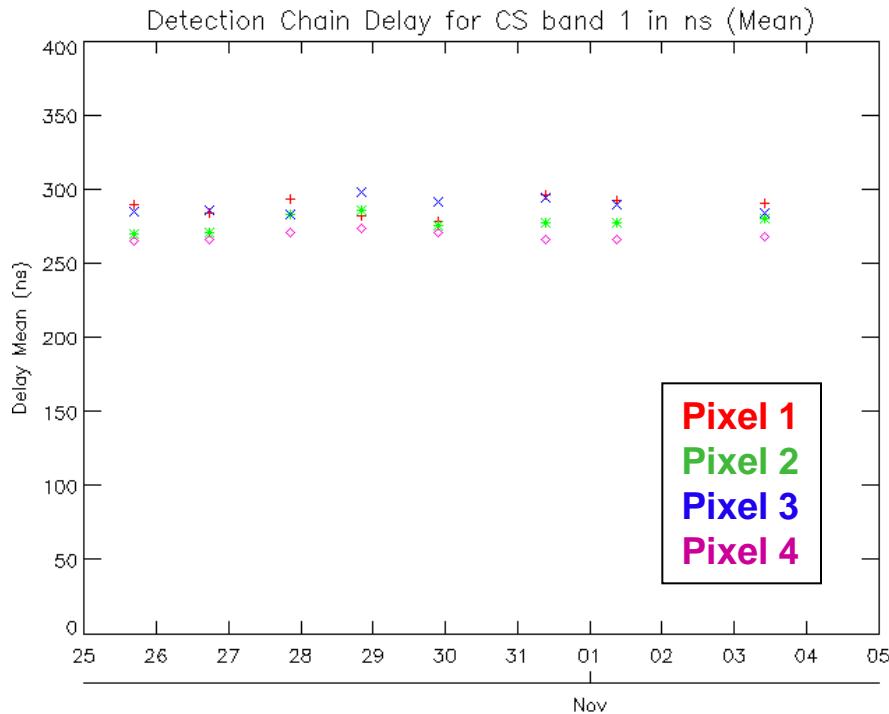
Detectors

Digital processing system



- optimize detector signal amplification (Signal / Noise)
- for each pixel, each spectral band, 8 gains available
- Interferogram dynamic monitoring have shown that there's no need of gain & offset update

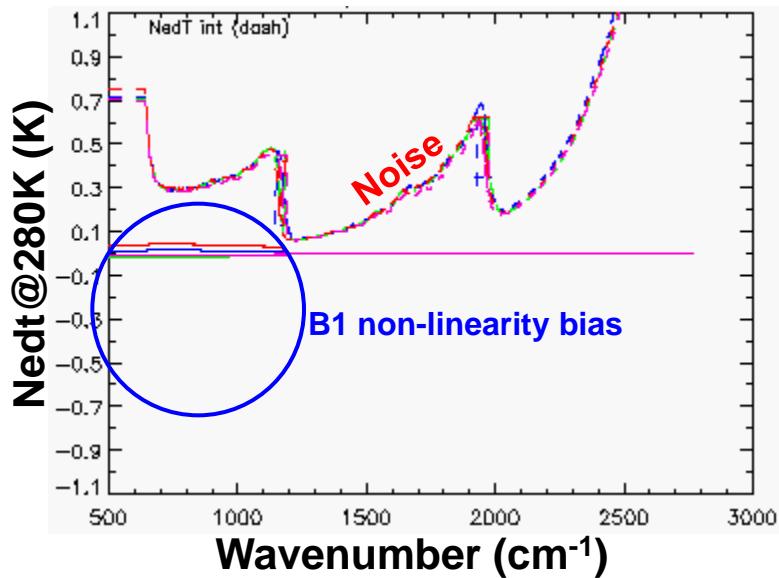
Detection chain



DELAY

- Monitored for calibration views (Black Body, Cold Space), Earth views
- Specification = max 400 ns
- On-board tuning
- Stable : no update needed

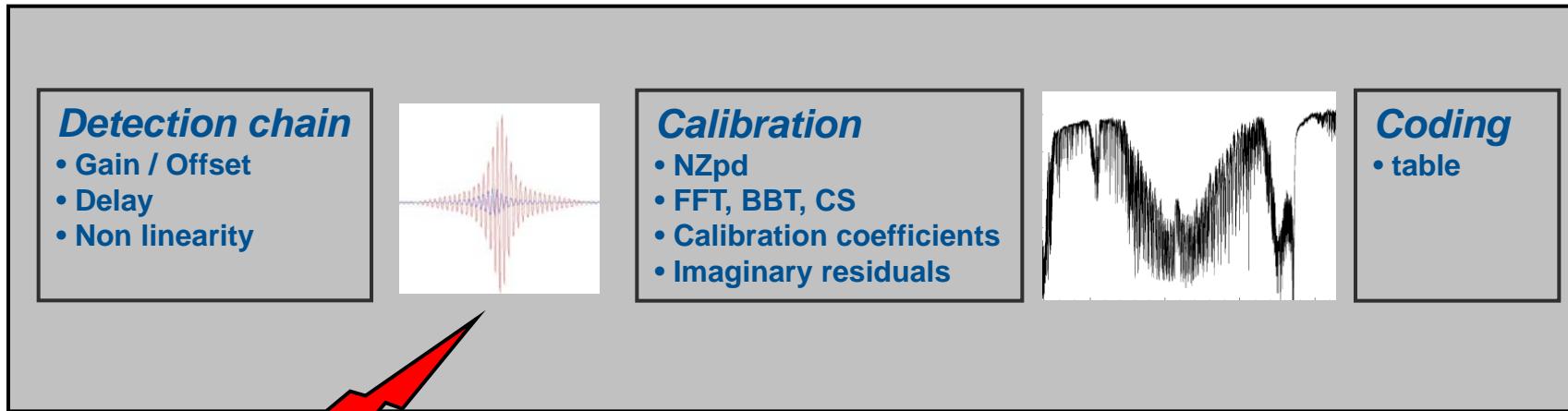
Detection chain



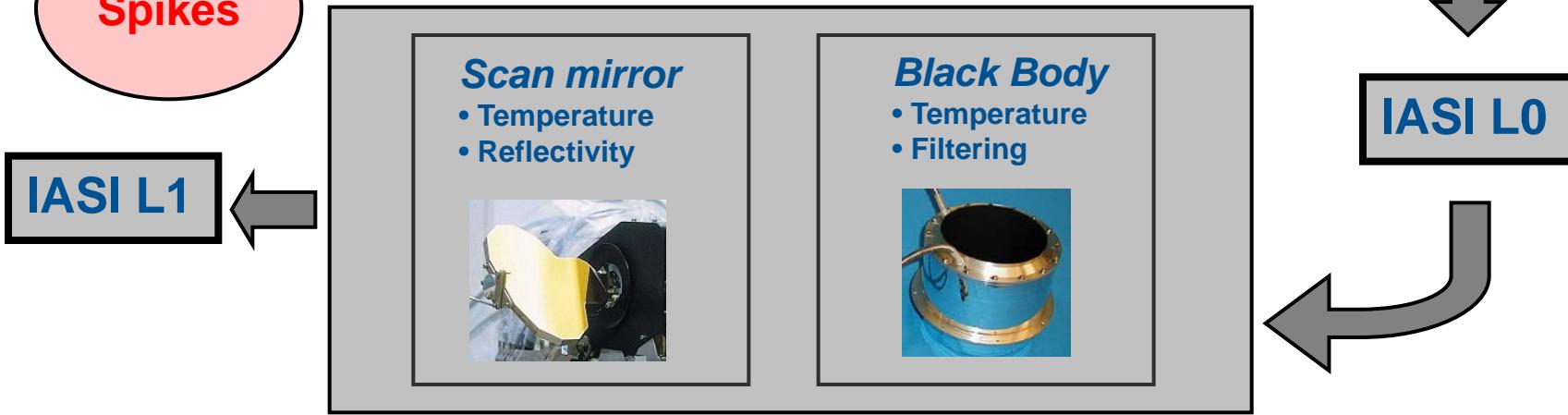
- Can cause radiometric error / bias !!
- Spectral Band 1 is the most non-linear
- Non linearity = f (detector temperature)
- Detector target T has changed
- Non linearity bias 0.3-0.5 K => < 0.05 K

