





Doc n°: IA-RP-2000-4263-CNE  
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Sheet: 1 Of: 57

**IASI QUARTERLY PERFORMANCE REPORT**  
**FROM 2015/12/01 TO 2016/02/29**

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**BY IASI TEC (TECHNICAL EXPERTISE CENTER)**  
**FOR IASI FM2 ON METOP A**



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<p align="center"><b>ANALYSE DOCUMENTAIRE</b></p> <p align="center"><b>Bordereau d'indexation</b></p>
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IASI TEC periodic report			
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IASI quarterly performance report			
<b>RESUME :</b>			
Quarterly report issued by the IASI TEC team to show trends and layout from the "long term synthesis" TEC function for flags and observables quality indicators			
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**LIST OF ACRONYMS**

4A/OP	Automatized Atmospheric Absorptions Atlas/ Operational
APO	Other Parameters OPS
AR	Anomaly Report
ASE	Acquisition Start End
AVHRR	Advanced Very High Resolution Radiometer
BB	Black Body
BRD	BoaRD configuration
CCFD	Cube Corner Functional Device
CCD	Cube Corner Direction
CCM	Cube Corner Mechanism
CD	Cube corner Compensation Device
CHART	Component Health Assessment and Reporting Tool
CGS	Core Ground Segment at EUMETSAT
CNES	Centre National d'Etudes Spatiales
CS	Cold Space
DA	Applicable document
DPS	Data Processing Subsystem
ECMWF	European Centre for Medium Range Weather Forecasts
EM	Engineering Model
EPS	EUMETSAT Polar System
EUMETSAT	European organisation for exploitation of METeorological SATellites
FM2 / FM3	Flight Model n°2 or 3
FOV	Field Of View
GRD	GRound configuration
IASI	Infrared Atmospheric Sounding Interferometer
IIS	Integrated Imaging Subsystem
IPSF	Instrument Point Spread Function
ISRF	Instrument Spectral Response Function
LFD	Locking Filtering Device
LN	Line Number
LSB	Least Significant Bit
METOP	METeorological OPERational satellite
MPF	Mission Planning Facility
NedT	Noise equivalent difference Temperature
NDVI	Normalized Difference Vegetation Index
NZPD	Number of Zero Path Difference
ODB	Operational Data Base
OPS	Operational Software
PC	Principal Component
PDD	Position Data Diagnostic

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PDU	Product Dissemination Unit
PL SOL	PayLoad Switch Off-Line
PN	Pixel Number
PTSI	Parameter Table Status Identifier
RMS	Root Mean Square
RD	Reference Document
SAA	South Atlantic Anomaly
SEU	Single Event Upset
TEC	IASI Technical Centre of Expertise (located in CNES, Toulouse)
TIGR	Thermodynamic Initial Guess Retrieval data set
VDS	Verification Data Selection
ZPD	Zero Path Difference



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## 1 INTRODUCTION



The IASI TEC is based at CNES Toulouse and is responsible for the monitoring of the IASI system performances, covering both instrument and level 1 processing sub-systems.

This document describes the activities and results obtained at the IASI TEC for instrument FM2 on METOP-A during the following period:

- Start Time: 2015/12/01                      Orbit: 47299
- End Time: 2016/02/29                      Orbit: 48591
- Duration: 3 months

Note that IASI ended the Calibration / Validation (commissioning) phase on July 2007.

## 2 RELATED DOCUMENTS

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### 3 SIGNIFICANT EVENTS

The following tables present a timeline of the various requests sent by TEC and the external IASI activities.

Those events are typically the configuration changes, programming requests, software update, but also any external operation or activity such as mission interruption, manoeuvre, dissemination problem, ...

#### 3.1 EXTERNAL CALIBRATION

Table 1 shows the External Calibration within the time period reported here. Note that the VDS files that come with each request are not described here.

Execution	TEC ref. <sup>(1)</sup>	Description	Activities
03/12/2015 from 5h13 to 9h09 orb. 47330 to 47332	RM-102	Monthly_MPF <sup>(2)</sup> Targets: Earth 15, Blackbody, 2 <sup>nd</sup> Deep Space, Mirror Backside	For routine monitoring (IIS and IASI NeDT, scan mirror reflectivity, ghost,...)
From 29/12/2015 00h24 to 30/12/2015 08h20 orb. 47697 to 47703 orb. 47709 to 47715	RL-28	Moon avoidance MPF <sup>(2)</sup> Targets: 1 <sup>st</sup> Deep Space	Monitoring of moon intrusion in CS1 FOV
01/01/2016 from 5h13 to 9h09 orb. 47742 to 47744	RM-103	Monthly_MPF <sup>(2)</sup> Targets: Earth 15, Blackbody, 2 <sup>nd</sup> Deep Space, Mirror Backside	For routine monitoring (IIS and IASI NeDT, scan mirror reflectivity, ghost,...)
30/01/2016 from 5h13 to 9h09 orb. 48154 to 48156	RM-104		
28/02/2016 from 5h13 to 9h09 orb. 48566 to 48568	RM-105		

Table 1: External Calibration TEC Requests



<sup>(1)</sup> TEC convention: R for Routine, M for Monthly and L for moon avoidance, followed by a chronological number

<sup>(2)</sup> An external calibration could be the result of:

- a TEC request or
- a “MPF” uploaded directly by EUMETSAT in full accordance with TEC. The reference “Monthly\_MPF” is based on the March 2008 TEC External Calibration request. The MPF for moon avoidance is based on the December 2008 TEC External Calibration request: “ICAL\_OCF\_xx\_M02\_20081216060000Z\_20090616060000Z\_20081209100934Z\_IAS\_EXTALIBRA.dts”

Moon external calibration on December 29<sup>th</sup> 0:24<sup>Z</sup> to 30<sup>th</sup> 8:20<sup>Z</sup> (orbits 47697 to 47703 and 47709 to 47715) detail:

External Calibration		External Calibration	
from	to	from	to
2015/12/29 00:24:12	00:47:07	2015/12/29 21:44:43	22:07:55
2015/12/29 02:04:11	02:30:51	2015/12/29 23:25:47	23:54:51
2015/12/29 03:44:11	04:16:27	2015/12/30 01:06:35	01:39:23
2015/12/29 05:23:23	05:58:03	2015/12/30 02:47:39	03:19:23
2015/12/29 07:05:31	07:38:51	2015/12/30 04:28:43	04:59:39
2015/12/29 08:49:15	09:19:23	2015/12/30 06:12:11	06:39:55
2015/12/29 10:34:03	10:59:39	2015/12/30 07:57:15	08:20:27

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### 3.2 ON BOARD CONFIGURATION

Table 2 presents the on-board processing configuration updates that had been made within the time period reported here:

PTSI	IASI on board parameter files	Delivery by TEC	activated on	TEC ref.	affected parameters of a DPS TOP configuration update

Table 2: DPS and MAS configuration TEC Requests

For information, Table 3 shows the delivery applicable at the beginning of the period:

PTSI	IASI on board parameter files	Delivery by TEC	activated on	TEC ref.	affected parameters of a DPS TOP configuration update
<b>15</b> 2.0	IDPS_OBP_xx_M02_20150717000000Z_20160117000000Z_20150716125458Z_IAS_T_DPSPARAMOD.tar	16/07/2015 12h54	05/08/2015 13:08 orbit 45630	R_60	Update of reduced spectra + ZpdQualIndexCutOff BB & CS

Table 3: DPS and MAS previous configuration

The associated ground configuration table (BRD file), necessary to handle coherent configuration at system level, is presented in the next section. These associated configuration table are necessary for L1 processing.

### 3.3 GROUND CONFIGURATIONS UPDATES FOR LEVEL 1 PROCESSING

Table 4 presents the on-ground processing configuration updates that had been made within the time period reported here:

IDef	IASI L1 auxiliary files	Delivery by TEC	Upload on GS1	Content

Table 4: IASI L1 Auxiliary File Configuration on the Operational EPS Ground Segment

For information, Table 5 shows the delivery applicable at the beginning of the period:

IDef	IASI L1 auxiliary files	Delivery by TEC	Upload on GS1	Content
<b>61</b>	IASI_BRD_xx_M02_20150918000000Z_XXXXXXXXXXXXZ_20150917130533Z_IAS_T_0000000015	17/09/2015 13h05	BRD activated on 26/10/2015 14:14 orbit 4676	Update of IIS/AVHRR Offset
<b>31</b>	IASI_GRD_xx_M02_20150522000000Z_XXXXXXXXXXXXZ_20150521083343Z_IAS_T_0000000031	21/05/2015 08h33	GRD activated on 17/06/2014 06:53 orbit 39745	Update of Scan Mirror Reflectivity (+ IdefOffsetSondIISModel )
<b>14</b>	IASI_ODB_xx_M02_20130417080000Z_XXXXXXXXXXXXZ_20130417074829Z_IAS_T_0000000014	17/04/2013 07h48	16/05/2013 07:08 orbit 34105	

Table 5: IASI L1 auxiliary file previous configuration



### 3.4 DATA BASES UPDATE FOR THE USERS

The Noise Covariance Matrix (NCM) and Spectral data base (SDB) are specific data bases for the users. They are updated according to the main ground level 1 evolutions.

Table 6 presents the updates of the NCM and SDB that had been made within the time period reported here:

IDef	Users Data-Base	Delivery by TEC	TEC ref.	Comments

Table 6: IASI Data Bases for the users

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For information, Table 7 shows the delivery applicable at the beginning of the period:

IDef	Users Data-Base	Delivery by TEC	TEC ref.	Comments
5	IASI_NCM_xx_M02_20141028000000Z_20141028000000Z_20141027140151Z_IA_ST_SPECTRESPO	27/10/2014	R_COV_5	Covariance matrix from L0 noise on BB (2014/09/24external calibration)
14	IASI_SDB_xx_M02_20130923140000Z_20130923140000Z_20130923124758Z_IASIST_IASISPECDB	23/09/2013	R_57	User database associated to ODB IDefSDB 14

Table 7: previous IASI Data Bases

### 3.5 ON GROUND HW/SW EVOLUTION

Table 8 presents the updates of PPF L1 software within the time period reported here:

IASI L1 PPF software version	Delivery by TEC	Date introduced on GS1	Comments
7.4	13/10/2015		AIX 7.1 upgrade

Table 8: IASI L1 PPF Configuration on the Operational EPS Ground Segment

For information, Table 9 shows the software version applicable at the beginning of the period:

IASI L1 PPF software version	Delivery by TEC	Date introduced on GS1	Comments
7.3	08/2015	29/09/2015 for sensing time 13:32 <sup>UTC</sup> Orbit 46412	CSQ impact on IIS

Table 9: Previous IASI L1 PPF

### 3.6 DECONTAMINATION

Table 10 presents decontaminations that have been made or requested within the time period reported here:



Last due date	Date of decontamination	Description

Table 10: Decontamination TEC Requests

For information, Table 11 shows the previous decontamination:

Last due date	Date of decontamination	Description
In 2015	From 8 to 13/09/2014	

Table 11: Previous decontamination

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### 3.7 INSTRUMENT

#### 3.7.1 External events

This category is for those activities/events that are external to IASI but still have an impact. It is broken down into classes of *PL-SOL* and *OOP* manoeuvre.

##### 3.7.1.1 Manoeuvres

Date	Type <sup>(*)</sup>	Description	IP flag	OoP mission Outage
2016/02/10	IP	IP manoeuvre #39 (orbit 48319) – collision avoidance		

Table 12: Overview of METOP manoeuvres in the reporting period

(\*): IP for In-Plane manoeuvres (IASI stays in NOP) and OoP for Out of plane manoeuvres (IASI is put in Heater 2)

##### 3.7.1.2 PL-SOL

Table 13 presents the PL-SOL events that have occurred within the time period reported here:

Dates	Orbits	Description

Table 13: PL-SOL

#### 3.7.2 Operation leading to mission outage

This chapter presents the intervention on IASI needing routine interruption that have occurred within the time period reported here.

Dates	Orbits	type	IASI mode	Description
14/01/2016	47932			IIS equalization

Table 14: Scheduled interruptions

#### 3.7.3 Anomaly leading to mission outage

Table 15 and Table 16 present the major and minor anomalies internal to IASI that have occurred within the time period reported here.

Note that, in this section minor anomalies are all identified and without any impact on the mission, and major anomalies only affect IASI instrument, and no other sub-systems of the spacecraft.



Dates	Orbits	Anomaly type <sup>(*)</sup>	IASI mode	Description
2016/01/08 13:48 to 17:48	47847 to 47849	SET	Heater Refuse	SET in DPS

Table 15: Major anomalies

(\*): SEU (LAS, CCM or DPS) anomalies or SET anomalies

Day	Orbits	error n°	Severity	Anomaly type	LN	SN	Description

Table 16: Minor anomalies

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## 4 PERFORMANCE MONITORING

### 4.1 *PERFORMANCE MONITORING*

In order to ensure that the IASI system is permanently running in good conditions, the CNES (IASI TEC) and EUMETSAT (CGS) are monitoring products at various temporal levels : at line, PDUs and DUMP (full orbit).

The on-board and ground processing performance algorithms issue more than one hundred quality indicators, called flags and simple parameters. Those are alarms for any bad functioning or local performance degradation.

According to the results, the TEC is also in charge of delivering new on-board or ground parameters to EUMETSAT when it is necessary. EUMETSAT is then in charge of uploading them on-board or as an input of the level 1 processing chain. During the whole instrument life, these parameter adjustments are necessary in order to take into account instrument evolution in the processing and finally to maintain a good data quality.

The Table 17 is the colour code used for the status report.

Status Colour	Meaning
GREEN	$\geq 95\%$ of good spectra
YELLOW	$< 95\%$ of good spectra
RED	Production interrupted
BLANK	No Status Reported

*Table 17: Functional status legend*

## 4.2 PERFORMANCE SYNOPSIS

Table 18 provides a synthetic view of all the indicators evaluated for L0/L1 data and their current status.

Section	Component	Description	Status	Comments
<a href="#">4.3</a>	L0	<b>Level-0 Data Quality</b> <ul style="list-style-type: none"> <li>Overall quality</li> <li>Main flag and quality indicator parameters <ul style="list-style-type: none"> <li>Spikes monitoring</li> <li>ZPD monitoring</li> <li>Overflows/Underflows monitoring</li> <li>Reduced Spectra monitoring</li> </ul> </li> <li>Second level flag and quality indicators</li> </ul>	GREEN	On-board processing
<a href="#">4.4</a>	L1	<b>Level-1 Data Quality</b> <ul style="list-style-type: none"> <li>Overall</li> <li>Main flag and quality indicator parameters</li> </ul>	GREEN	On ground processing
<a href="#">4.5</a>	L1	<b>Sounder radiometric performances</b> <ul style="list-style-type: none"> <li>Radiometric noise</li> <li>Radiometric calibration</li> <li>Acquisition chain delay</li> <li>Optical transmission <ul style="list-style-type: none"> <li>Ice</li> <li>Prediction of decontamination date</li> </ul> </li> <li>Interferometric contrast</li> <li>Interferogram baseline</li> <li>Detection chain</li> </ul>	GREEN	
<a href="#">4.6</a>	L1	<b>Sounder spectral performances</b> <ul style="list-style-type: none"> <li>Dimensional stability <ul style="list-style-type: none"> <li>Position of axis</li> <li>Cube Corner constant offset</li> <li>Cube Corner velocity</li> <li>Optical bench temperature</li> </ul> </li> <li>Spectral calibration</li> </ul>	GREEN	
<a href="#">4.7</a>	L1	<b>Geometric performances</b> <ul style="list-style-type: none"> <li>Sounder/IIS co-registration</li> <li>IIS/AVHRR co registration</li> </ul>	GREEN	
<a href="#">4.8</a>	L1	<b>IASI-A inter-calibration</b> <ul style="list-style-type: none"> <li>IASI-A inter-calibration with CRIS</li> <li>IASI-A inter-calibration with AIRS</li> </ul>	GREEN	
<a href="#">4.9</a>	L1	<b>IIS radiometric performances</b> <ul style="list-style-type: none"> <li>IIS radiometric noise monitoring</li> <li>IIS radiometric calibration monitoring</li> </ul>	GREEN	

Table 18: IASI product components functional status

### 4.3 LEVEL 0 DATA QUALITY (L0)

#### 4.3.1 Overall quality

The IASI L0 data quality (orbit average) through IASI engineering products is shown in Figure 1.

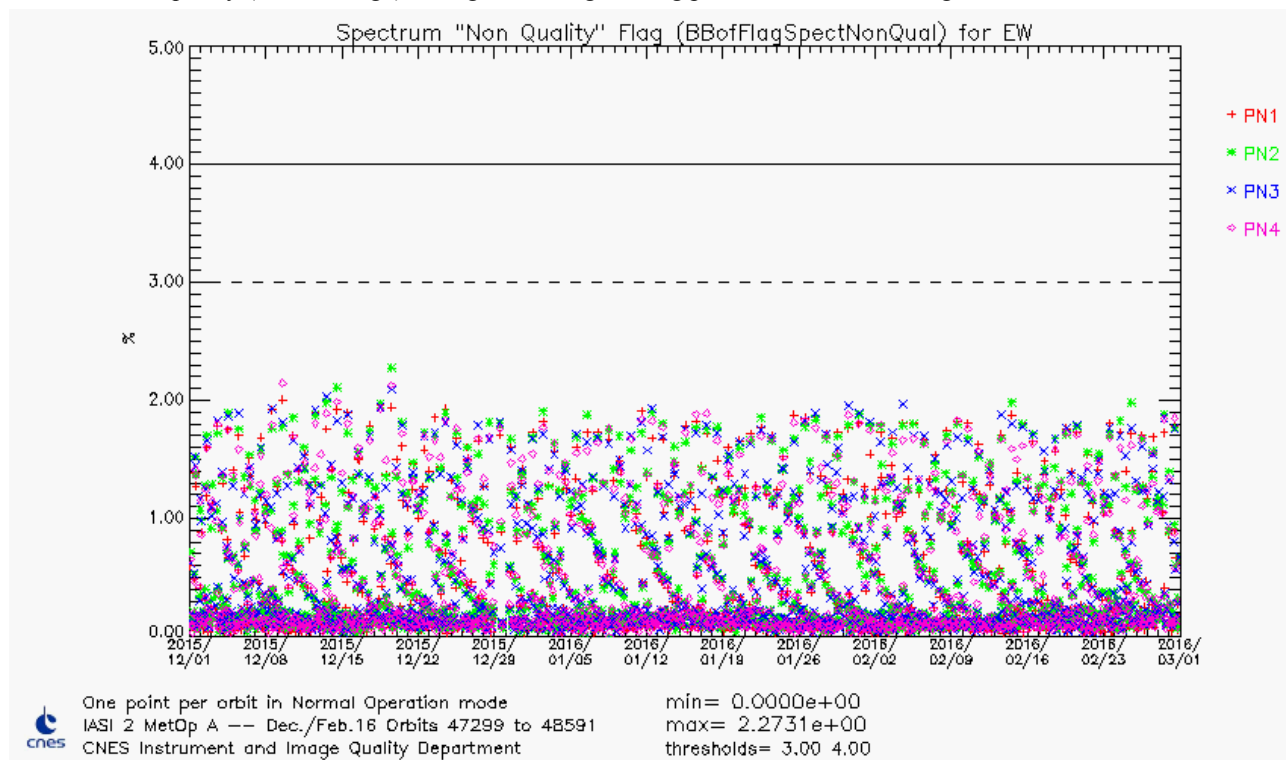


Figure 1 : IASI L0 data quality orbit average (per pixel and CCD)

The geographical distribution of the overall L0 (board) quality flag for the 4 pixels is shown in Figure 2.