OUTLINE

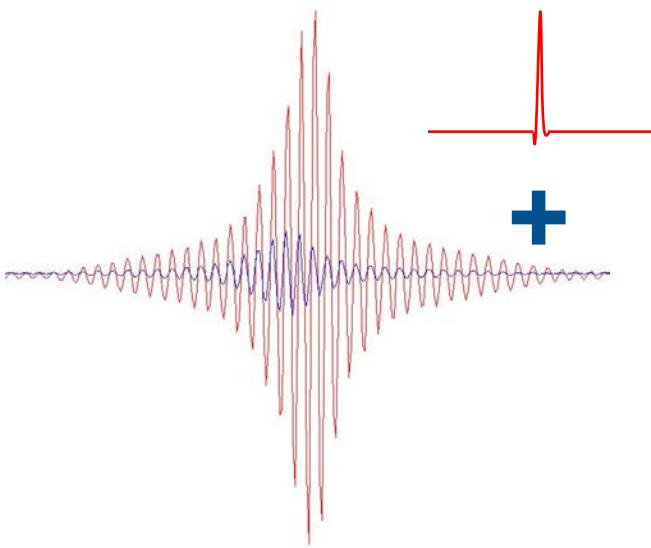
Space Segment – IASI instrument



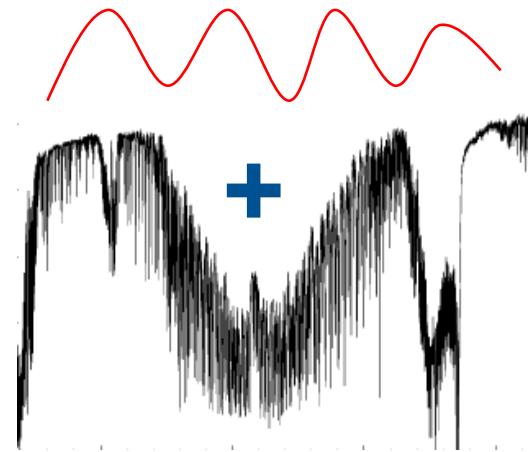
Ground Segment – L1 processing chain



Spikes



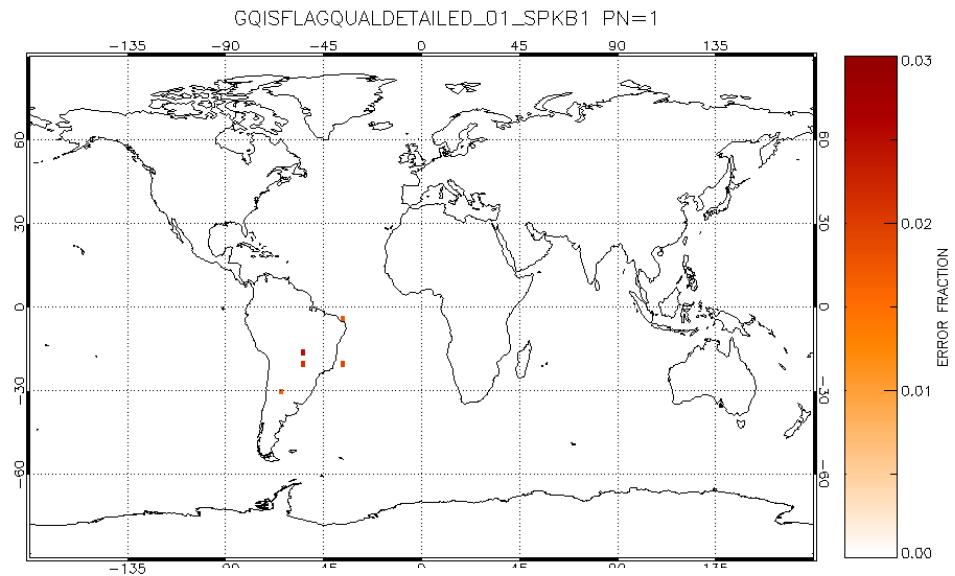
Fourier Transform
→



On-board spike filter :

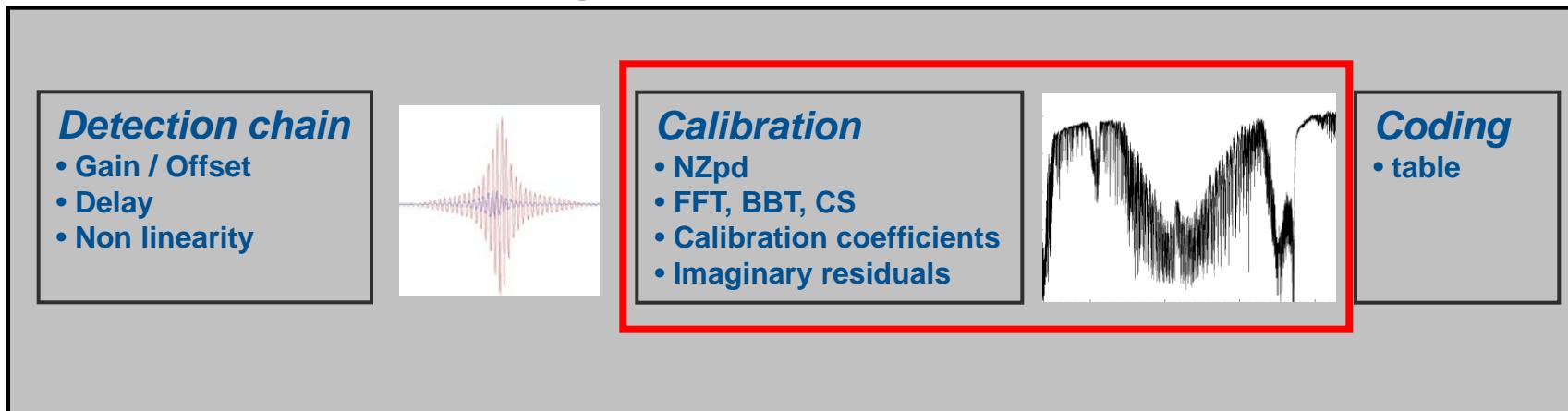
Method :

- Determine spike model
- Simulate spike effect in the output spectrum and determine the threshold above which the signal in the output spectrum becomes greater than $\frac{1}{4}$ instrument noise.
- Load the threshold into on-board configuration.

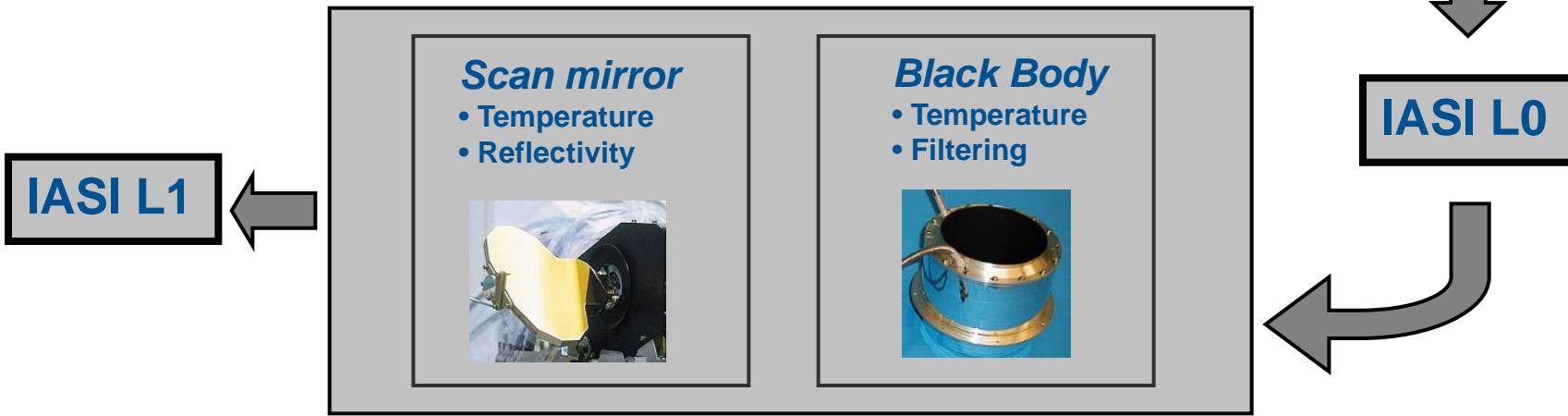


OUTLINE

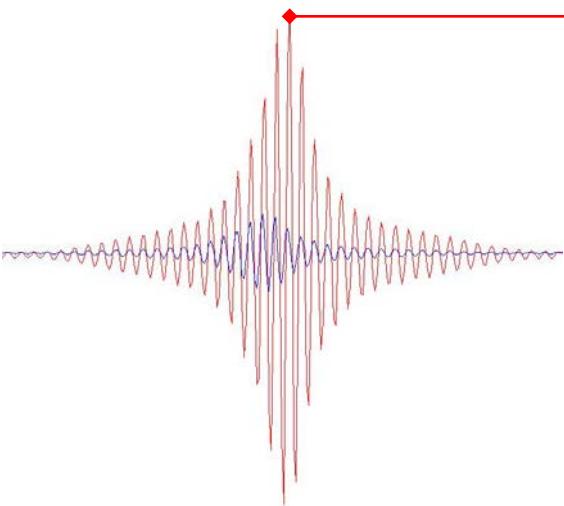
Space Segment – IASI instrument



Ground Segment – L1 processing chain



NZpd determination



Method : 3 steps determination (soft)

- 1- Interferogram maximum
- 2- Central fringe barycenter
- 3- Search for the optimal NZpd with respect to the calibration line

Effect of a wrong ZPD:

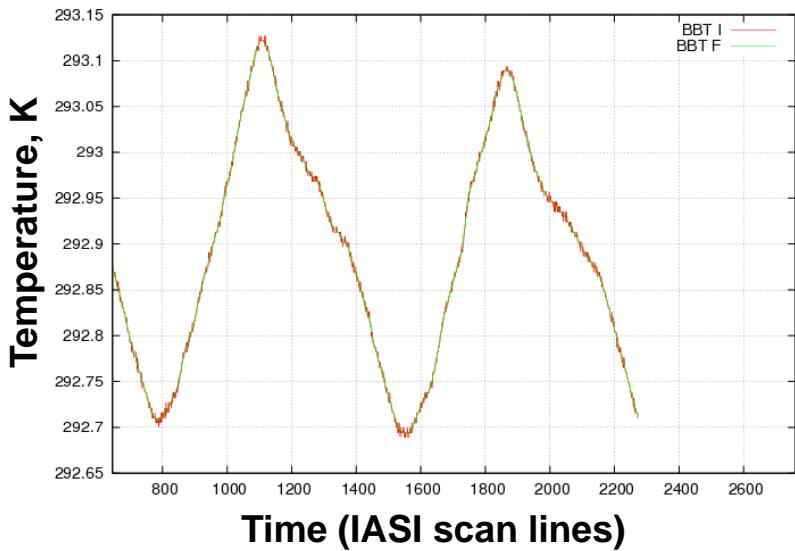
- ↔ Shift of interferogram
- ↔ Rotation of output spectrum in complex space
- ↔ Radiometric calibration error

NZpdQualityIndex :

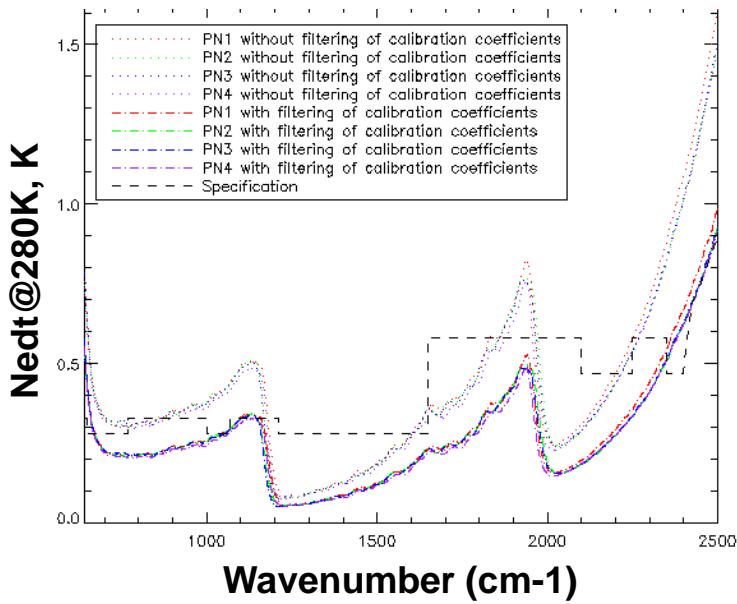
Quality indicator of NZpd determination

Complex calibration method & coefficients

Black body temperature over 2 orbits



Simulated instrument noise

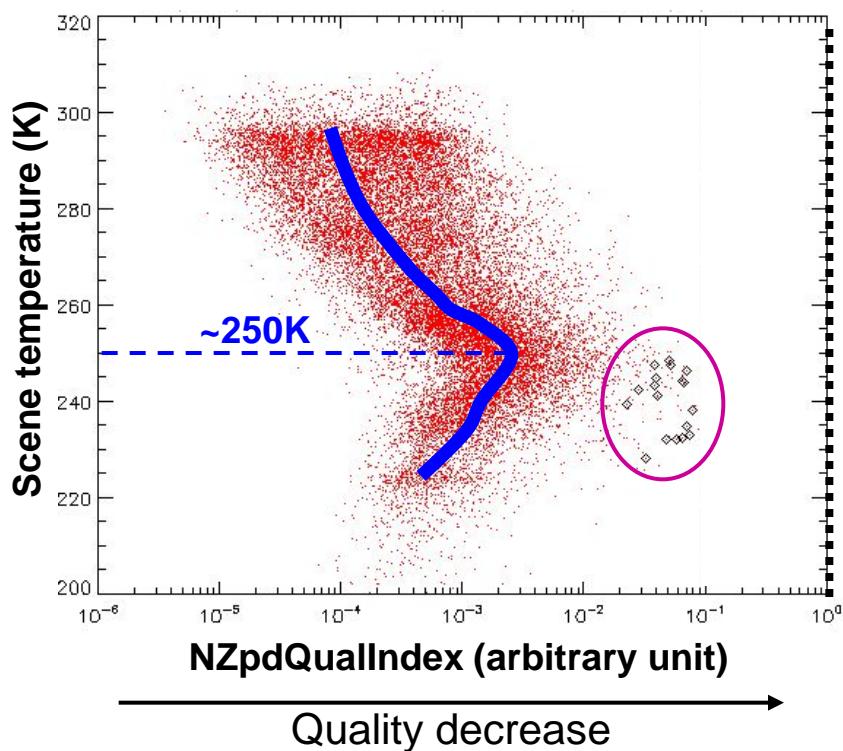


- Calibration coefficients are computed every scan line using BB and CS
- Tunable filter on-board
- Simulation with instrument numerical prediction model
- Optimum => 20 scan lines (160 s)