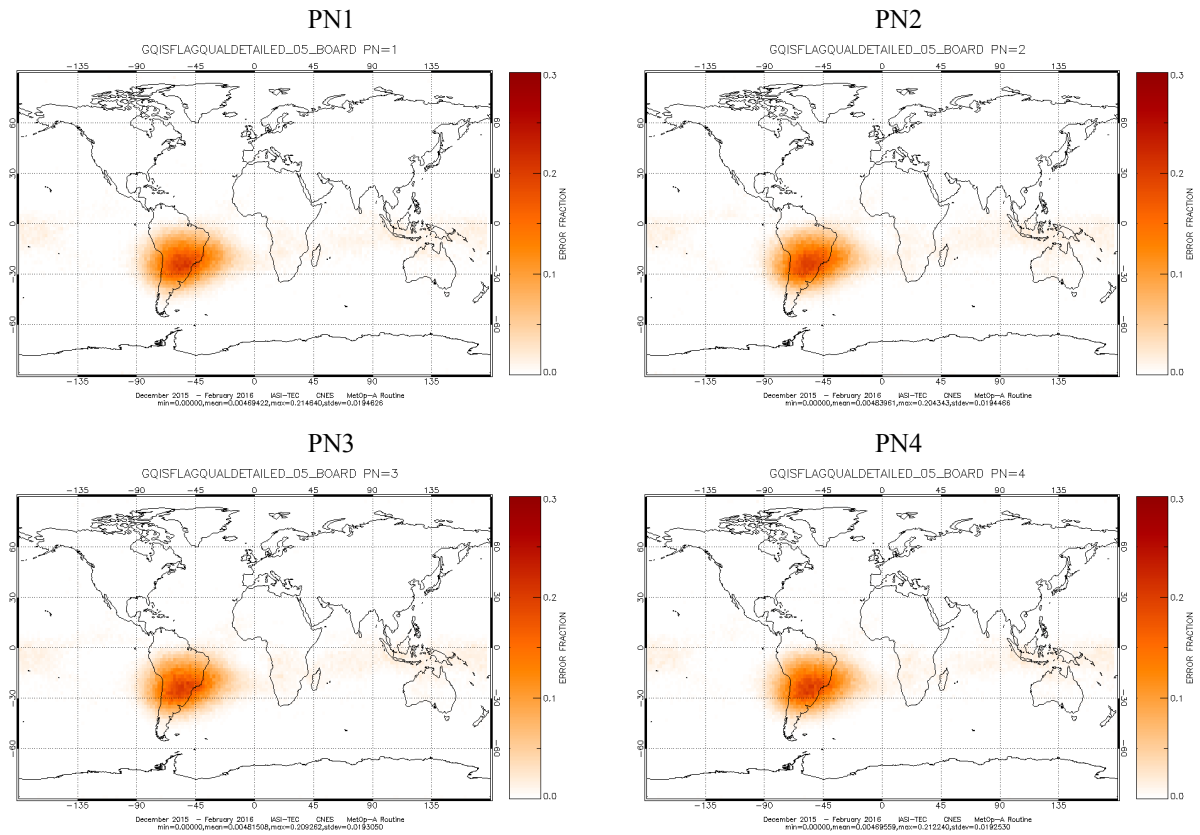




Figure 2 : IASI L0 data quality spatial distribution (per pixel)

The IASI L0 quality and on-board processing are nominal.

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### 4.3.2 Main flag and quality indicator parameters

The main contributors to the rejected spectra by on-board processing are: spikes (proton interaction on detectors), failure of NZPD algorithm determination and over/underflows (measured data exceeding on-board coding tables capacity). They are analysed in details hereafter.

#### 4.3.2.1 Spikes monitoring

Spikes occur when a proton hits a detector. This very high energetic particle disrupts the measure of the interferogram and then corrupts the spectrum.

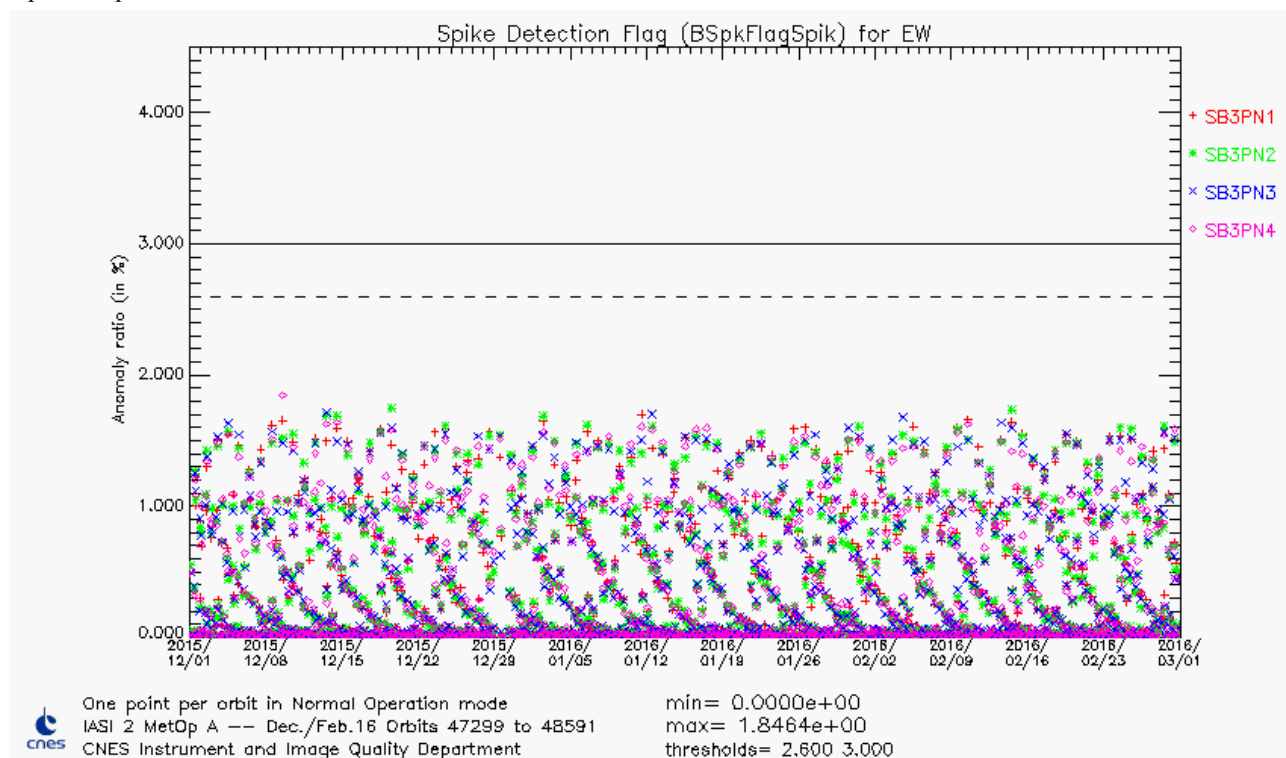


Figure 3 : Temporal evolution of spikes anomaly ratio in % for all pixels (orbit average)

An example of the geographical distribution of spikes occurrences on band 3 for the 4 pixels is shown in Figure 4.



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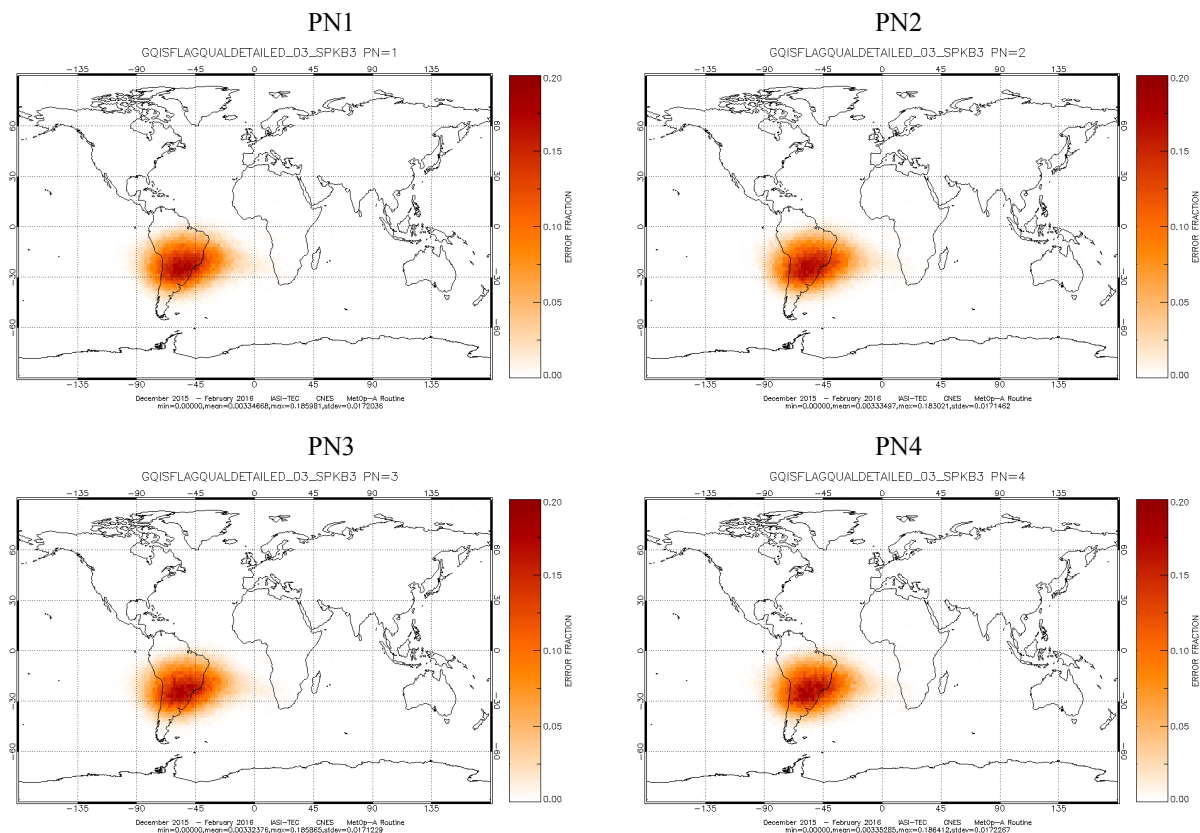




Figure 4 : Geographical distribution of spikes occurrences in % for band 3 and all pixels

Spikes are mainly located in the regions of Earth where the magnetic field doesn't protect the satellite from the energetic particles: the poles and the SAA (South Atlantic anomaly).

Spike anomaly ratio is nominal for the reported period.

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#### 4.3.2.2 ZPD monitoring

The ZPD (“Zero Path Difference”) is the position of the central fringe of the interferogram. The NZPD is the number of the sample detected as the ZPD. On IASI, it is determined by a software. This is a special feature of IASI in comparison to other instruments for which NZPD determination is done by hardware.

NZPD variations are governed by two phenomena:

1. ASE fluctuations which have the same effect on each pixel and can produce NZPD variation of 30-40 samples over month. This is the first order phenomena.
2. Mechanical deformation of the interferometer or evolution of detection chain delays. These phenomena affect the 4 pixels in different way. However this phenomenon has a second order effect in comparison to the first one.

**We monitor both NZPD determination quality flag and interpixel homogeneity. We expect stability.**

BZPDFlagNZPDNonQualEW: Temporal evolution of NZPD determination quality flag for earth view

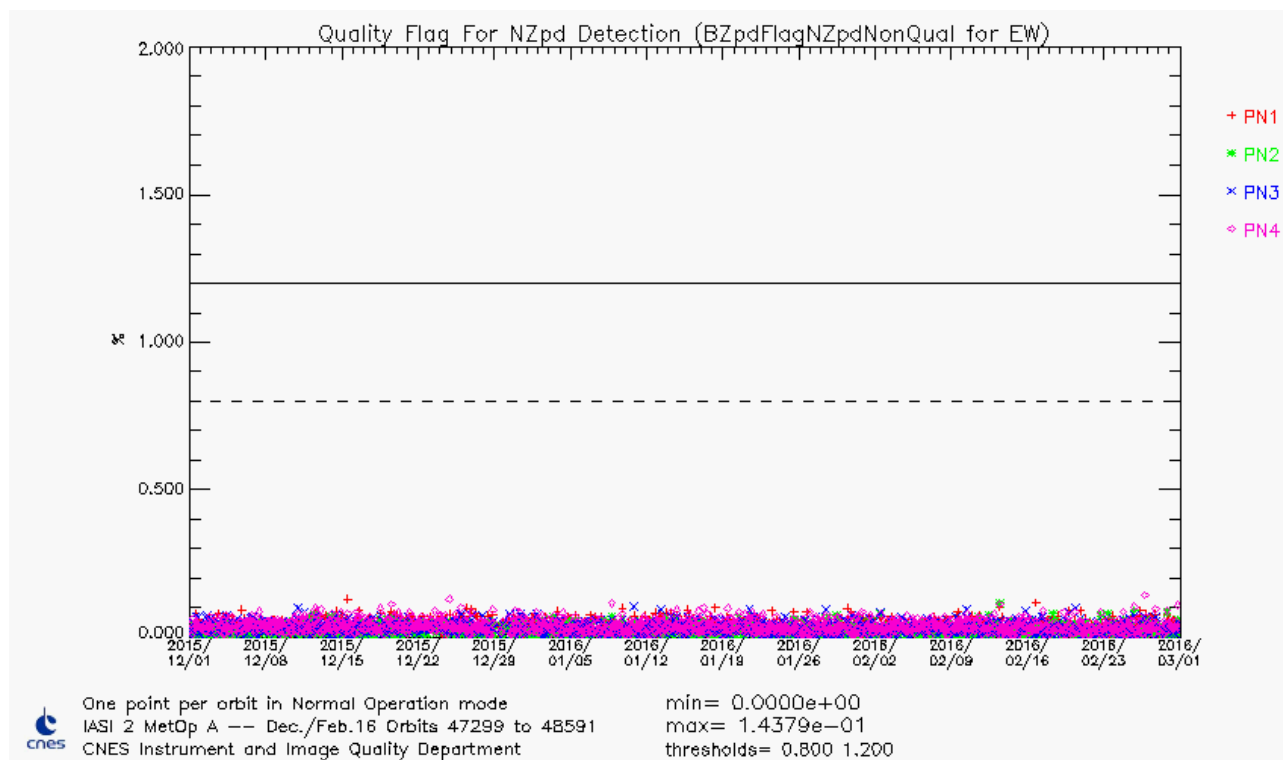


Figure 5 : Temporal evolution of NZPD determination anomaly ratio in % for all pixels (orbit average)

NZPD determination anomaly ratio is nominal for the reported period.

The geographical distribution of the NZPD determination quality flag for the 4 pixels is shown in Figure 6.

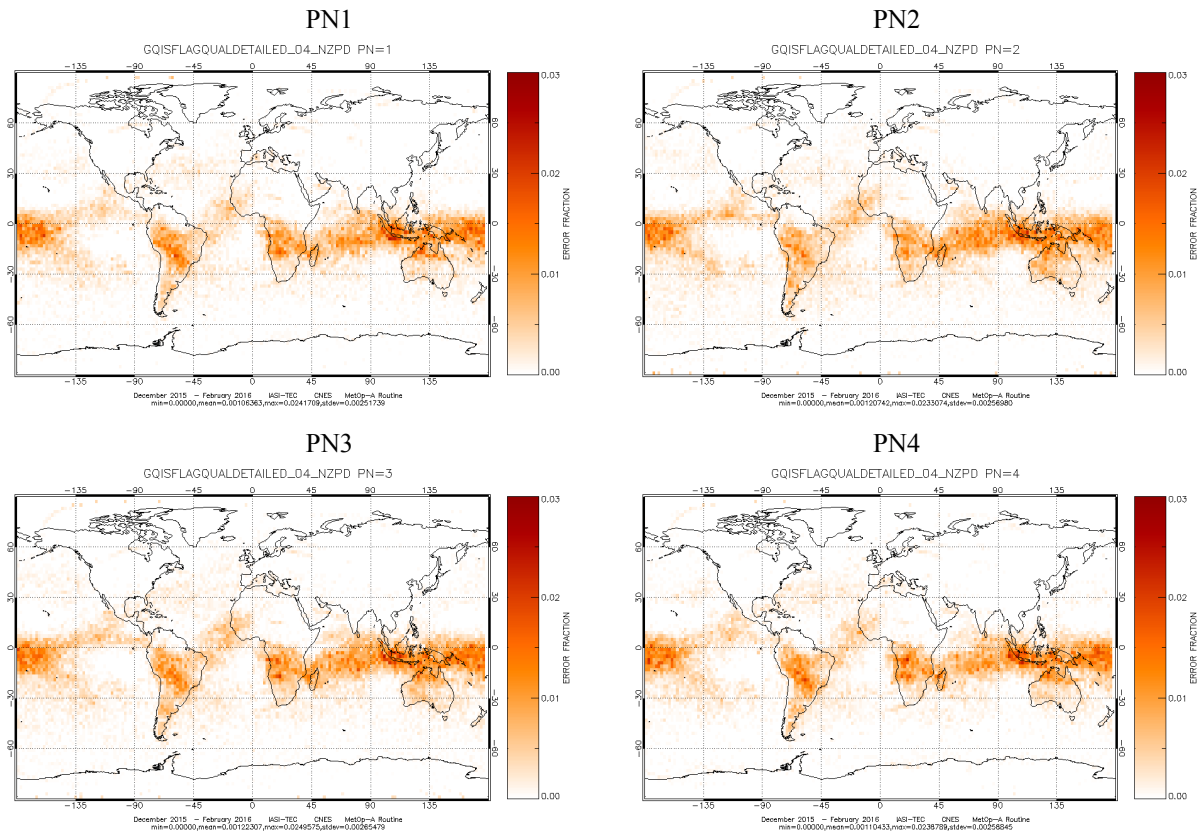


Figure 6 : IASI NZPD determination quality flag spatial distribution (per pixel)

The NZPD determination fails over some clouds that have a temperature that induces no energy in the central fringe of the interferogram.