Calibration residuals : on-board data filtering

NZpdQualIndex

Quality Indicator of NZpd determination

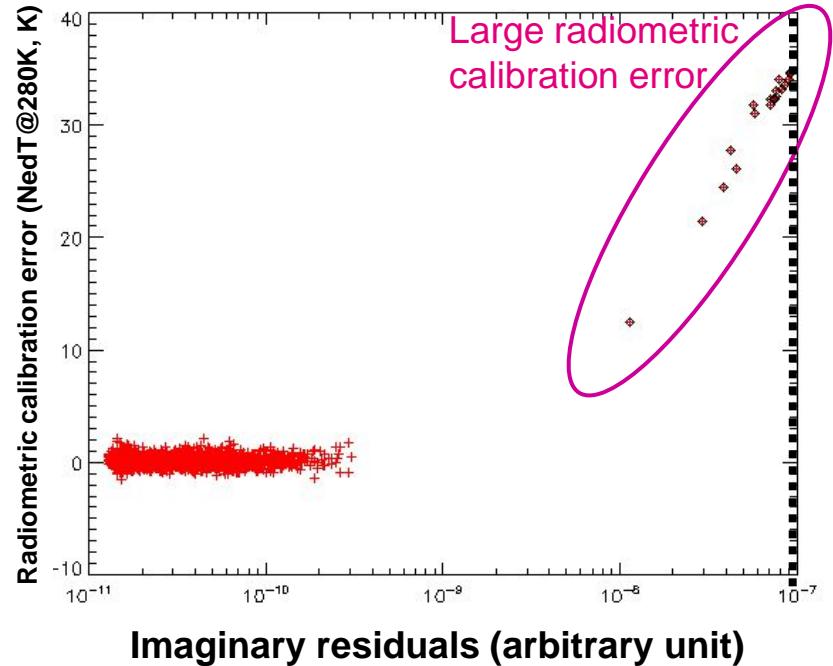


- Bias around instrument background
- => relaxed on-board filtering

=> Imaginary residuals have a higher discrimination power

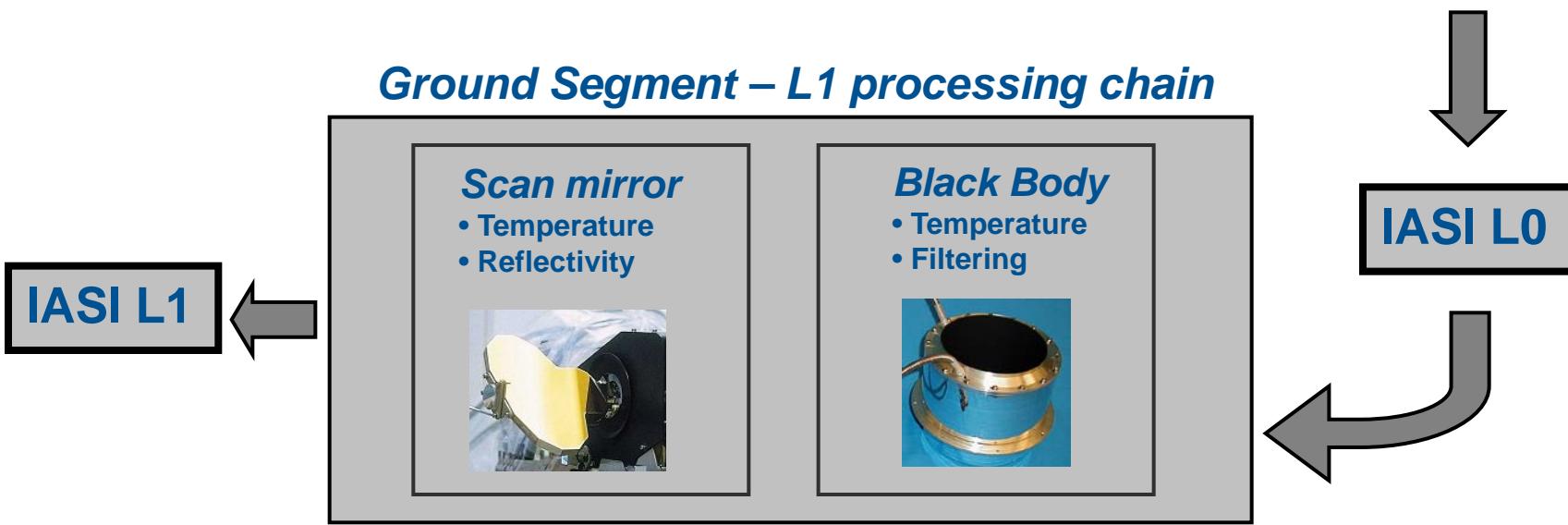
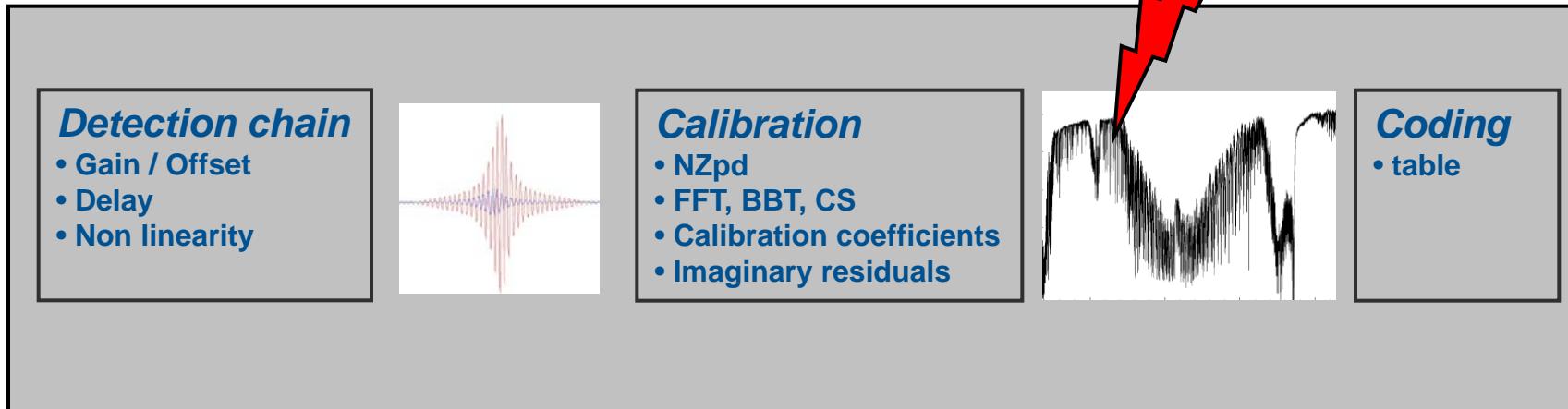
BArcImagMeanRms

Calibration imaginary residuals



- efficient on-board filtering
 - No bias
 - No false alarm

OUTLINE

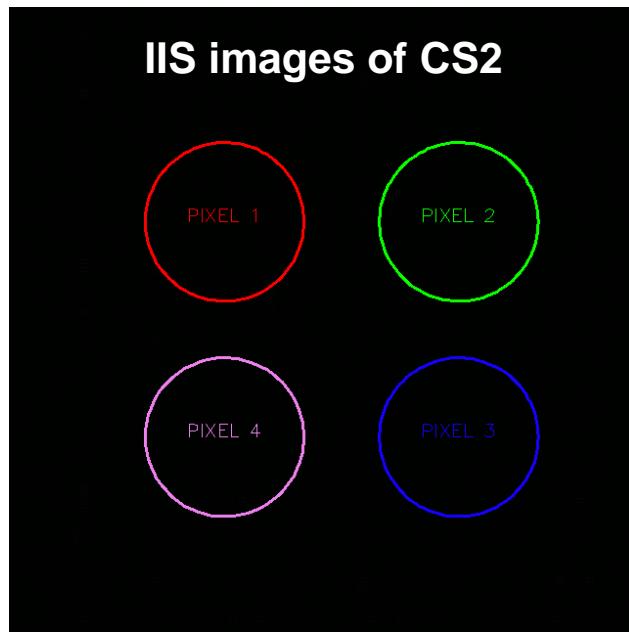
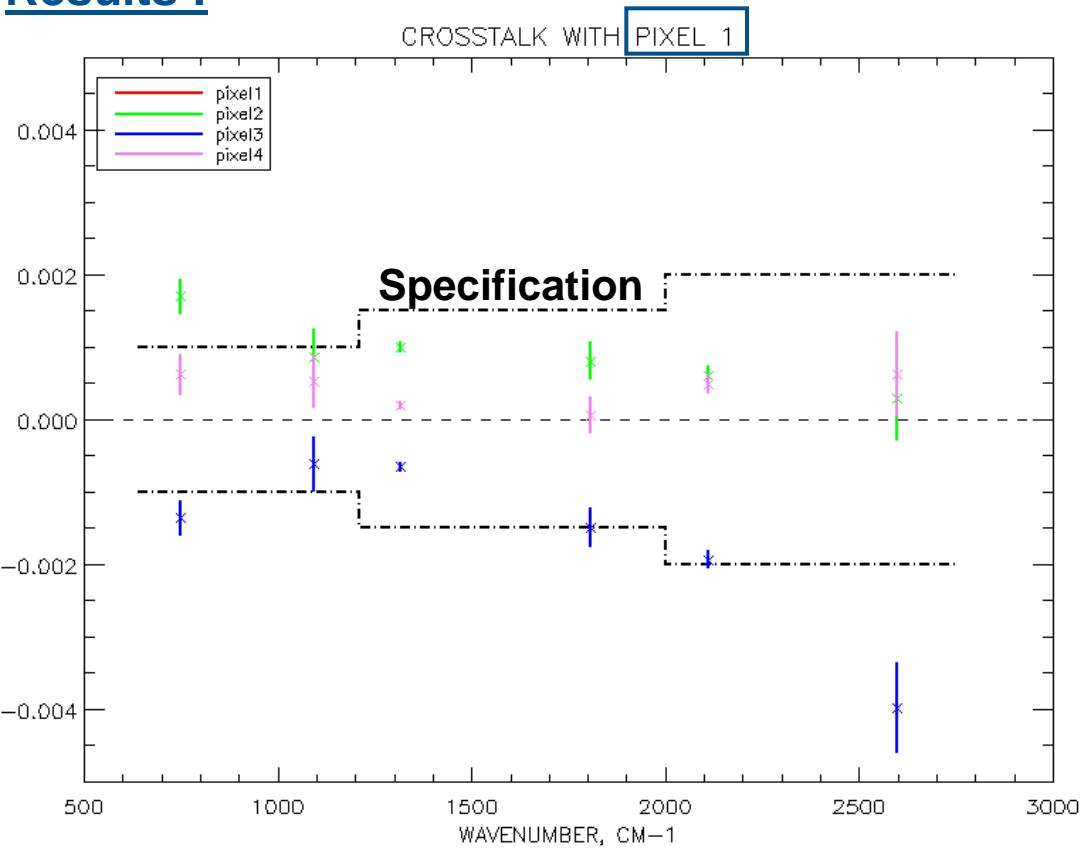


Straylight

Method : To use Moon transit in Cold Space 2 target with the instrument in external calibration.

Crosstalk PN_x => PN_y = F_{_PNy} / F_{_PNx}

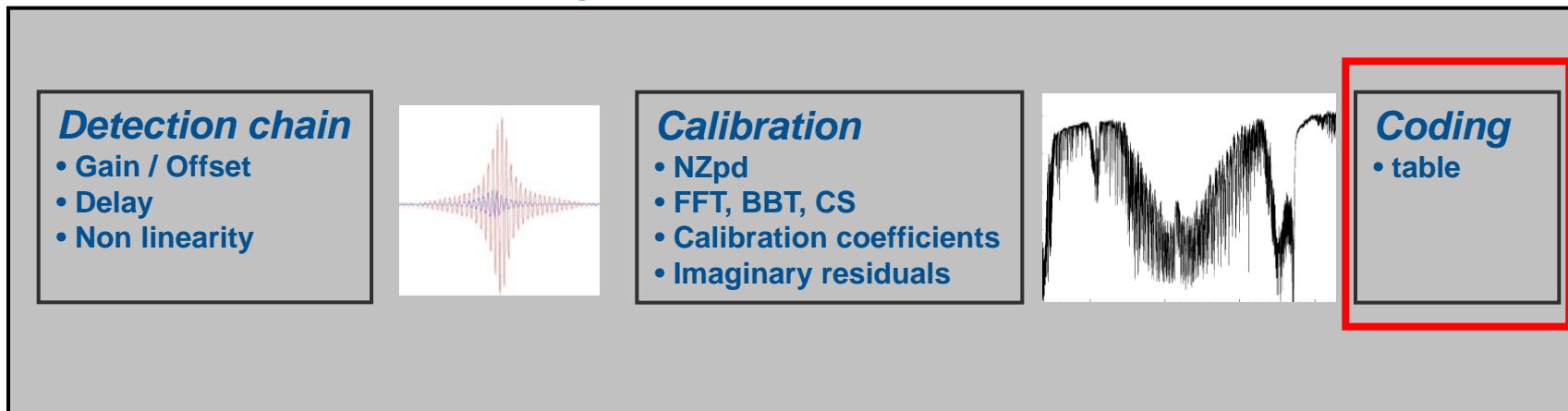
Results :



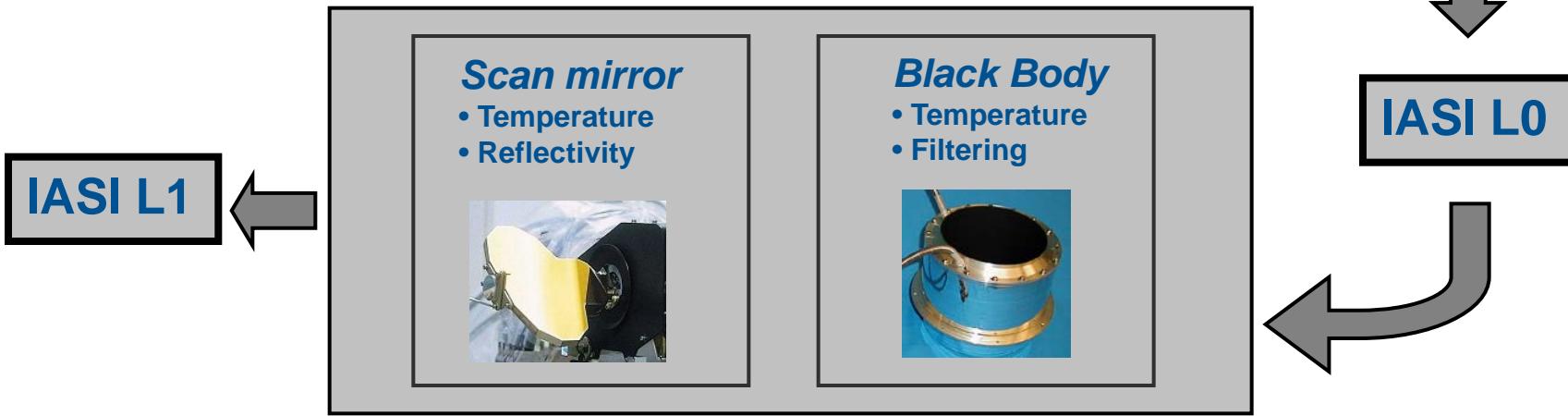
- No significant crosstalk for Pixel 2 and Pixel 4
- Pixel 1 and Pixel 3 seems to be weakly anti-correlated (~1/1000)