### NZPD inter-pixel homogeneity monitoring

This monitoring is necessary in order to follow potential deformation of the interferometer or evolution of detection chain delay.

The NZPD inter-pixel homogeneity is nominal over the reported period. Consequently, these parameters are perfectly stable and in-line with the specification.

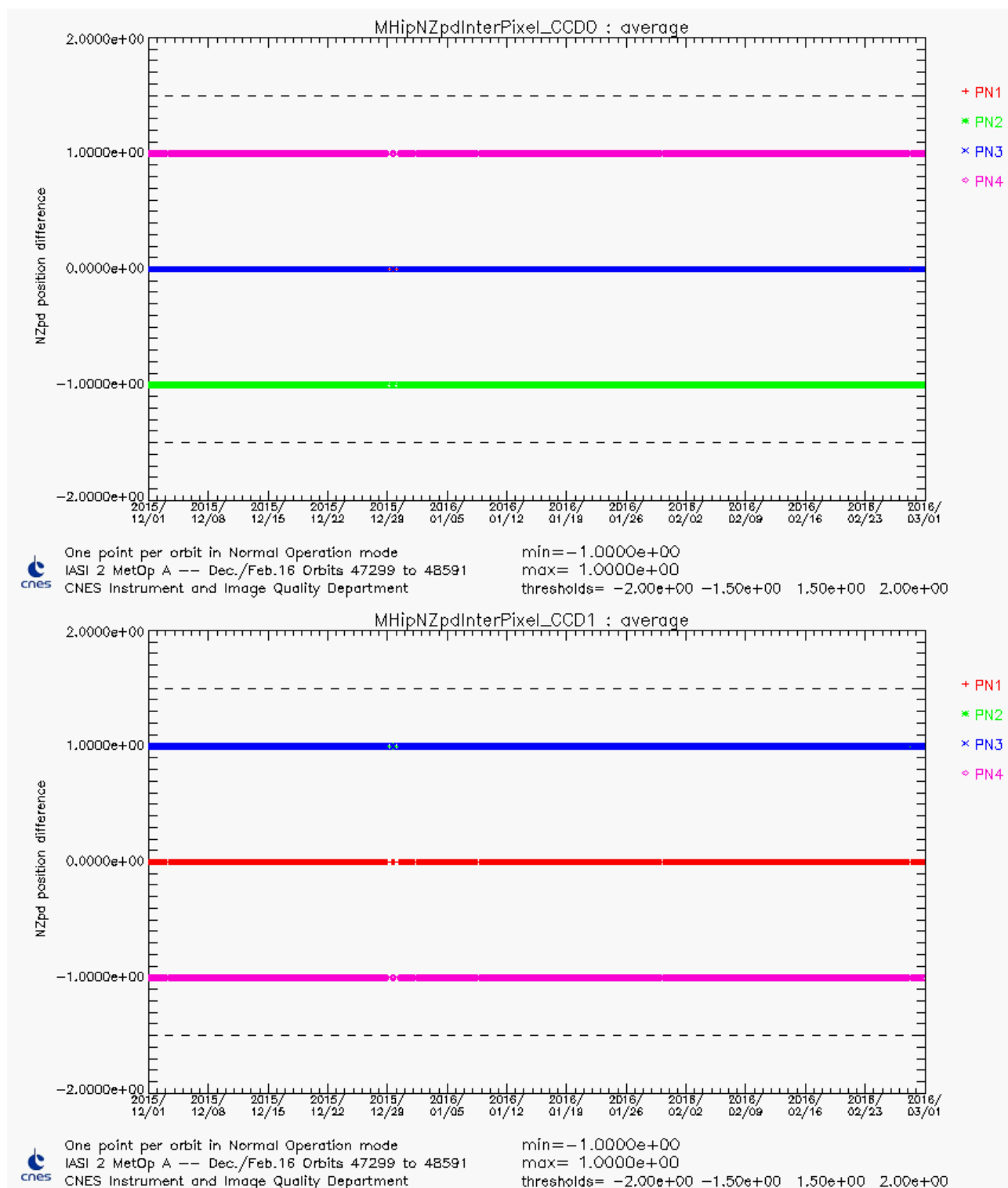




Figure 7 : NZPD inter-pixel for all pixels and CCD calculated with respect to pixel 1 (orbit average)

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#### 4.3.2.3 Overflows / Underflows monitoring

The total number of bits available for a spectrum to be transmitted to the ground is limited. For that reason, we have defined coding tables to encode each measured spectrum. These tables have been designed by using “extreme spectrum” corresponding to known drastic atmospheric conditions. The coding step is also set to not introduce additional noise into the spectrum. However for very extreme atmospheric conditions (sunglint in B3, very high stratospheric temperature...) a measurement can exceed on-board coding tables’ capacity and causes an over/underflow.

Over/underflows occurrences are monitored and stability is expected. As long as they remain to low levels, the coding table is not changed. Note that changing the coding tables requires compromises. Indeed, increasing the encoding capacity can be achieved by two different ways. A first solution consists in an increase of the coding step without changing the number of bits. However, that leads to an increase of the digitalization noise. Then, a second solution consists in keeping the coding step constant while increasing the number of bits available for a particular band. But, the total amount of bits available for the entire spectrum is limited and constant. So, that requires to decrease the encoding capacity in another spectral band.

Time series of Overflows and Underflows (orbit average) are shown in following figure for all pixels.

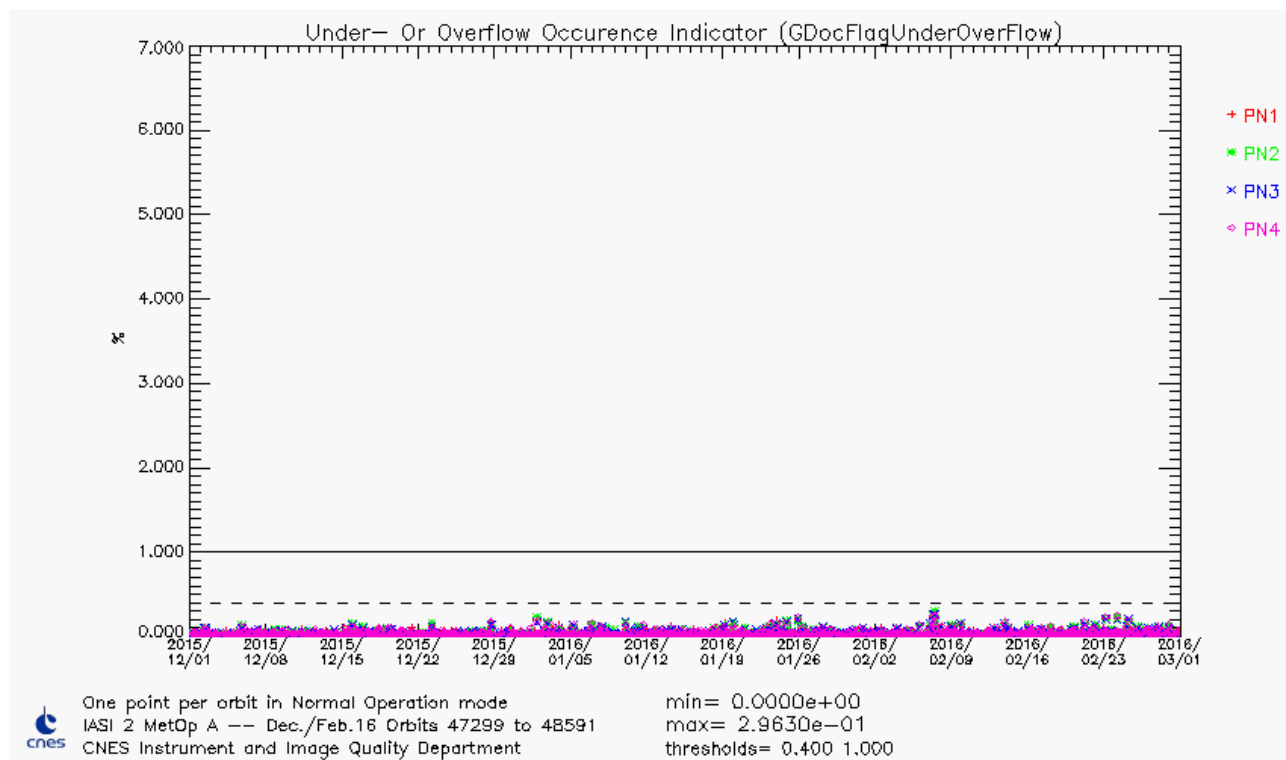


Figure 8 : IASI L0 over/under-flows orbit average of all pixels

Over/underflows ratio is nominal for the reported period.

The geographical distribution of the Overflows and Underflows for the 4 pixels is shown in Figure 9.



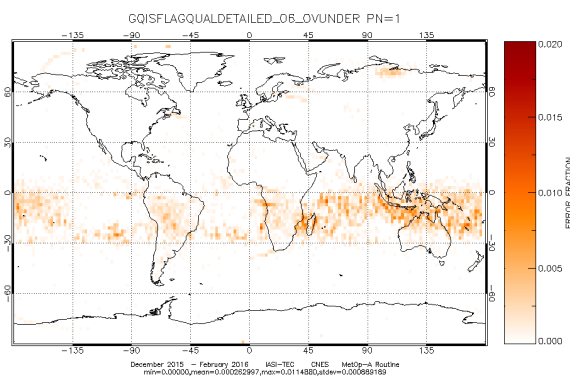
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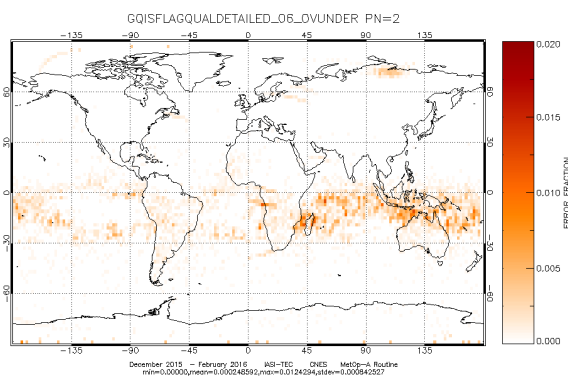
Date: 2017-07-25

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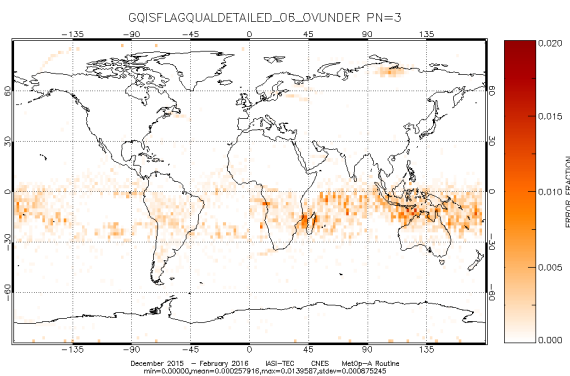
PN1



PN2



PN3



PN4

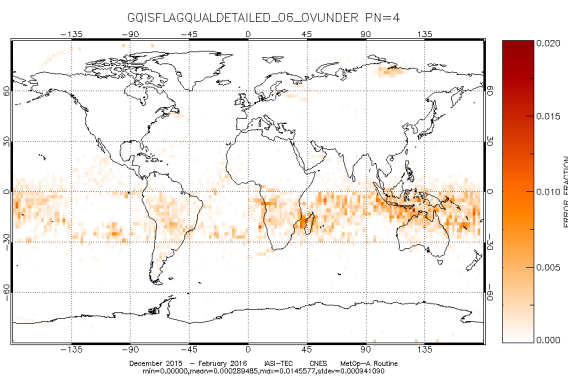




Figure 9 : IASI Overflows and Underflows spatial distribution (per pixel)



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#### 4.3.2.4 Reduced Spectra monitoring

On-board Reduced Spectra is one of the most important parameter to monitor. It ensures that on-board spectra still have a good radiometric calibration when on-board configuration reduced spectra are reloaded. This is the case, for instance, after an instrument mode change.

Reduced spectra are slightly evolving with respect to potential deformation of the interferometer (optical bench).

In order to prevent from a large difference between current and on-board configuration reduced spectra, we apply the DPS processing on the verification interferograms using the reduced spectra from the on-board configuration (TOP) instead of the filtered reduced spectra computed on-board with the current calibration views. These reduced spectra from the on-board configuration are used as initialisation each time there is mode change. If they are too far from the reality, no spectra can be computed on-board after a mode change. We monitor the evolution of ZPD determination quality index for calibration views (BZpdNZpdQualIndexBB and CS) obtained by this DPS processing at TEC, results of this monitoring are given hereafter.