OUTLINE

Space Segment – IASI instrument



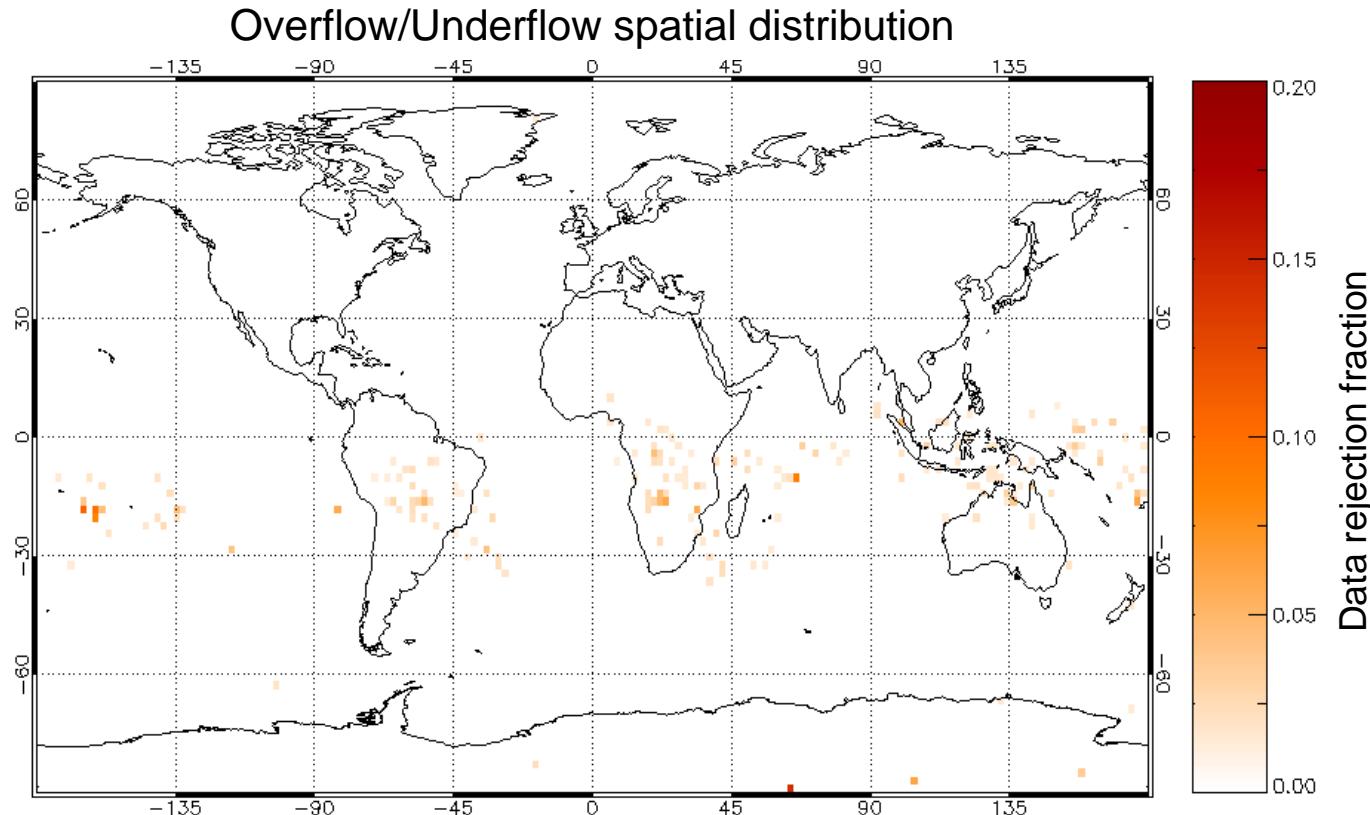
Ground Segment – L1 processing chain



Coding table

Goal : optimize digitalization noise = f (wavenumber), avoid saturation

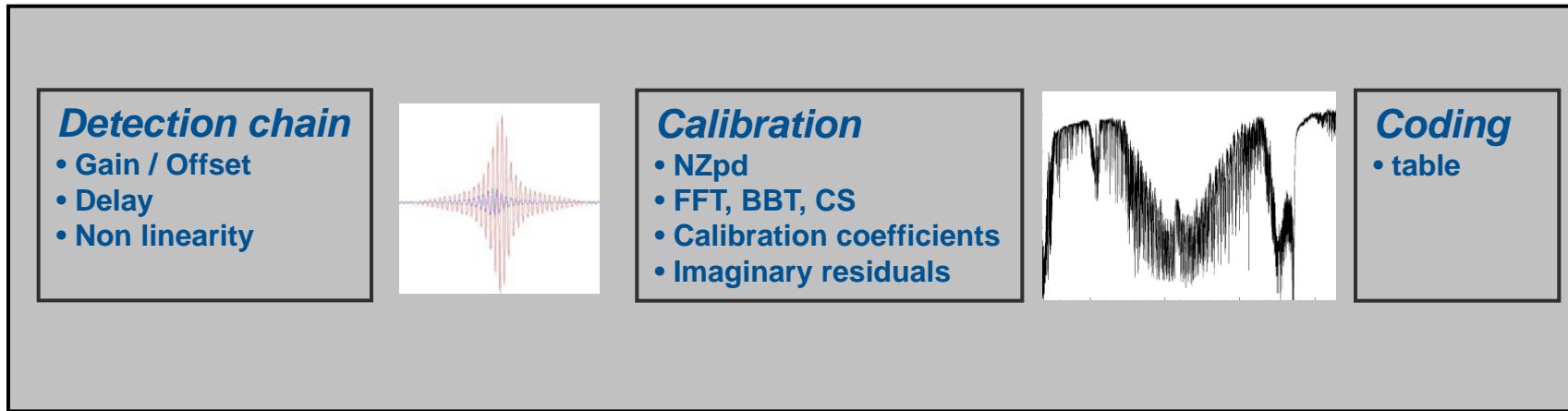
Constraints : available bandwidth



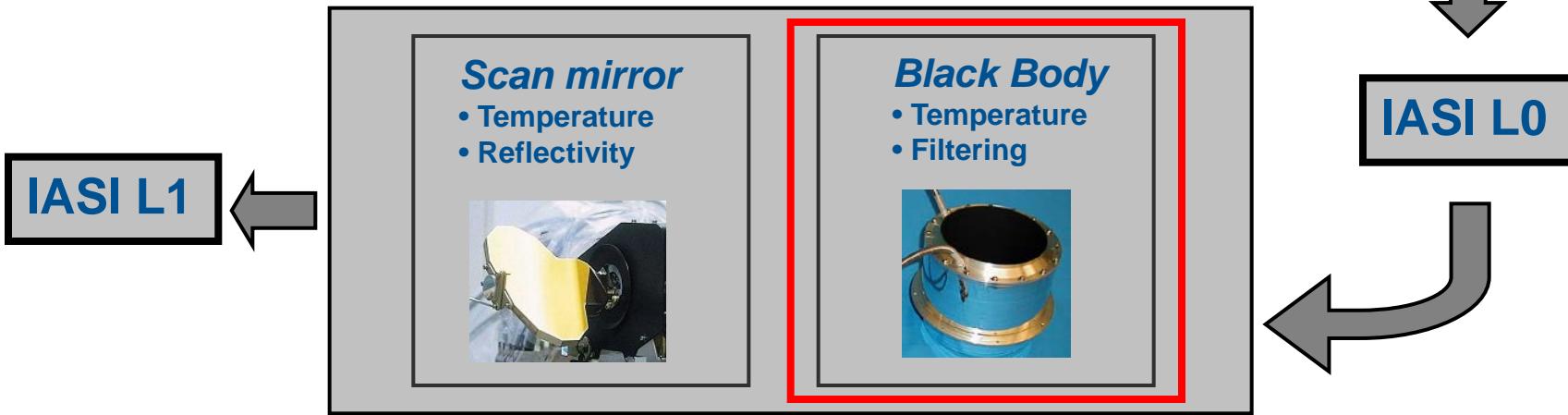
Sun reflection in high thick clouds in tropical region
Overall data rejection < 0.02%

OUTLINE

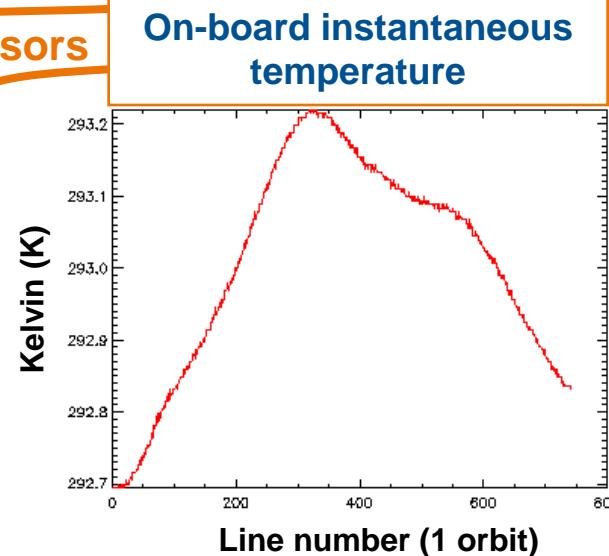
Space Segment – IASI instrument



Ground Segment – L1 processing chain



Black Body

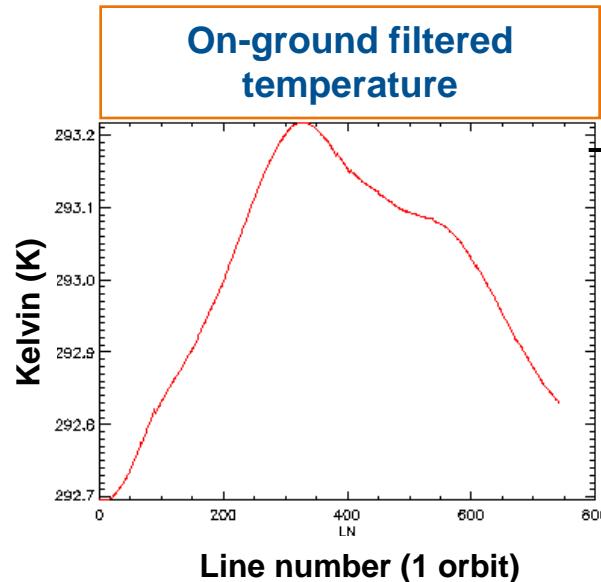


Orbital variation
~ 0.5 K

IASI instrument

L1 processing chain

Filtering



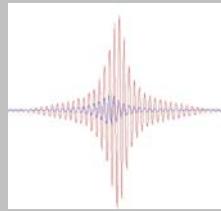
Radiometric post-calibration

OUTLINE

Space Segment – IASI instrument

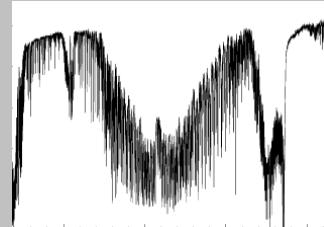
Detection chain

- Gain / Offset
- Delay
- Non linearity



Calibration

- NZpd
- FFT, BBT, CS
- Calibration coefficients
- Imaginary residuals



Coding

- table

Ground Segment – L1 processing chain

IASI L1

Scan mirror

- Temperature
- Reflectivity



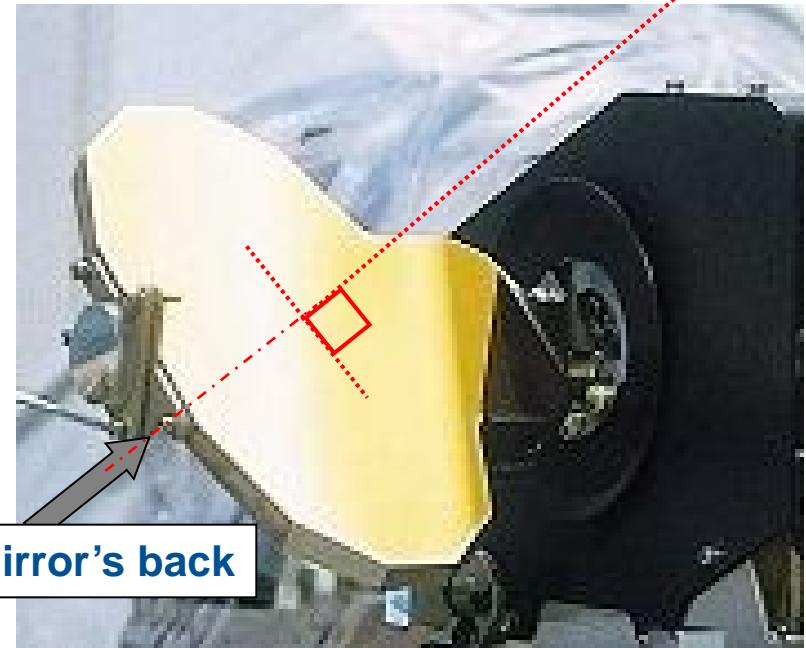
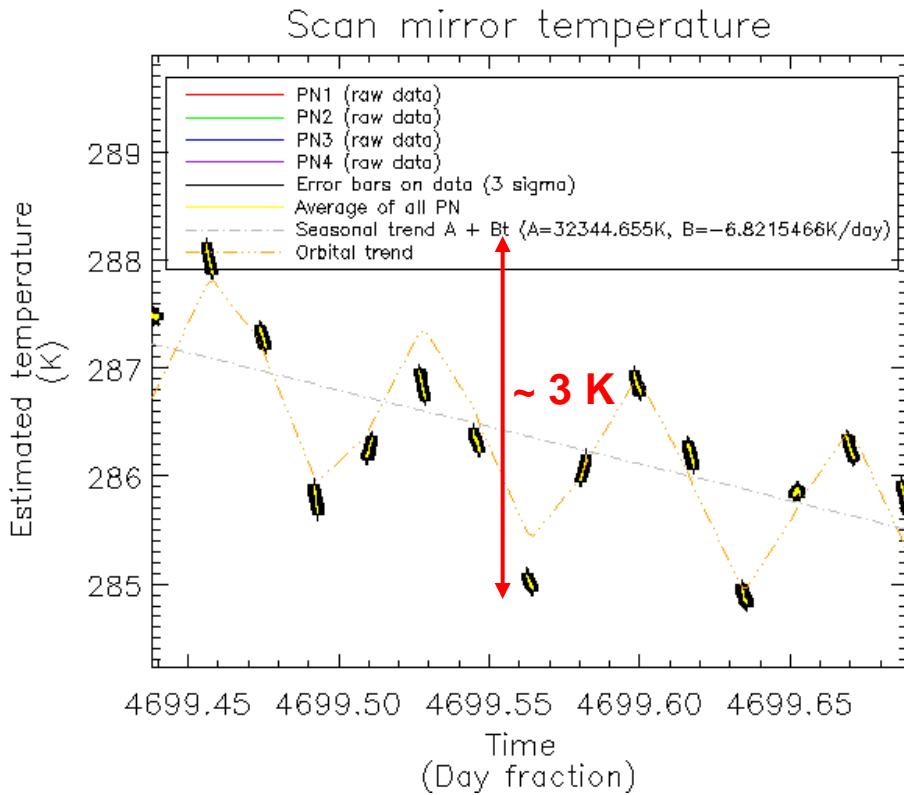
Black Body