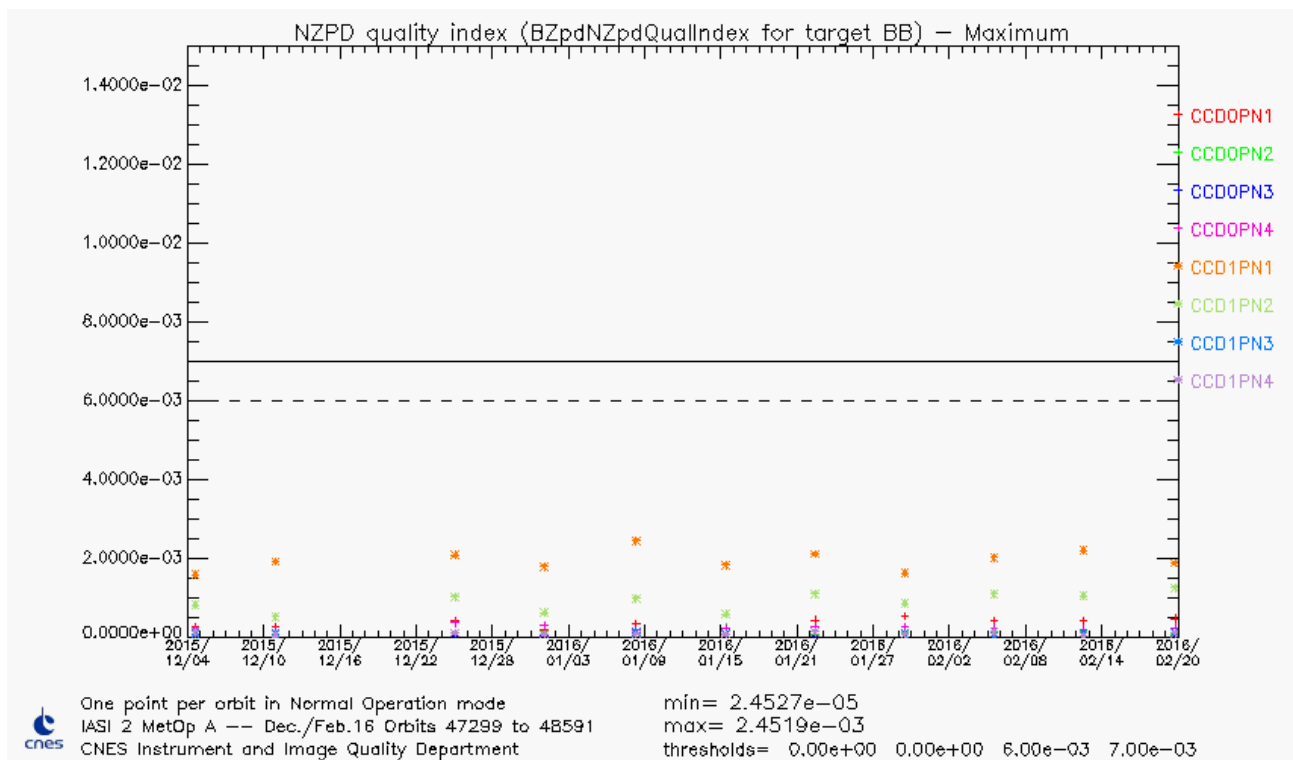


Figure 10 : Max of NZPD quality index for all pixels and CCD - BB

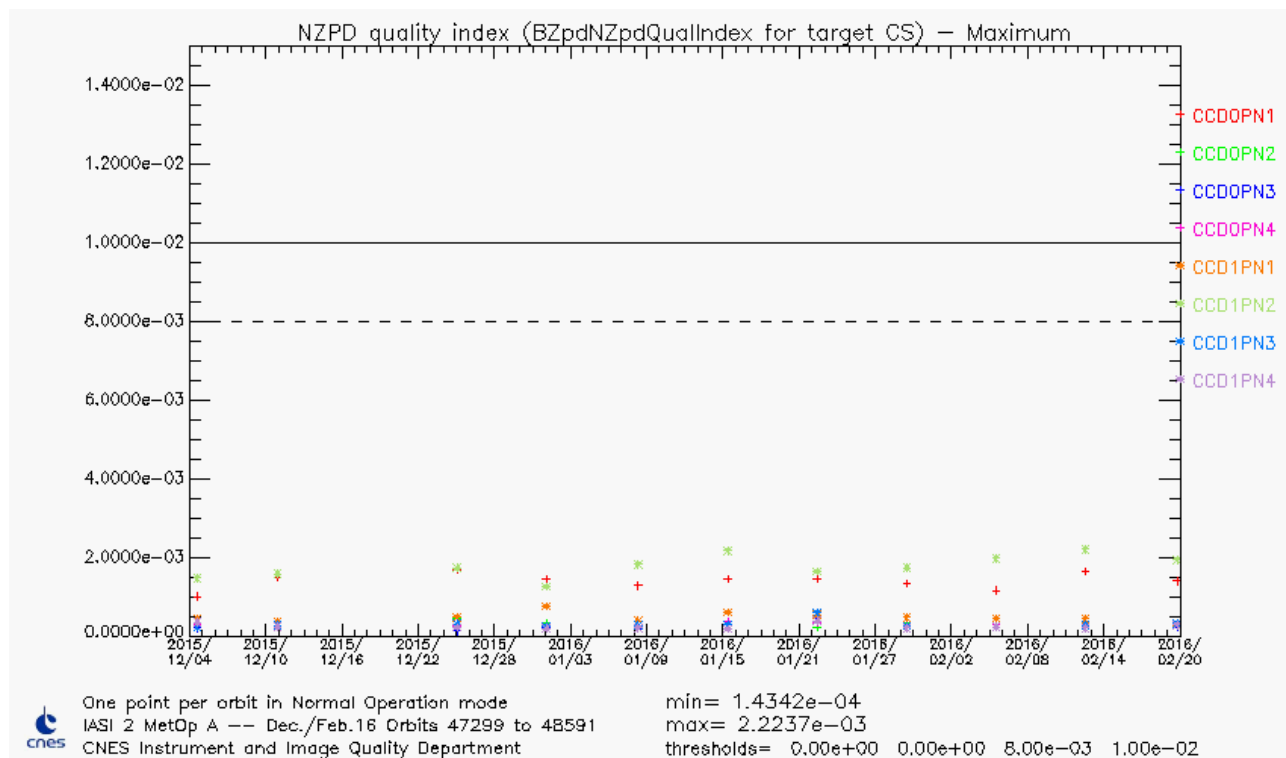


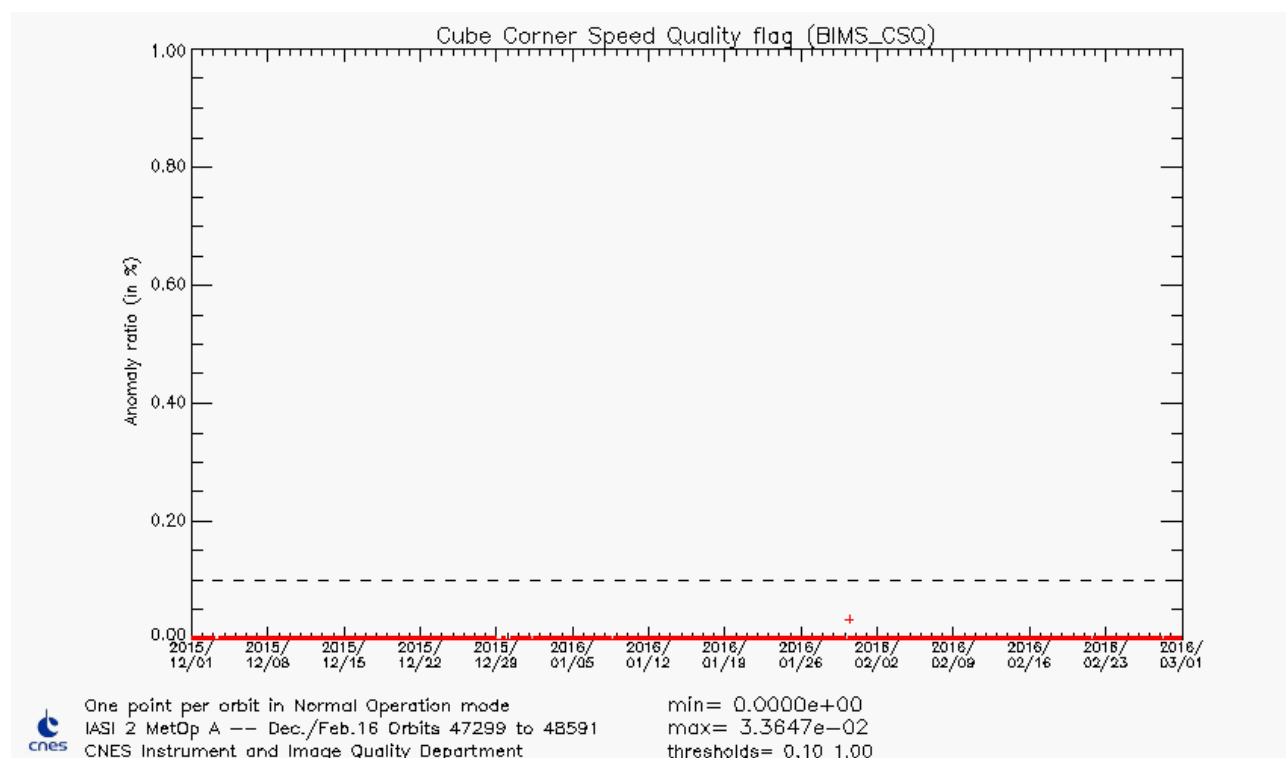
Figure 11 : Max of NZPD quality index for all pixels and CCD - CS



As soon as average BZPDNZPDQualIndexBB and CS remain below 0.02 on-board reduced spectra are robust to an instrument mode change.

The reduced spectra quality is well within specification since the last update of the on-board reduced spectra performed in August 2015.

#### 4.3.2.5 Cube corner Speed Quality (CSQ) monitoring

From verification products (BB & CS):



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From engineering products (EW):

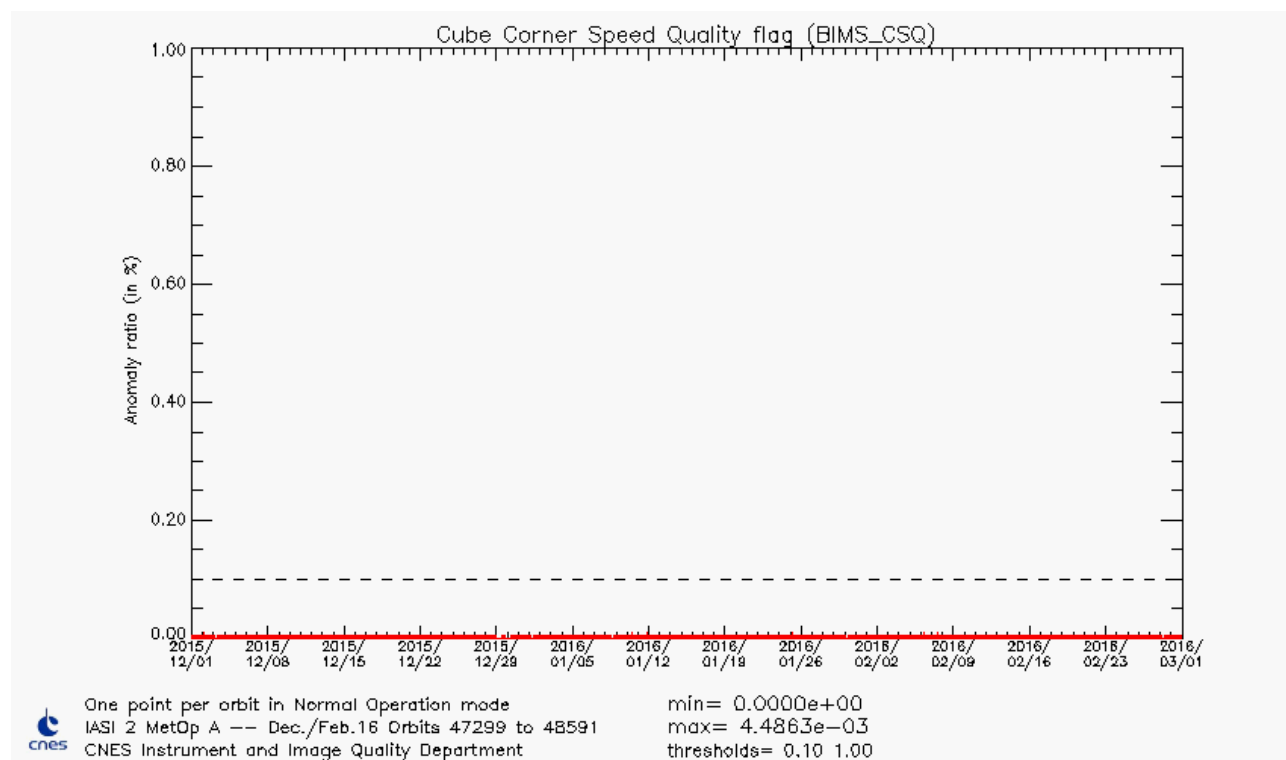




Figure 12 : Number of CSQ

### 4.3.3 Second level flags and quality indicators

L0 Flag and quality indicators are stable

### 4.3.4 Conclusion

L0 Flag and quality indicators are stable.

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## 4.4 LEVEL 1 DATA QUALITY (L1)

### 4.4.1 Overall quality

The IASI overall quality is shown as the orbit averages of the quality indicator for the individual pixels in the next figure.

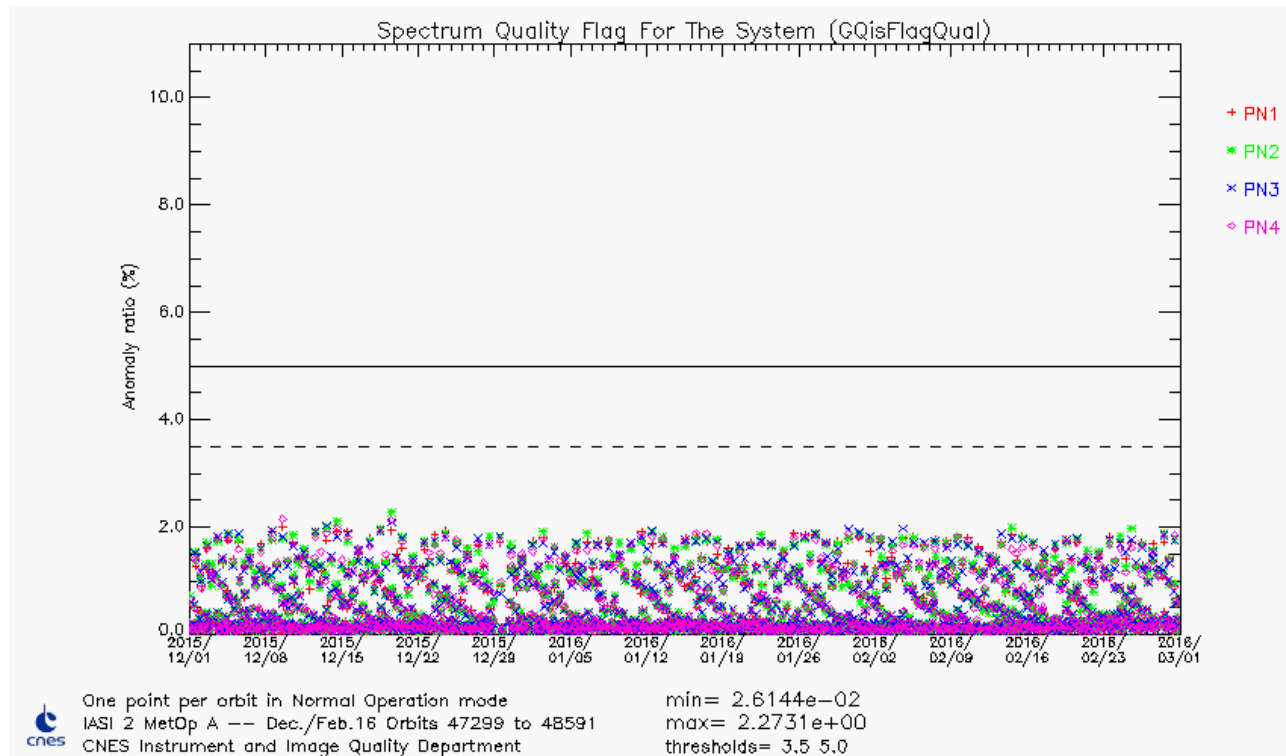




Figure 13 : IASI L1 data quality orbit average (% of bad by PN)

One should note that, over the period covered by the present document, the averaged data rejection ratio is less than 1%. We clearly see that data quality is better on the bands B1 and B2 in comparison to band B3 (which is the most affected by spikes).

The geographical distribution of the IASI product overall quality for the 4 pixels is shown in Figure 14.

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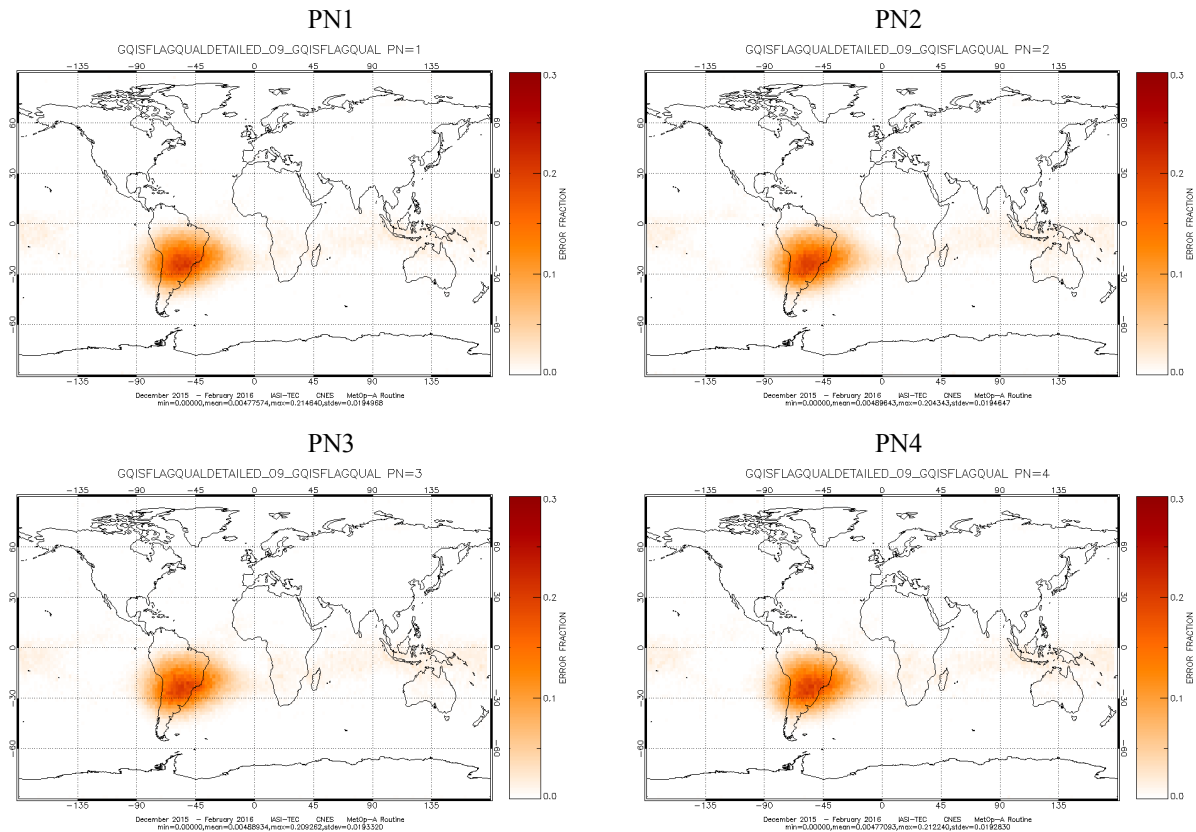


Figure 14 : IASI product overall quality spatial distribution (per pixel)

The main contributors are the spikes (mainly in band 3, which is the band the most sensitive to the spikes).

#### 4.4.2 Main flag and quality indicator parameters

All the quality indexes that follow are general L1 quality indexes of sounder products.

GQisQualIndex – average – is the average general quality index of the sounder products.

GQisQualIndexIIS is the IASI integrated imager (IIS) images quality index.

GQisQualIndexSpect is the spectral quality index of the sounder products.

GQisQualIndexRad is the radiometric quality index of the sounder products.

GqisQualIndexLoc is the ground localisation quality index of the sounder products.

MDptPixQual is a quality index for IASI integrated imager (IIS) that represents a fraction of not dead pixels.

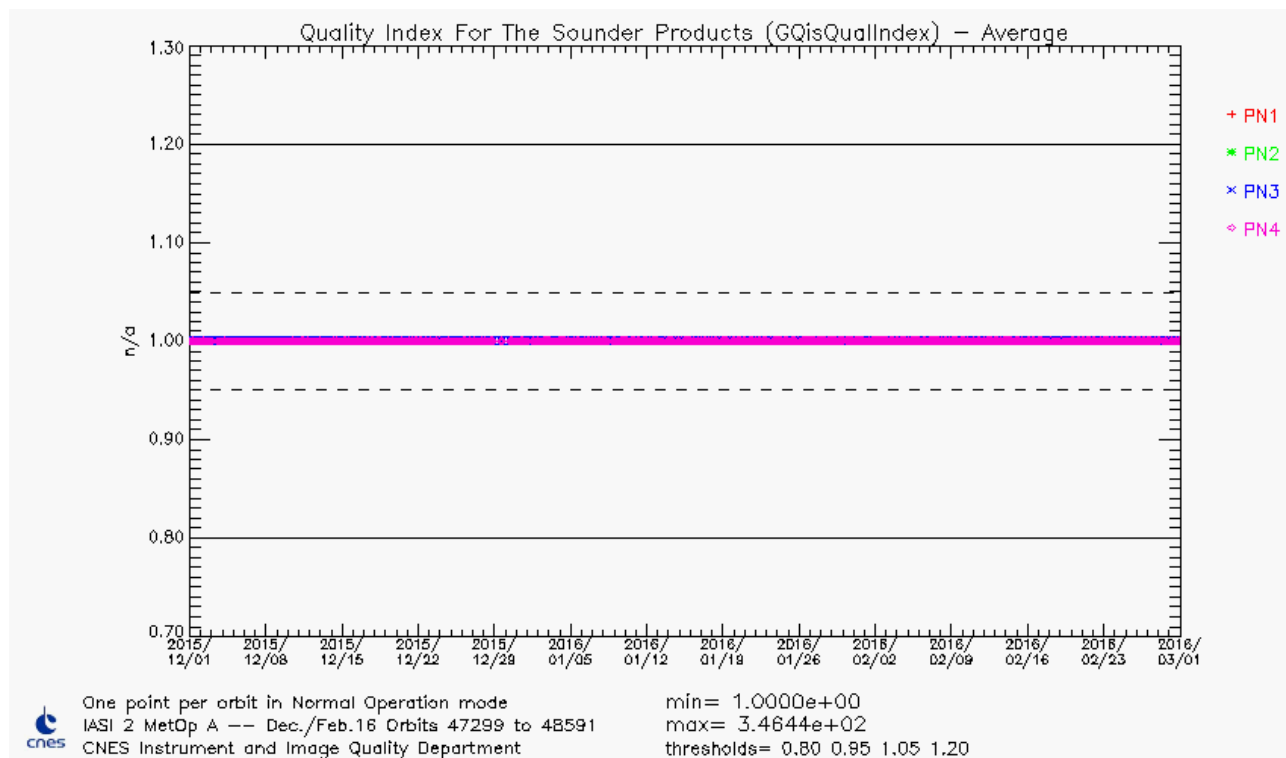


Figure 15 : GQisQualIndex average (L1 data quality index for IASI sounder)