- Temperature
- Filtering



IASI L0

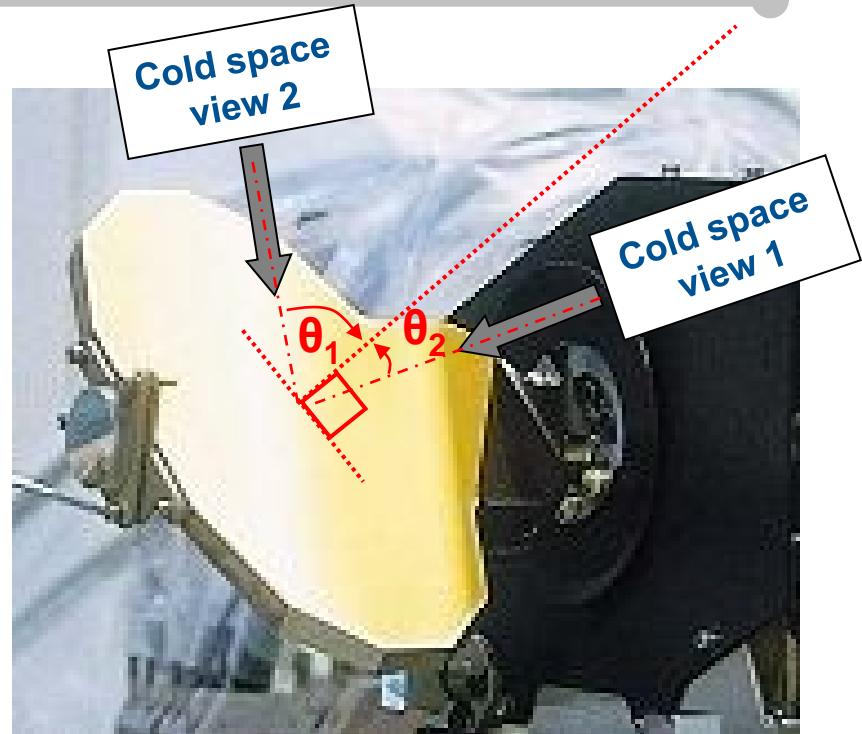
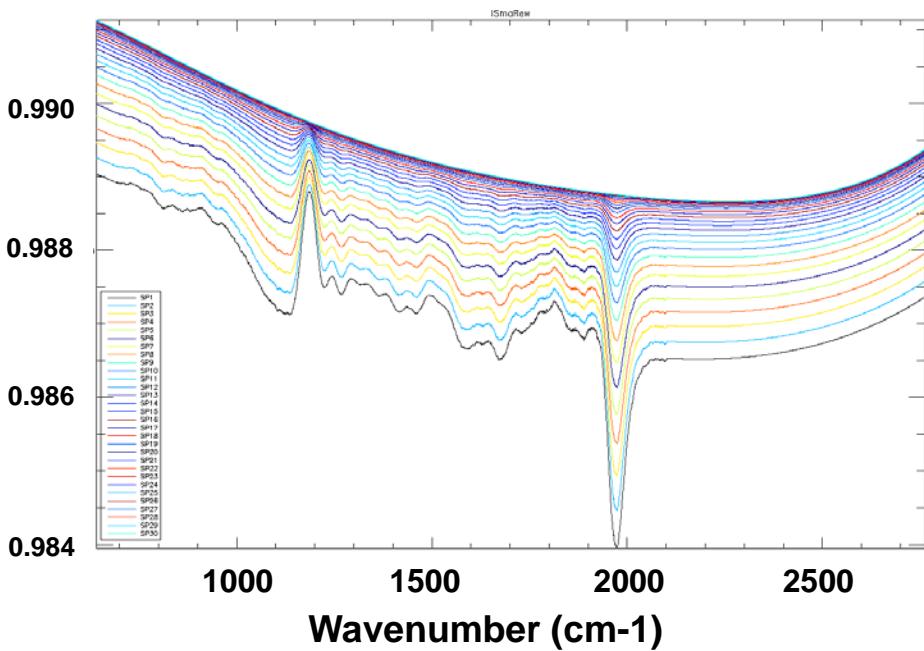
Scan Mirror : Temperature



- Simulation demonstrates that T_{scan} variations of 3 K implies < 0.01 K on the radiometric calibration
- => No need of orbital model, use mean temperature, $T_{\text{scan}}=286.5$ K, for post radiometric calibration

Scan Mirror : Reflectivity

Reflectivity = f (wn, scan position)



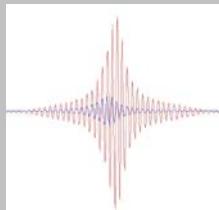
- External calibration targeting CS1 and CS2 allows to model the scan mirror reflectivity = f (wn, SP), part of L1 processing chain parameters
- reflectivity slowly evolves with time, monitored every month.
- updated ~ 1 time per year when impact on radiometric calibration is too high (> 0.1 K)

OUTLINE

Space Segment – IASI instrument

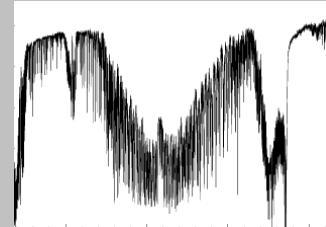
Detection chain

- Gain / Offset
- Delay
- Non linearity



Calibration

- NZpd
- FFT, BBT, CS
- Calibration coefficients
- Imaginary residuals



Coding

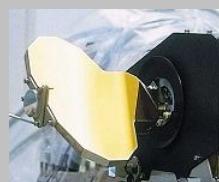
- table

Ground Segment – L1 processing chain

IASI L1

Scan mirror

- Temperature
- Reflectivity



Black Body

- Temperature
- Filtering

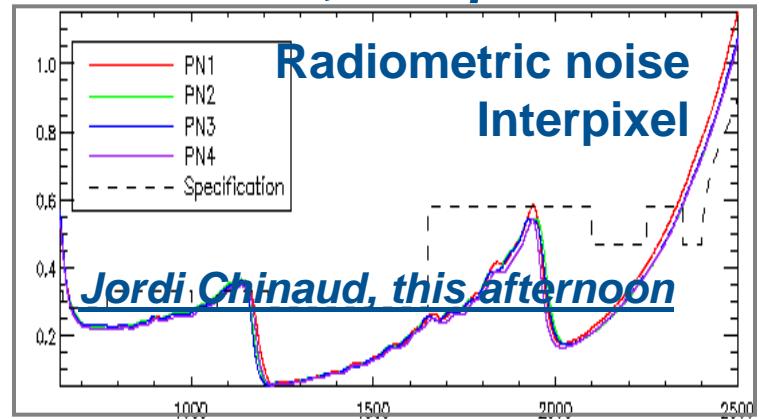


IASI L0

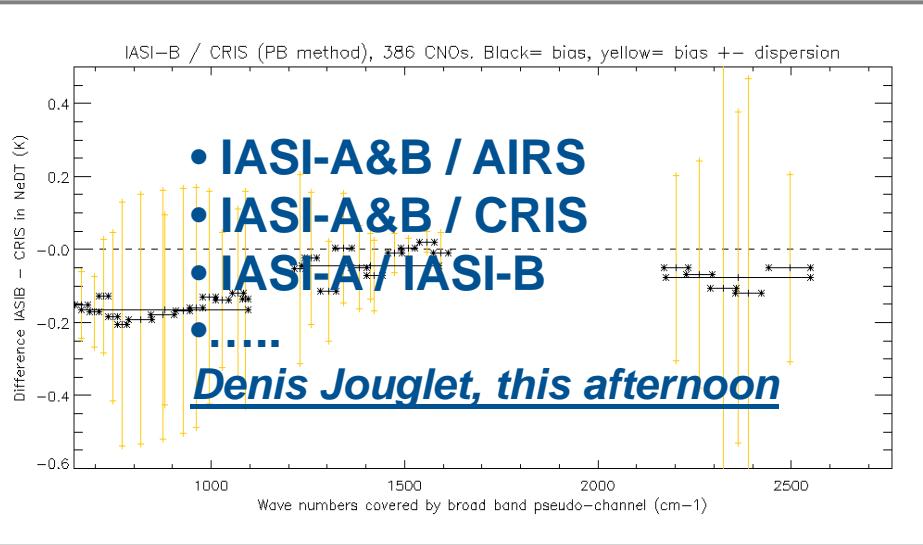
Conclusions

- All parts of the radiometric calibration chain behave nominally
- Data availability improvement / IASI-Metop-A

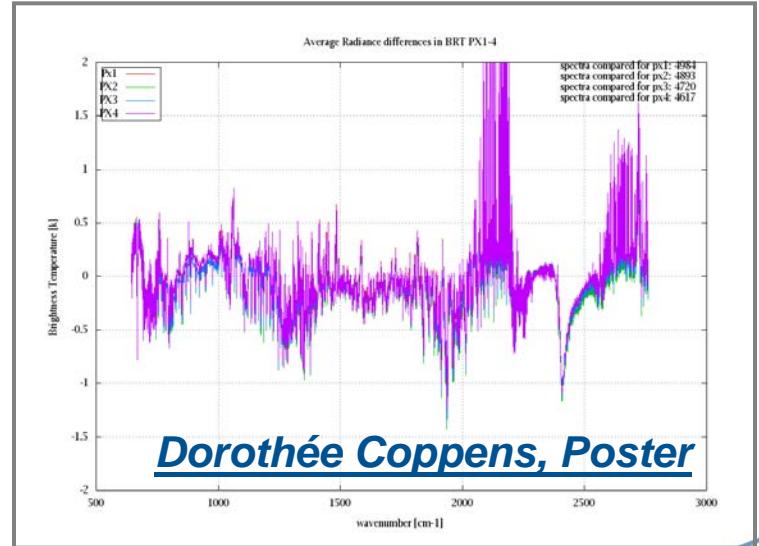
Noise, interpixel



Intercalibration



Obs - Calc.