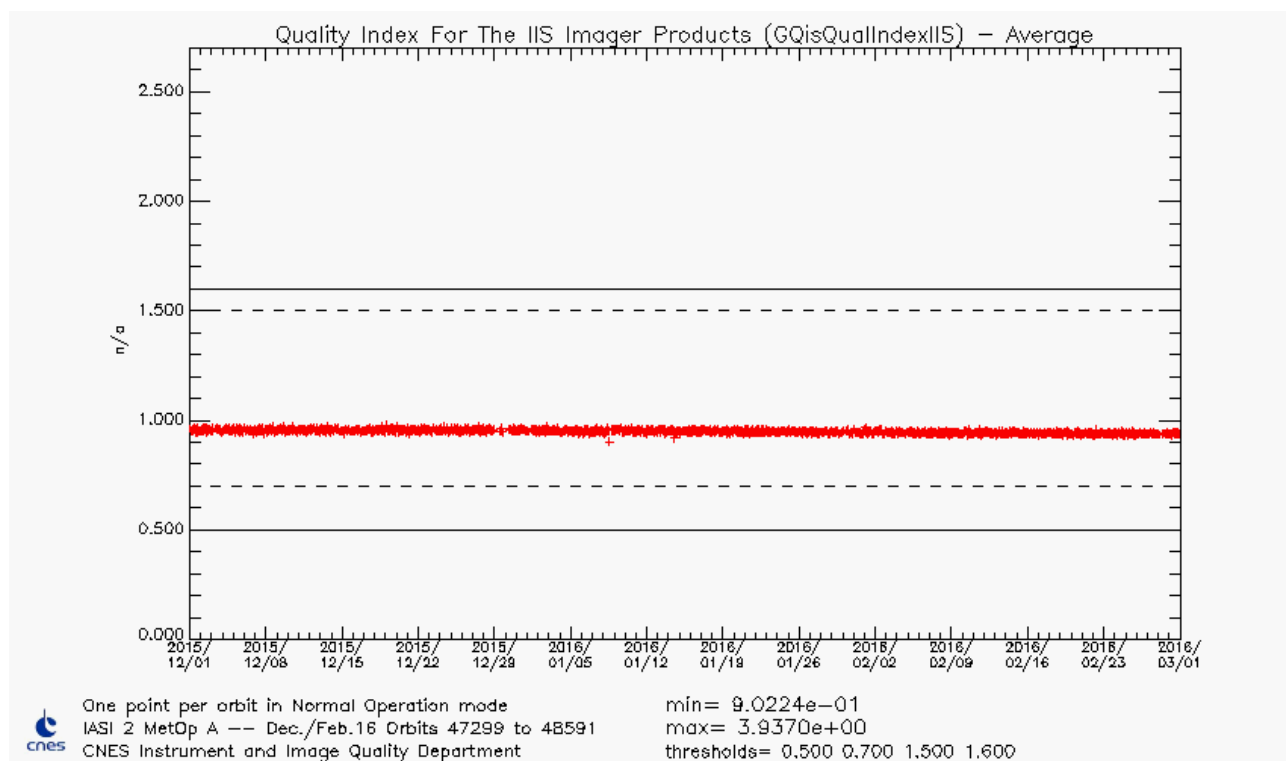


Figure 16 : GQisQualIndexIIS average (L1 data quality index for IASI Integrated Imager)

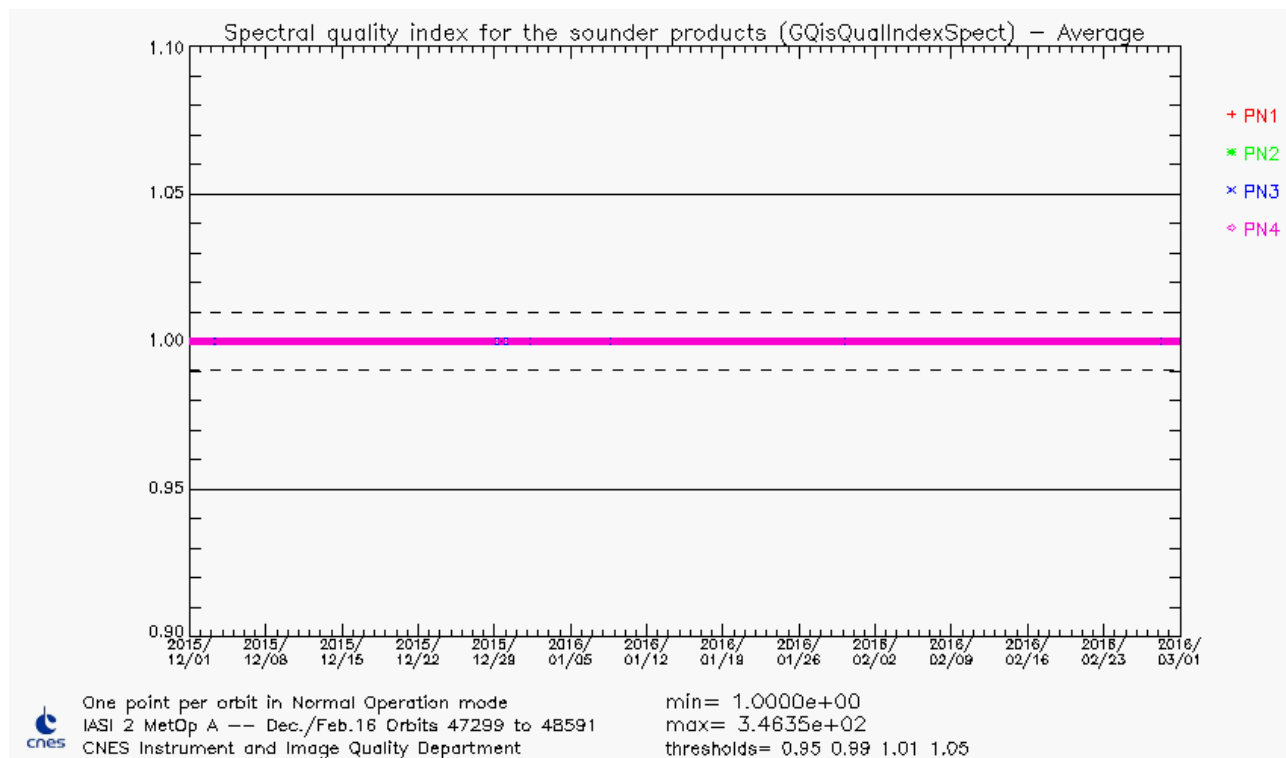


Figure 17 : GQisQualIndexSpect average (L1 data index for spectral calibration quality)

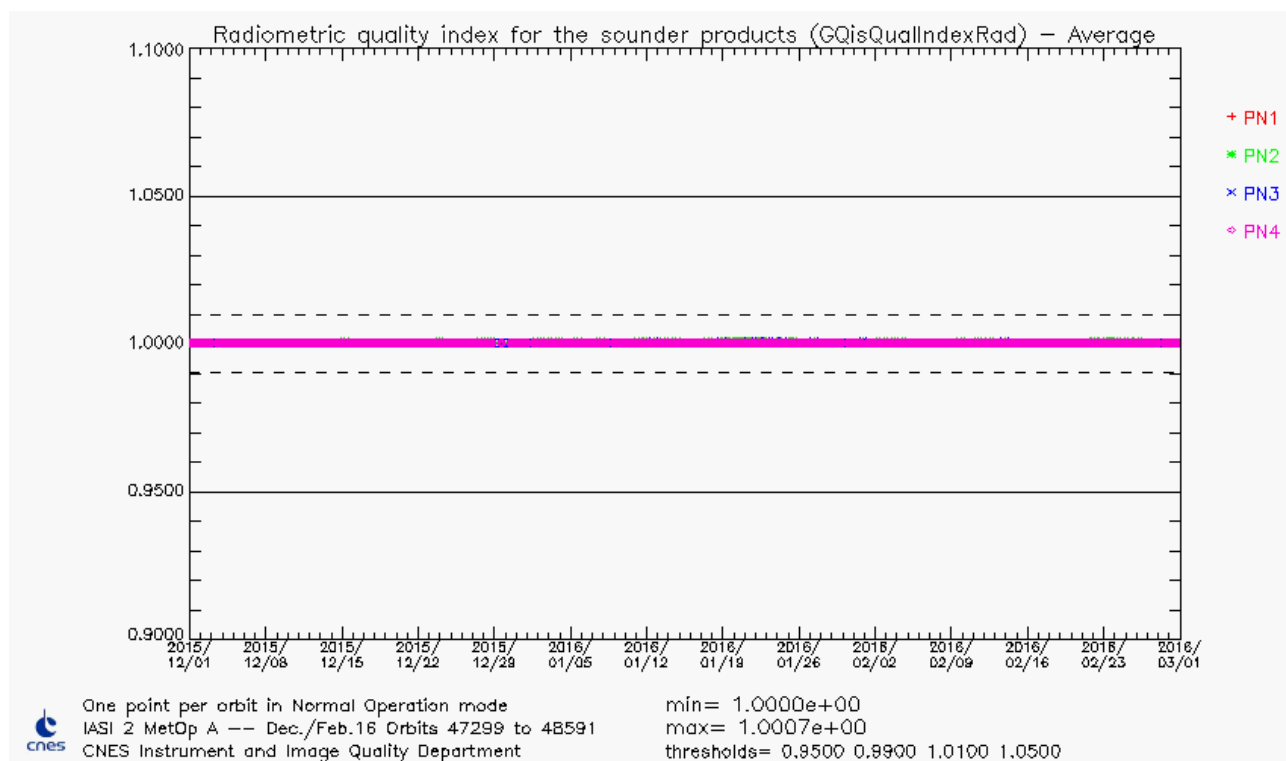


Figure 18 : GQisQualIndexRad average (L1 data index for radiometric calibration quality)

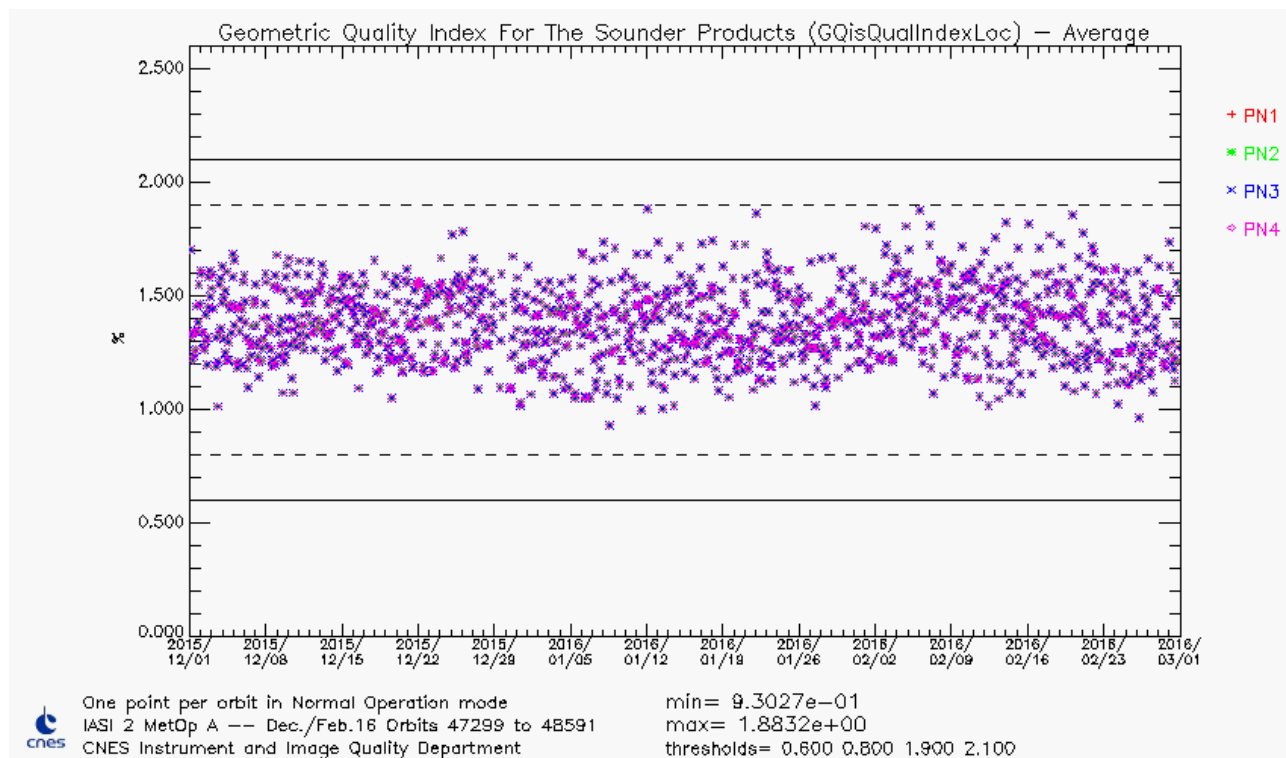


Figure 19 : GQisQualIndexLoc average (L1 data index for ground localisation quality)

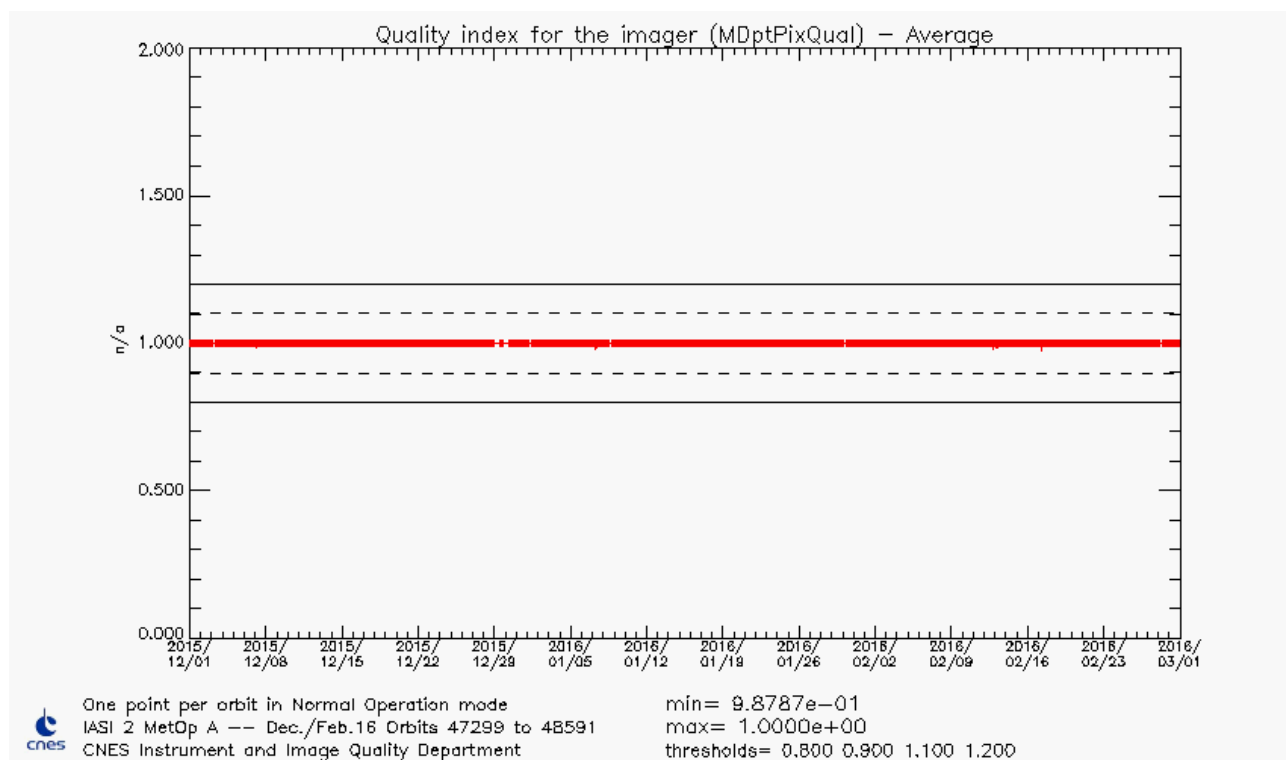


Figure 20 : MDptPixQual average (L1 quality index for IASI integrated imager; fraction of not dead pixels)

#### 4.4.3 Conclusion

L1 Flag and quality indicators are stable and meet the specifications.



## 4.5 SOUNDER RADIOMETRIC PERFORMANCES

### 4.5.1 Radiometric Noise

Monitoring the radiometric noise allows to monitor the long term degradation of the instrument as well as to look for punctual anomaly of IASI or other component of METOP.

#### Monthly L0 noise estimation (CE)

This monthly estimation is performed during routine External Calibration on BB views.

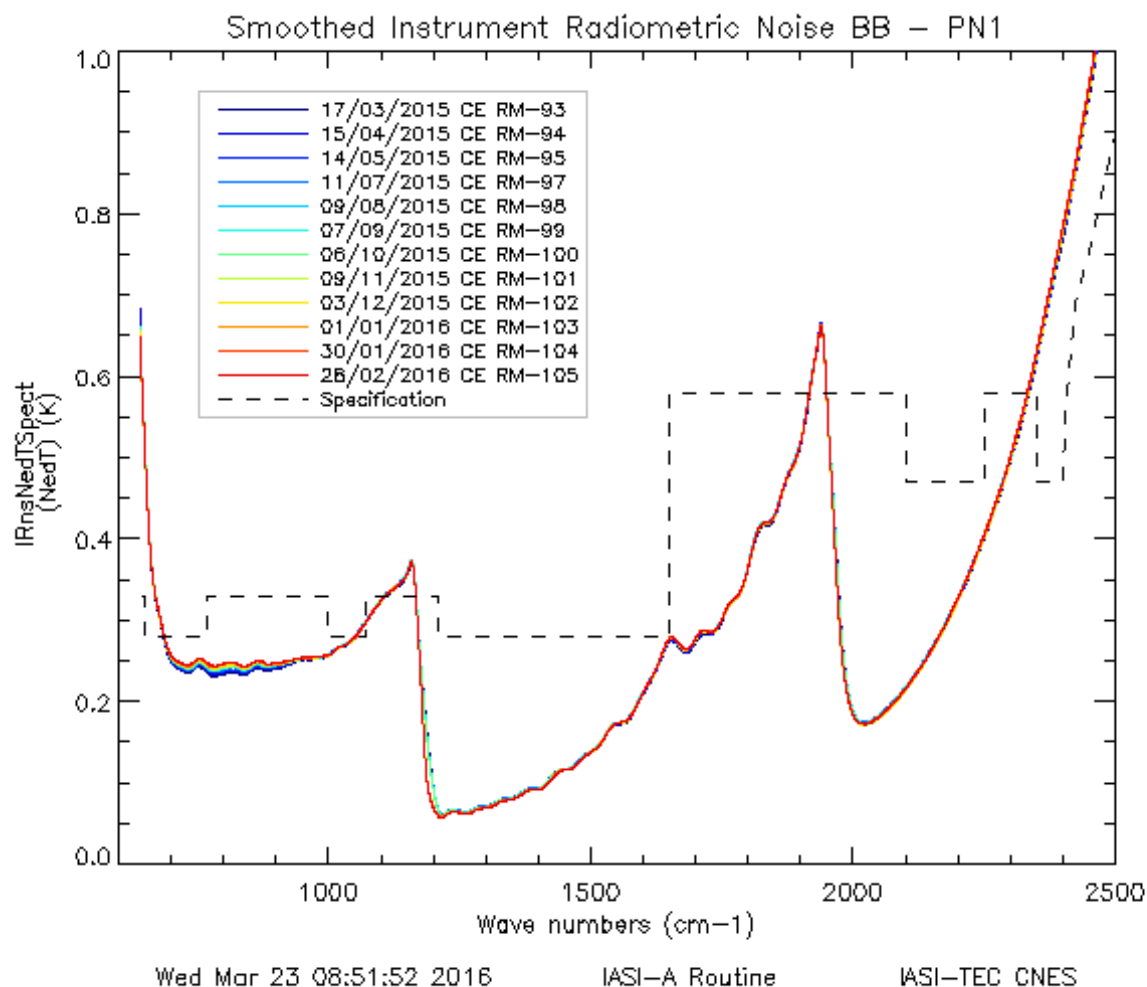


Figure 21 : Instrument noise evolution between start and end of the period

The instrument noise is very stable apart from ice effect between 700 and 1000  $\text{cm}^{-1}$ . This point will be developed in section 4.5.4.1.

## 4.5.2 Radiometric Calibration

The radiometric calibration allows to convert an instrumental measurement into a physical value. The radiometric calibration is used to convert an interferogram into an absolute energy flux by taking into account instrument discrepancies. Even if the calibration has been studied on ground, it has to be continuously monitored in-flight in order to follow any potential degradation of the instrument (optics, detectors ...).

Approach: Radiometric fine characterization has been done during on-ground testing and Cal/Val. All parameters likely to cause a failure in radiometric calibration process have been identified and are continuously monitored. As long as they remain stable, there is no problem with radiometric calibration.

### *Evolution of scanning mirror reflectivity*

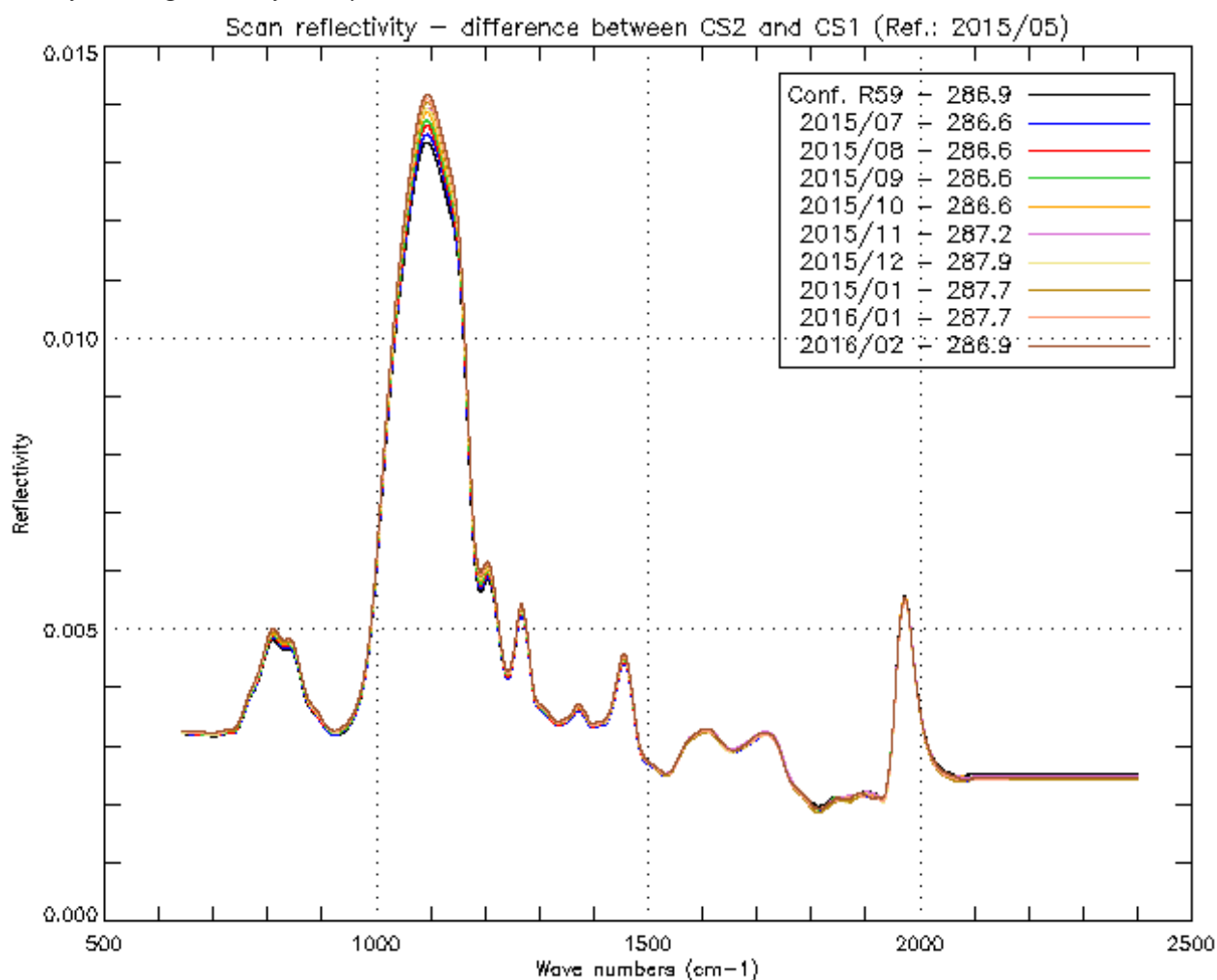
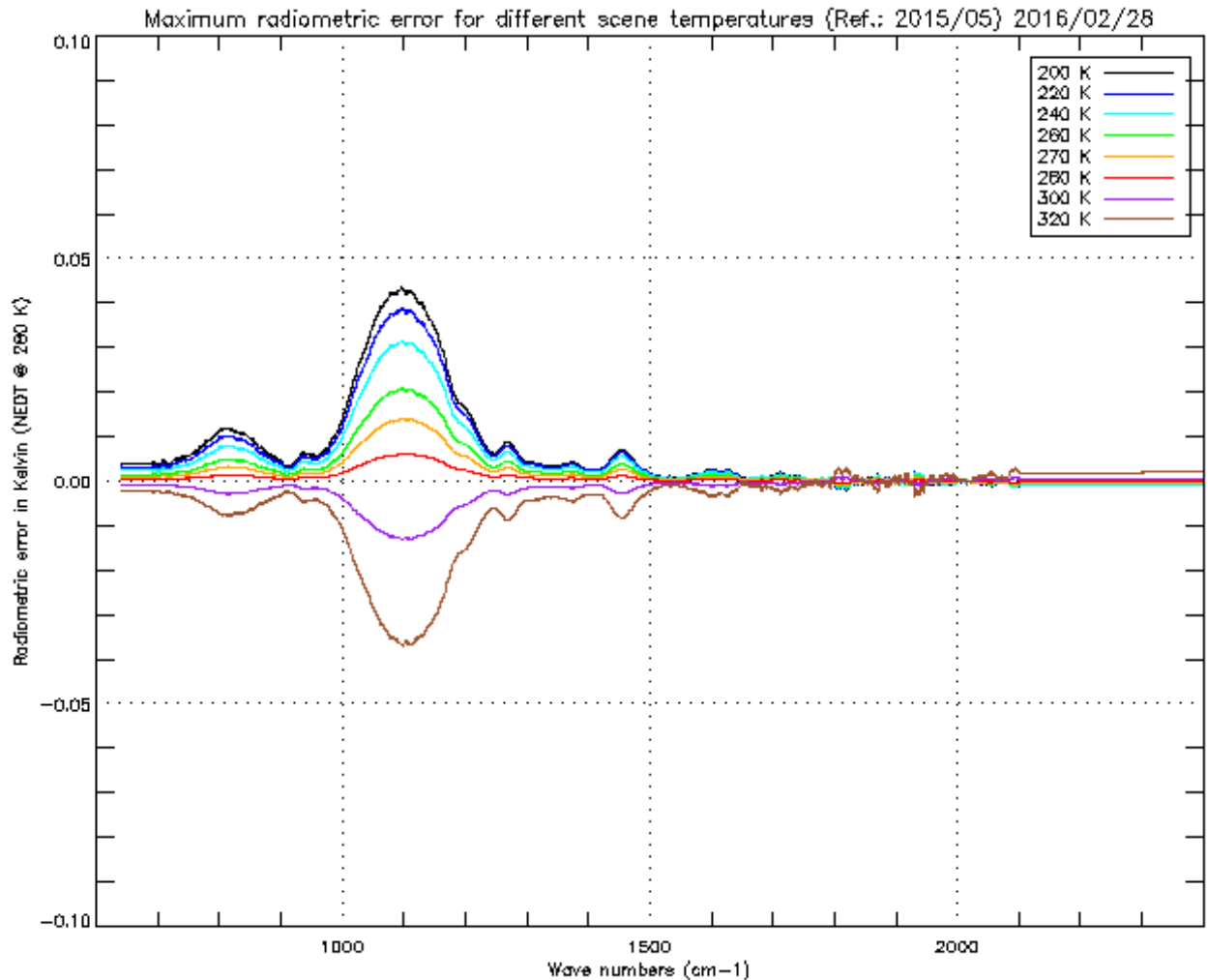


Figure 22 : Scan mirror reflectivity evolution



The reference reflexivity (in black) is the one computed on data from May 14<sup>th</sup> 2015. We see a slight evolution within [1000-1100 cm<sup>-1</sup>] band. Values for wavenumbers greater than 2400 cm<sup>-1</sup> are not significant because of instrument noise.

The next figure shows the translation of scan mirror reflectivity in terms of maximum radiometric calibration error for different scene temperatures.



*Figure 23 : Radiometric calibration error due to scan mirror reflectivity dependency  
with viewing angle Maximum effect on SN1 for different scene temperature.  
Done with the period May 2015/ February 2016*

In any cases radiometric calibration maximum error is lower than the specification (0.1K). The scan mirror reflectivity law (on ground configuration), prepared with May 2015 routine External Calibration data, has been updated in the operational ground segment in June 2015.

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Internal black body

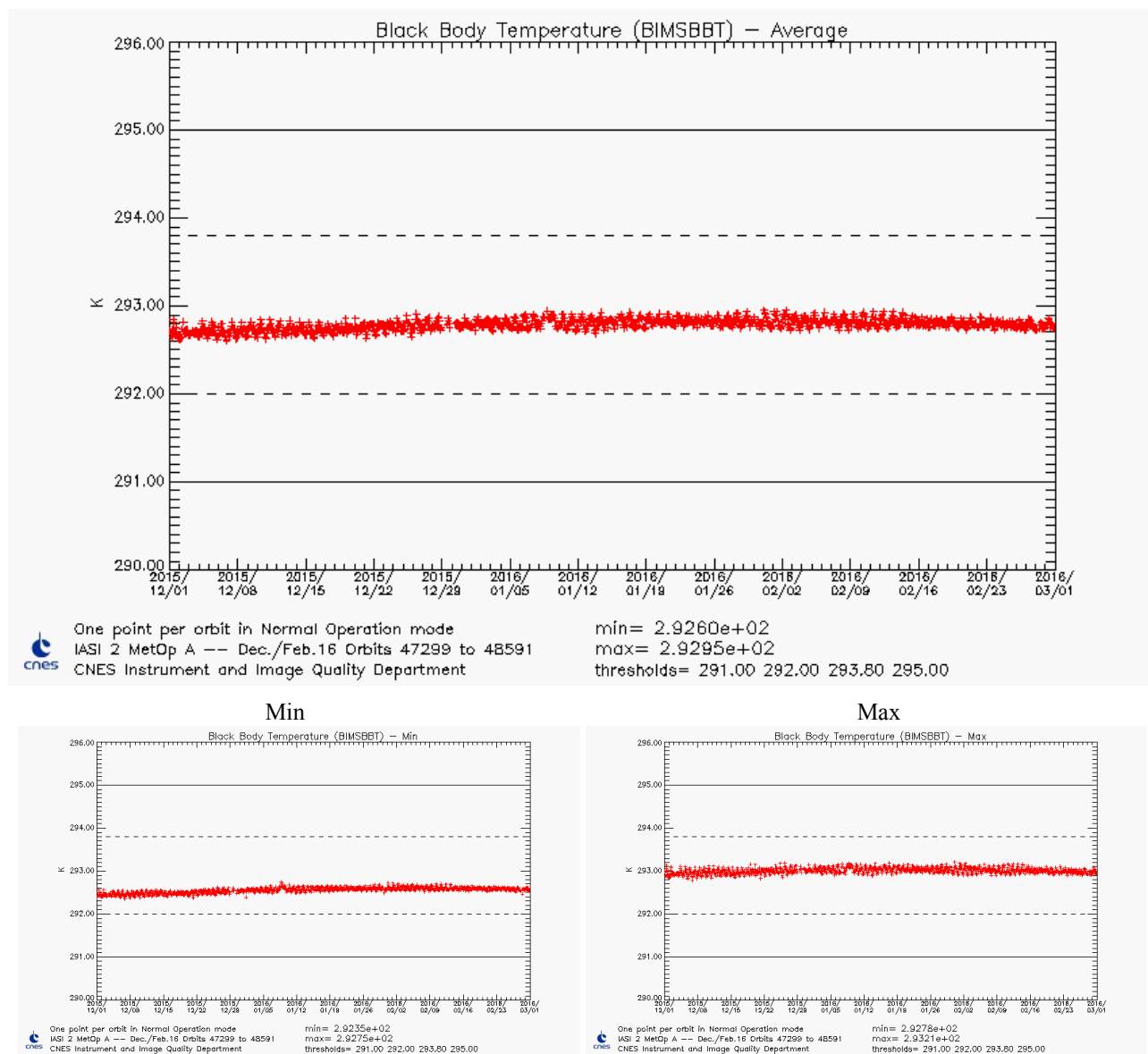




Figure 24 : Black Body Temperature

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### Non linearity of the detection chains

Non-linearity tables of the detection chains are still nominal as long as sounder focal plane temperature variation amplitude is lower than 1K.

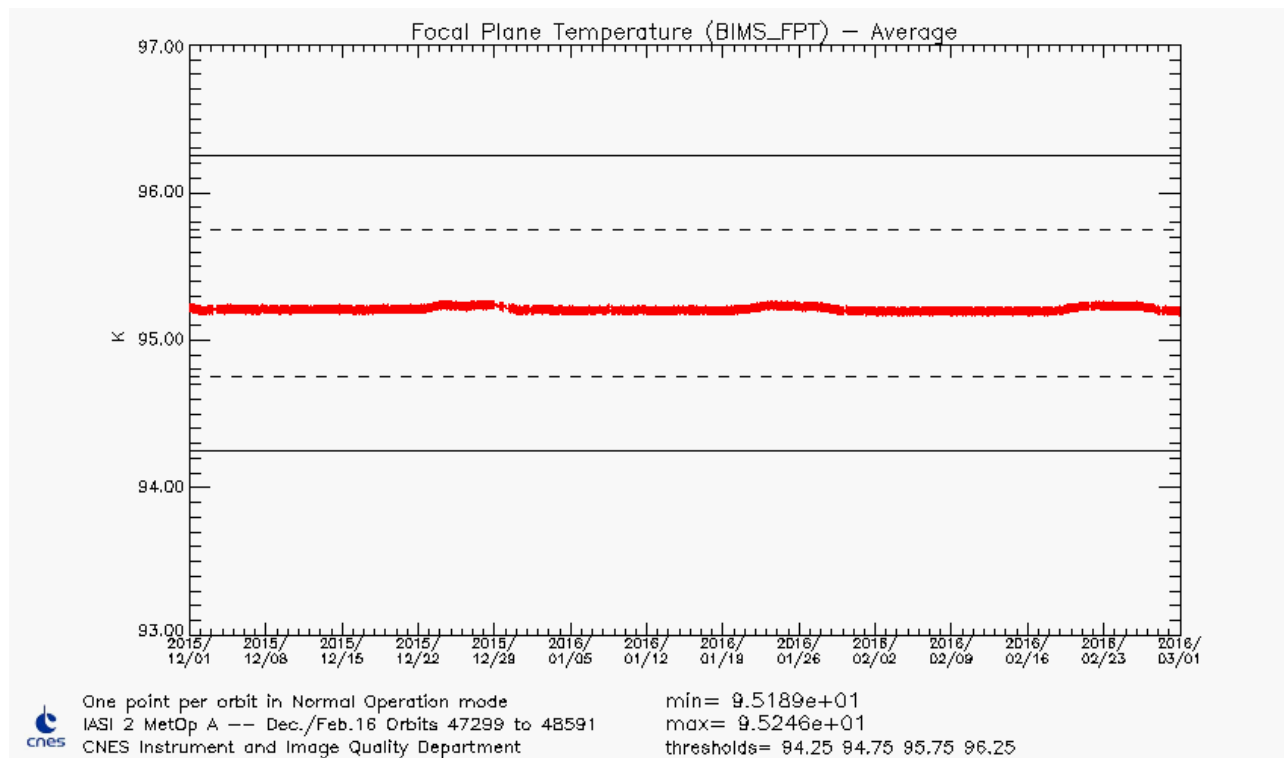




Figure 25 : Focal Plane Temperature

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### 4.5.3 Delay of detection chains

Long term stability and values lower than 400 ns are required in order to properly take into account cube corner velocity fluctuations.