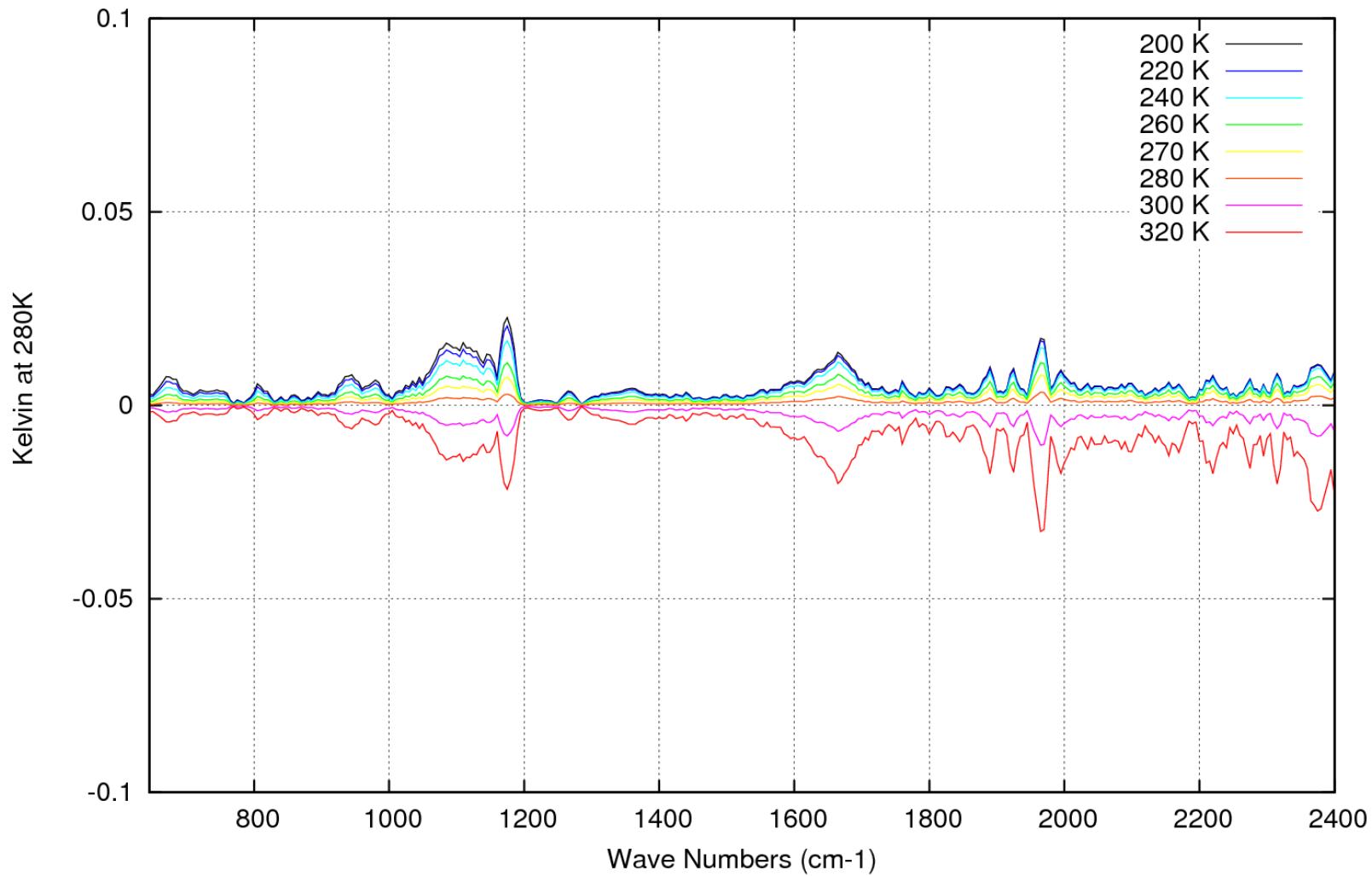




BACKUP SLIDES

Scan Mirror

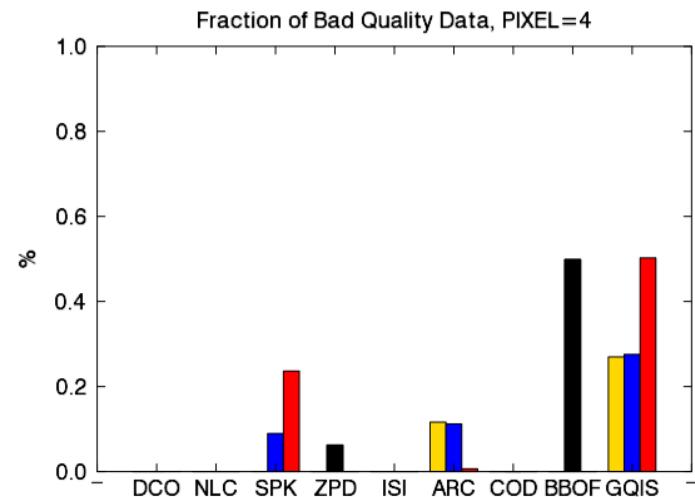
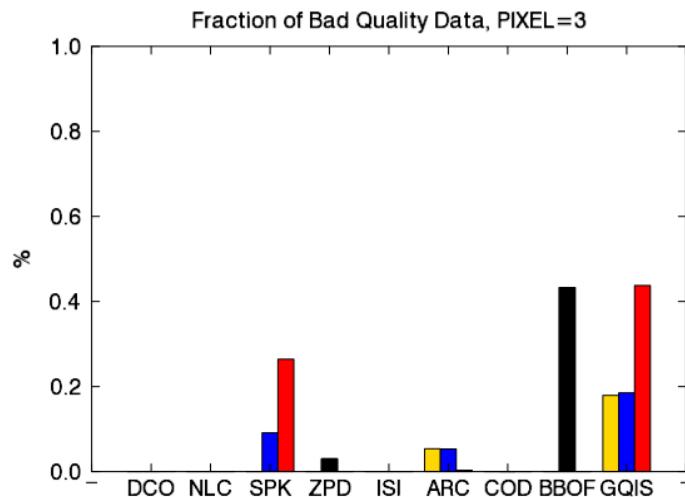
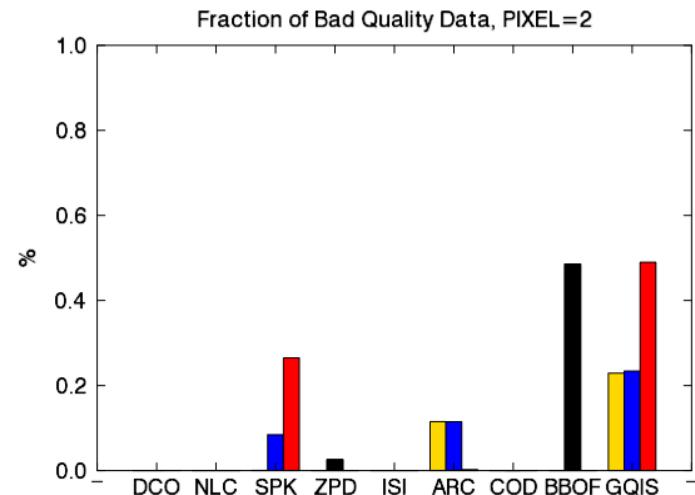
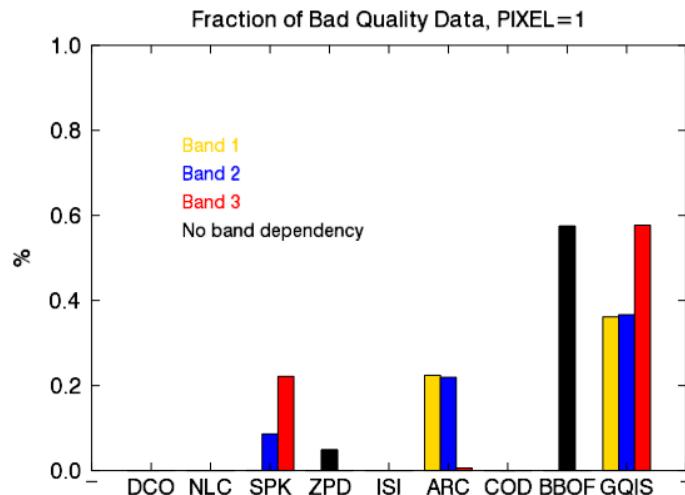
Maximum radiometric calibration error for different scene temperature (Ref : May 12)



28-01-13 14:31 IASI PFMR ROUTINE

CNES DCT/SI/MO

On-board processing quality monitoring



Coding table