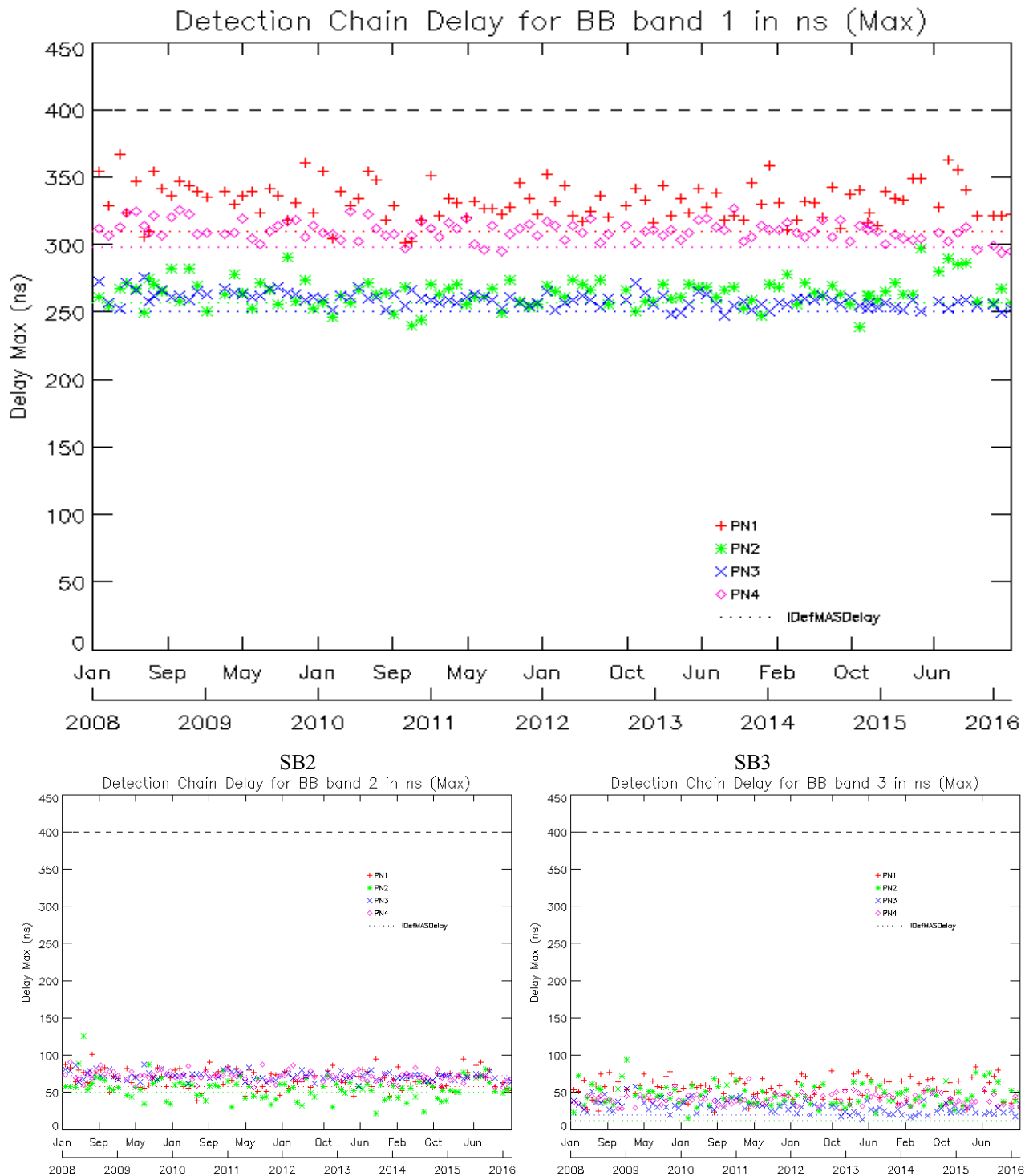




Figure 26 : Monitoring of detection chain maximum delays for all bands

Since the switch to IMS redundant side, the detection chain delays have changed a little bit, especially for pixel 2. Values are still as expected.

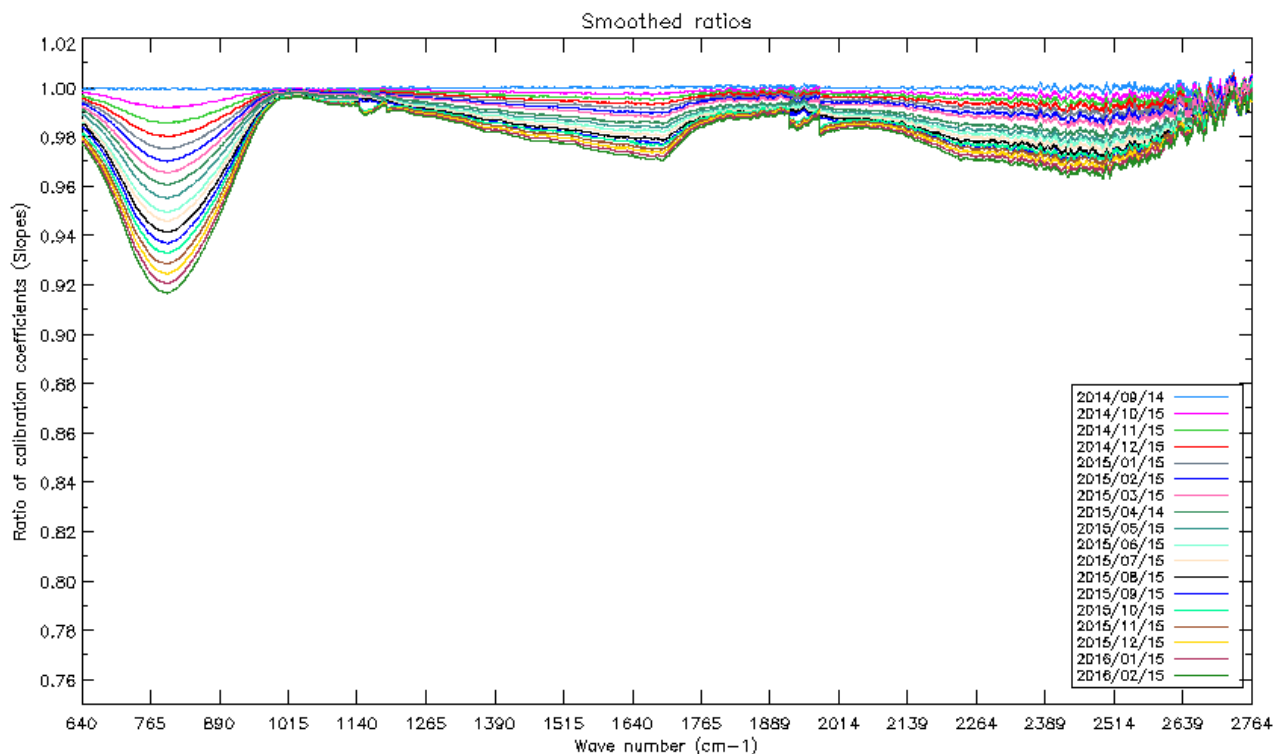
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#### 4.5.4 Optical Transmission



##### 4.5.4.1 Ice

The IASI interferometer and optical bench are regulated at 20°C temperature, while the cold box containing cold optics and detection subsystem is at about -180°C. Water desorption from the instrument causes ice formation on the field lens at the entrance of IASI cold box. This desorption phenomenon is particularly important at the beginning of the instrument in-orbit life. That's why one of the very first activities of IASI in-orbit commissioning was an outgassing phase consisting in heating the cold box up to 300 K during 20 days. This operation allows removing most of the initial contaminants coming from IASI and other MetOp instruments. A routine outgassing is then needed from time to time to remove ice contamination, but less and less frequently as the desorption process becomes slower. A first run of this routine outgassing procedure (shorter duration and at 200 K), was done for validation purpose during commissioning phase in December 2006. The second one, which was actually the first in routine phase, was done in March 2008. The third one was done in August 2010 and the last was done on September 2014.

The maximum acceptable degradation of transmission is about 20% loss at 850 cm<sup>-1</sup> (which corresponds to an ice deposit thickness of about 0.5 µm).



*Figure 27 : Ratio of calibration coefficient slopes as a function of wave number and time after the last decontamination*

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#### 4.5.4.2 Prediction of decontamination date

The transmission degradation rate is regularly monitored by CNES TEC through gain measurements given by calibration coefficients ratios.

The loss of instrument gain due to ice contamination is, as expected, decreasing over time. The decontamination that occurred from 8<sup>th</sup> to 13<sup>rd</sup> September could be the last one since the next one is not predicted before end 2019.

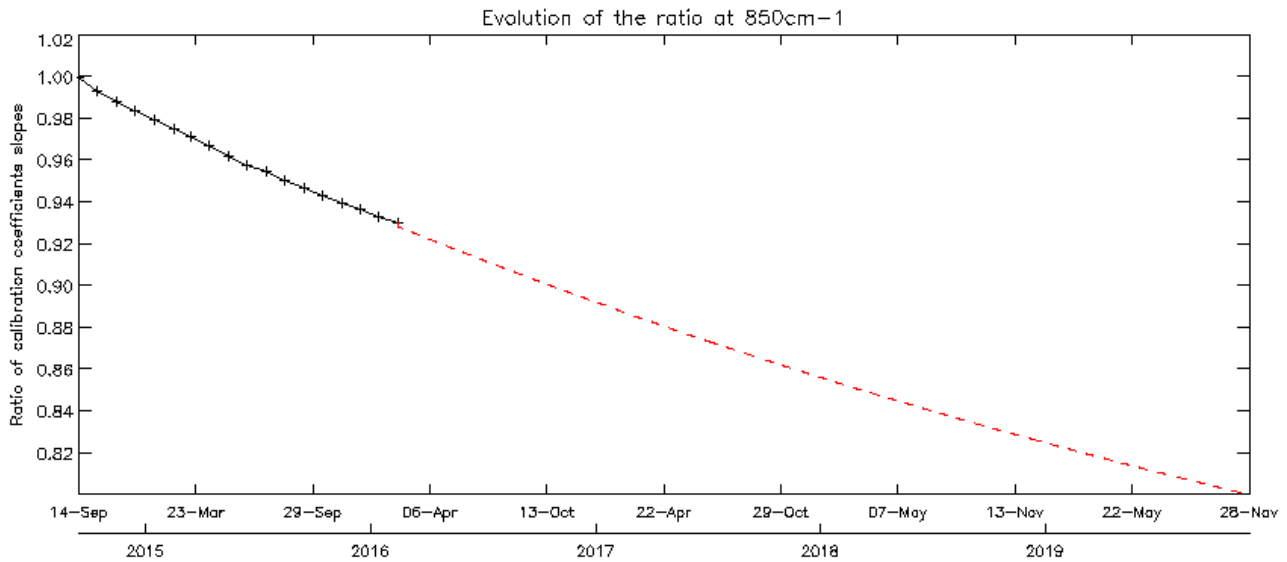


Figure 28 : Temporal evolution of calibration ratio coefficient slopes since the last decontamination. The curve was fitted with a decreasing exponential function to determine a rough date for the next decontamination (relative gain evolution of 0.8)



#### 4.5.5 Interferometric Contrast

The interferometric contrast is defined as the interferogram fringe discrimination power. Figure 29 shows temporal evolution of instrument contrast on the quarter for all pixels and all CCD.

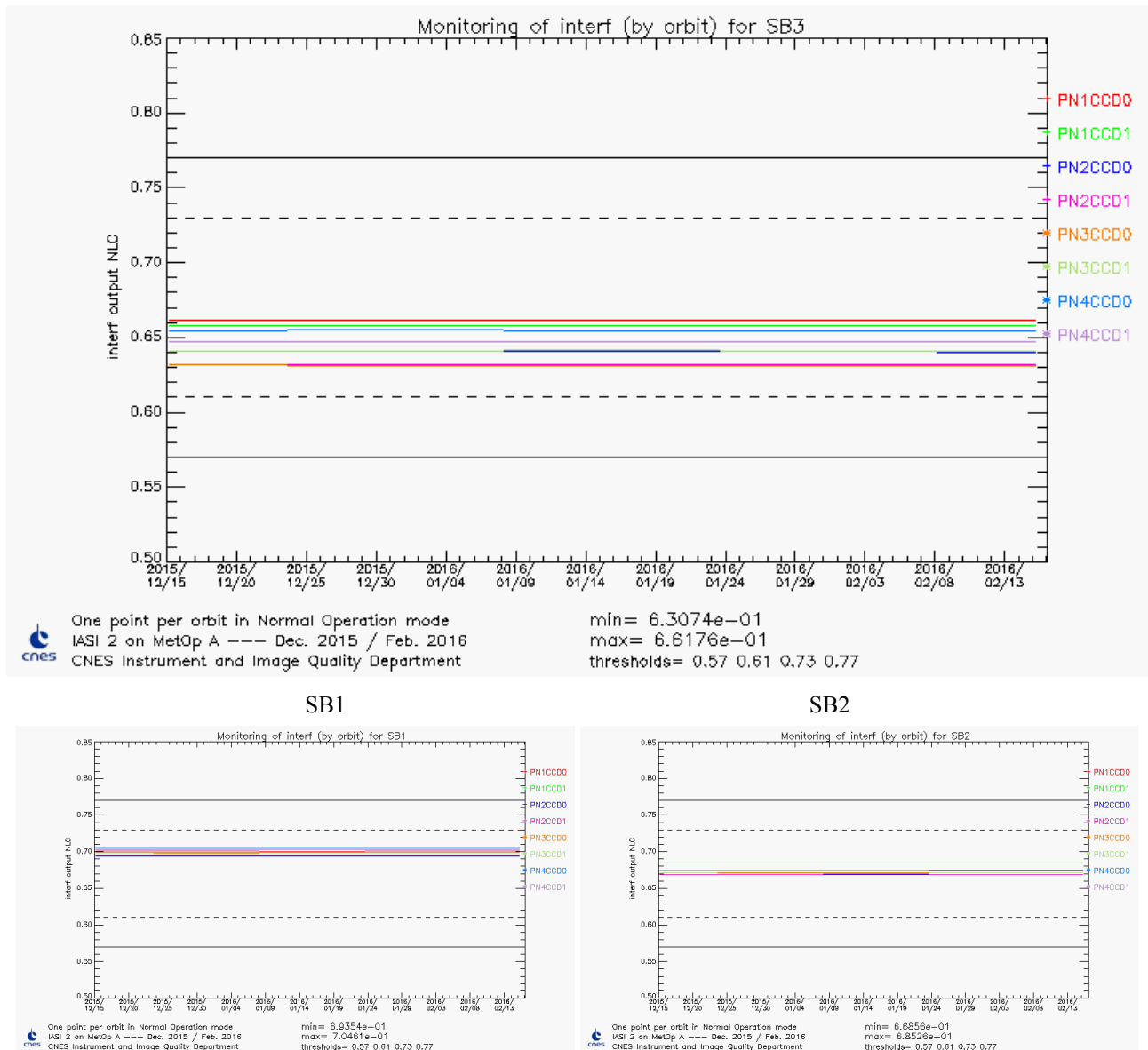


Figure 29 : Contrast Monitoring

#### 4.5.6 Interferogram baseline

The interferogram baseline is the mean value of the interferogram. Figure 30 shows temporal evolution of the baseline of the raw interferograms on calibration targets (BB and CS). The values are raw values, they are not physical, but the evolution is interesting: as the values are proportional to the energy received from a target and calibration targets are stable, the evolution can show the decrease of instrument transmission or events due to energetic particles.

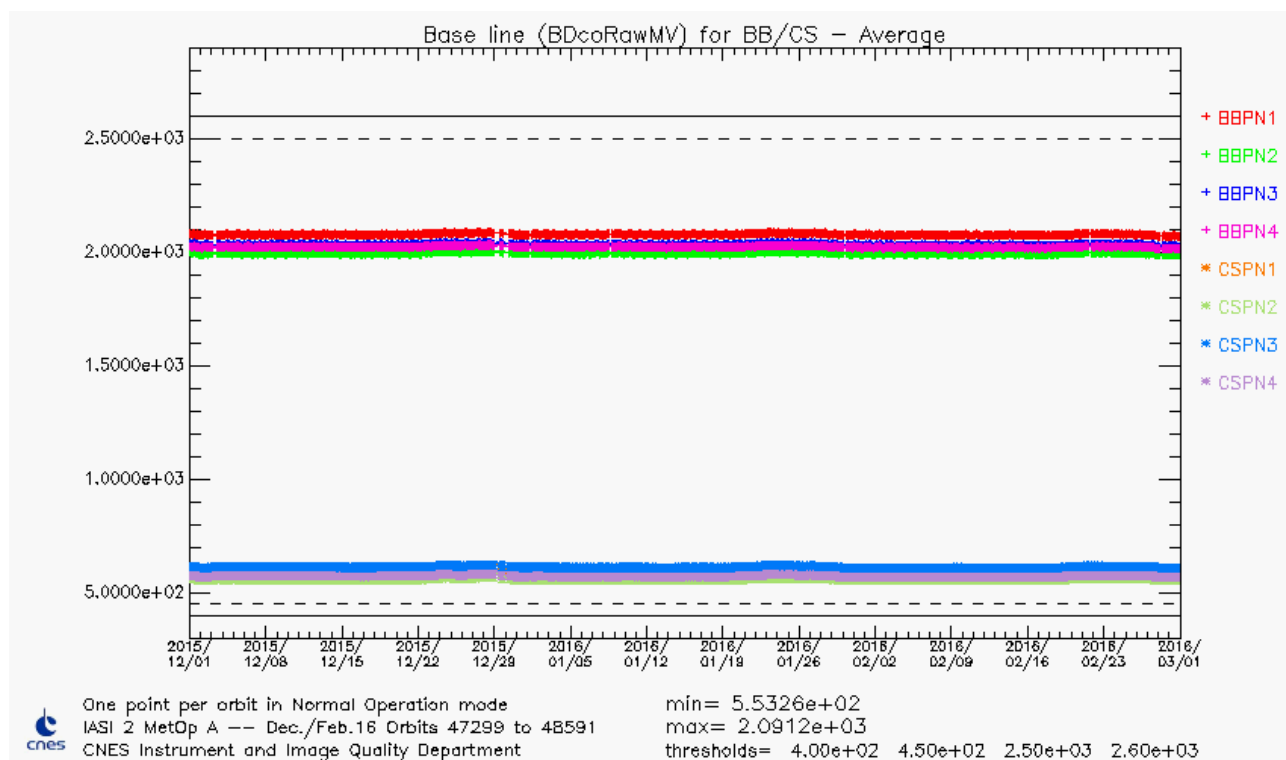


Figure 30 : Monitoring of interferogram baseline

#### 4.5.7 Detection Chain

Detection chains are tuned in gain and offset via telecommand. The goal is to avoid saturation while conserving the maximum dynamic to limit digitalization noise.

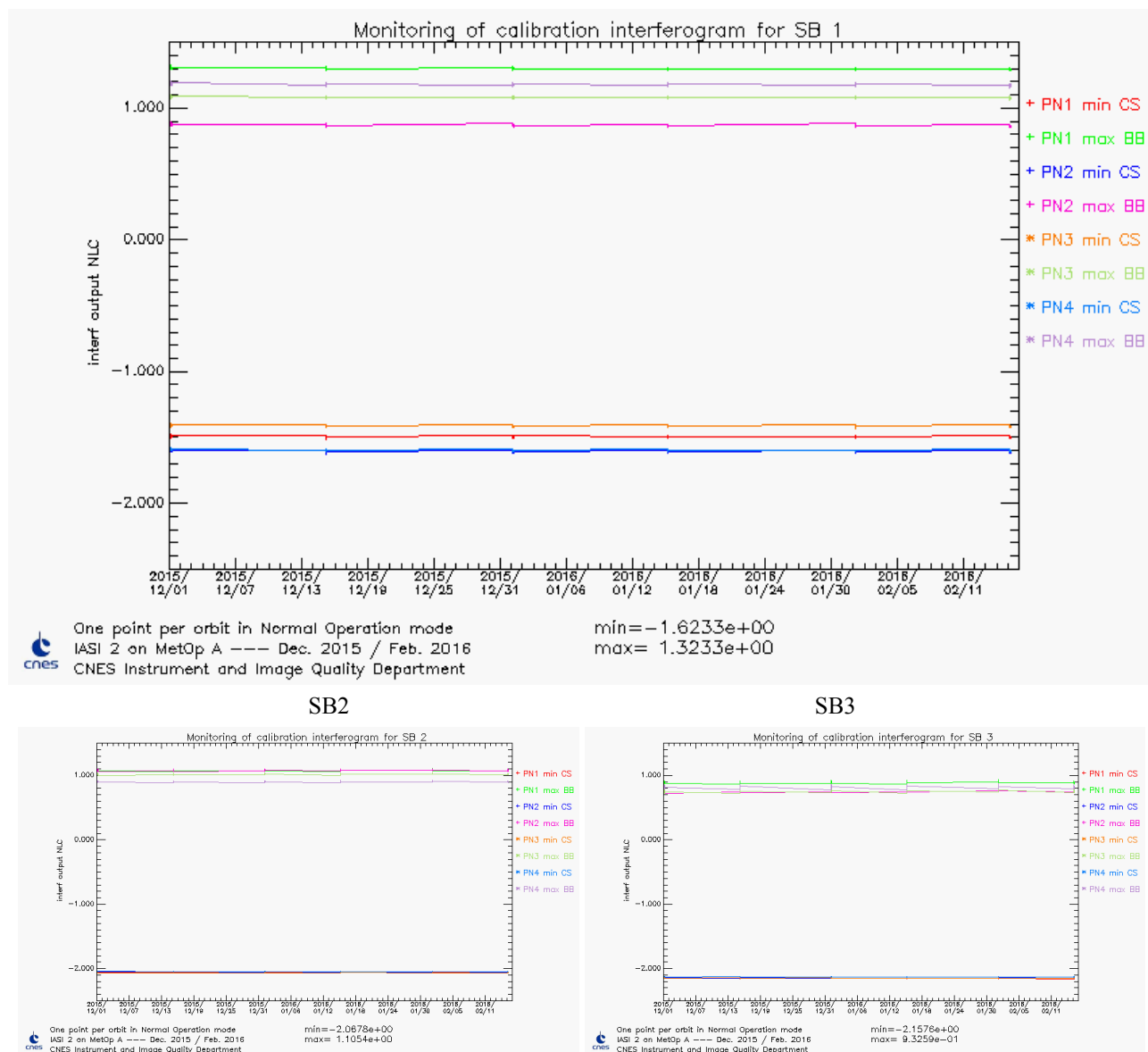




Figure 31 : Monitoring of detection chain margins

Margins are sufficient for the moment.

#### 4.5.8 Conclusion

The radiometric performances of IASI are nominal and stable. The last decontamination was performed in September 2014. Scan mirror reflectivity was updated in June 2015 with May 2015 data. The next update is foreseen end 2016.

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## 4.6 SOUNDER SPECTRAL PERFORMANCES

The goal of the spectral calibration is to provide the best estimates of spectral position of the 8461 IASI channels.

The large sensitivity of infrared spectrum to spectral calibration errors has led to stringent specifications:

- A prior knowledge of spectral position better than of  $2 \cdot 10^{-4}$  (design)
- A posterior maximum spectral calibration relative error of  $2 \cdot 10^{-6}$  (after calibration by OPS)

In order to reach the specification of  $2 \cdot 10^{-6}$ , we need an accurate Instrument Spectral Response Function (ISRF) model. This model have been done and validated in the early time of IASI development.

For sake of operational time constrain, complete ISRF calculation is not done in real-time by OPS software but pre-calculated and stored in a database called “spectral database”. OPS processing determine on-line the most relevant instrument function to be used by OPS with respect to current values of a set of parameters (interferometric axis, cube corner offset...).

The approach to monitor IASI spectral performances is very similar to the one used for radiometric calibration. Spectral calibration fine characterization has been done during ground testing and Cal/Val. All parameters likely to cause a failure in spectral calibration process have been identified and are continuously monitored. As long as they remain stable, there is no problem with IASI spectral calibration.

In addition, a spectral calibration assessment is done over homogeneous scenes when IASI is in external calibration, nadir view.

Since the permanent compensation device stop on 7<sup>th</sup> October 2015, the monitoring of the ghost was cancelled as the data are no more affected by the ghost.

### 4.6.1 Monitoring of the ISRF inputs

#### 4.6.1.1 Position of the interferometric axis

The interferometric axis is the cube corner displacement direction. Its value was around ( $Y = -160 \mu\text{rad}$ ;  $Z = -450 \mu\text{rad}$ ) since the beginning of the life of IASI-A. Since the change of IPSF positions in 2013/05/16, its value has changed and is now stable around ( $Y = 445 \mu\text{rad}$ ;  $Z = 195 \mu\text{rad}$ ). The central position used in the “spectral database” generation, are  $400 \mu\text{rad}$  and  $200 \mu\text{rad}$ , respectively for Y and Z axis.

Since the drift of the interferometer axis is lower than  $300 \mu\text{rad}$ , there is no need to update the “spectral database”.



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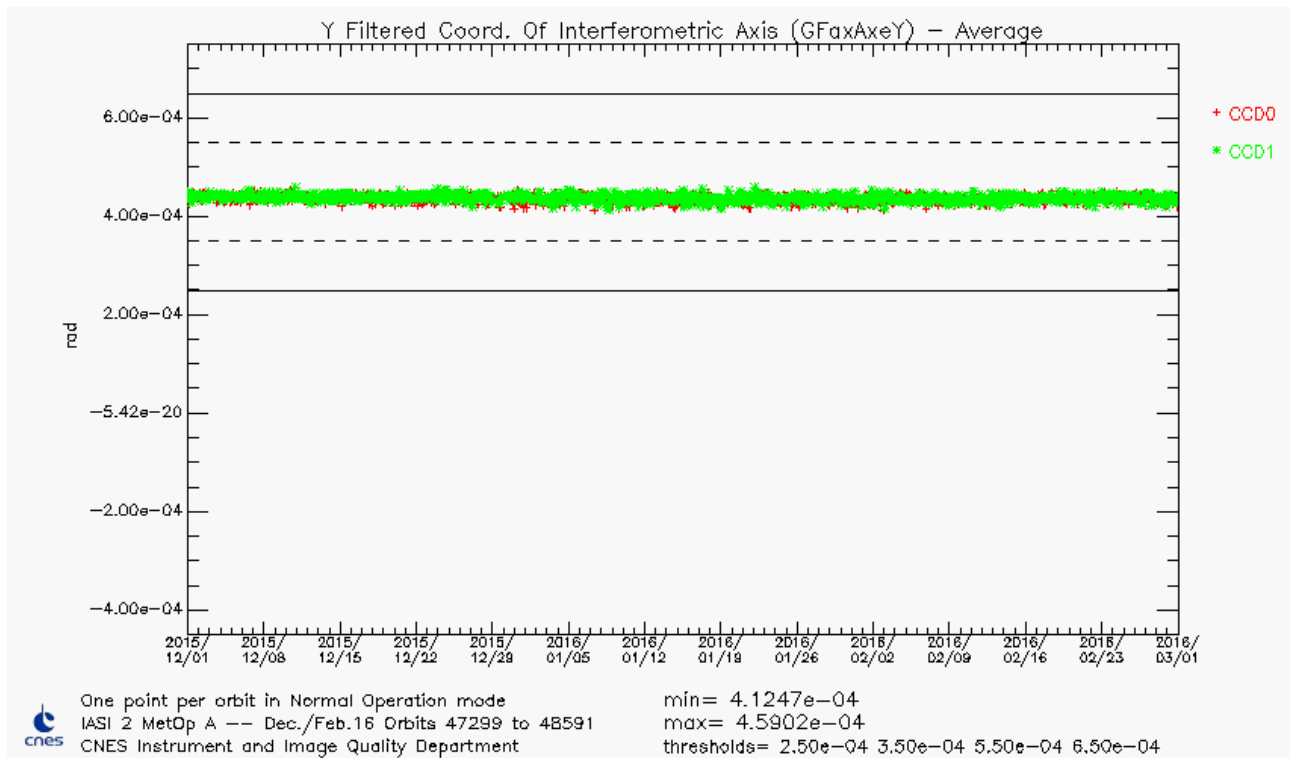
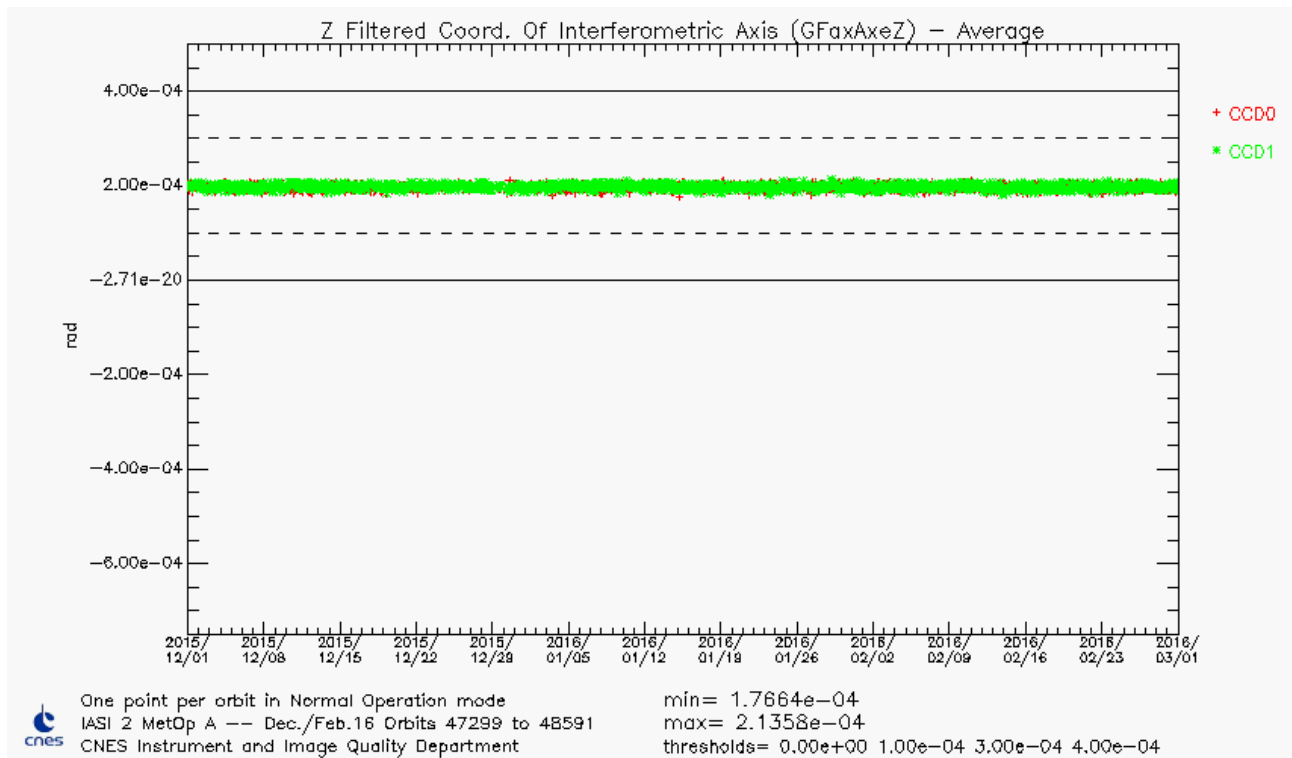
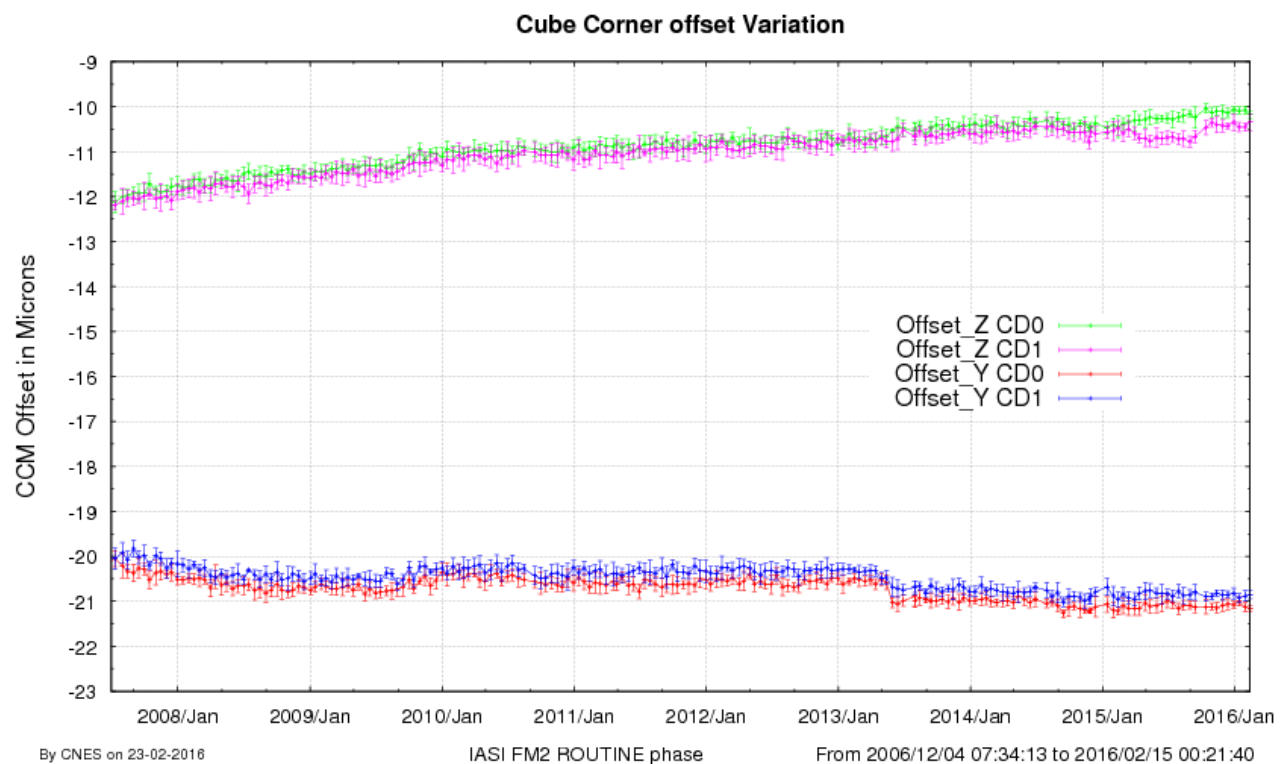




Figure 32 : GFaxAxeY average (Y filtered coordinates of sounder interferometric axis)



#### 4.6.1.2 Cube Corner constant offset



Reference cube corner offsets, used in the spectral database of the period (ODB14), are -21.08  $\mu\text{m}$ , -20.69  $\mu\text{m}$ , -10.49  $\mu\text{m}$  and -10.77  $\mu\text{m}$ , respectively for Y CD0, Y CD1, Z CD0 and Z CD1.

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The slight change of the cube corner constant offset that happened in May 2013 is due to the update of the IPSF positions in the ground configuration.

Since the drift of cube corner offset is lower than 4  $\mu\text{m}$ , there is no need to update the “spectral database”.

#### 4.6.1.3 Cube corner velocity

Refer to REVEX, paragraph 5.5.

#### 4.6.1.4 Interferometer optical bench temperature

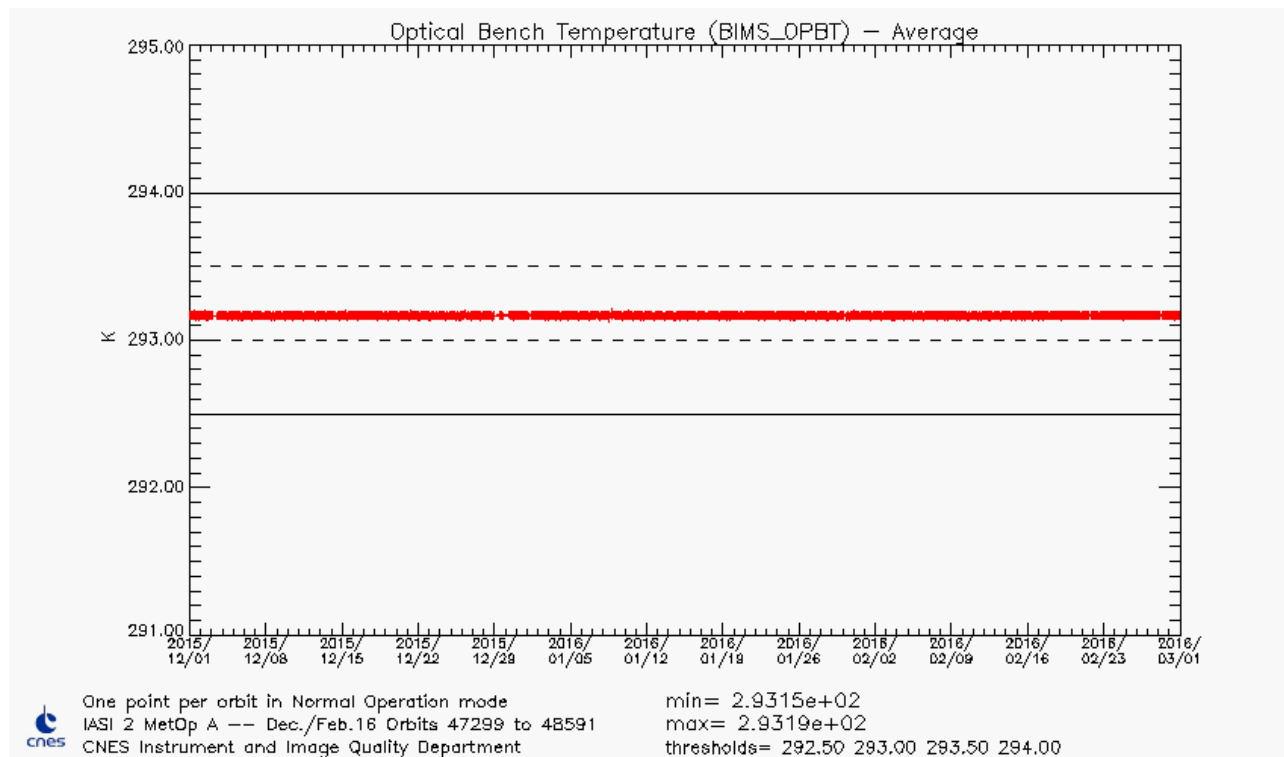


Figure 35 : Optical bench Temperature

#### 4.6.2 Spectral calibration assessment

Absolute spectral calibration assessment and Interpixel spectral calibration assessment are performed during routine External Calibration on Earth views at nadir (SP 15) and synthesized once a year.

Refer to REVEX, paragraph 6.6.2.

#### 4.6.3 Conclusion

All parameters impacting IASI spectral calibration are stable and within specifications.

## 4.7 GEOMETRIC PERFORMANCES

The geometric calibration is performed on ground (level 1 processing). Most of the analyses of geometric performances require being in external calibration mode.

Specifications are the following: the IIS/AVHRR co-registration has to be better than 0.3AVHRR pixel while the IIS/sounder co-registration has to be better than 0.8mrad.

### 4.7.1 Sounder / IIS co-registration monitoring

This monitoring is performed one time a year, generally around September for REVEX and march for mid-REVEX.

The sonder/IIS coregistration error is lower than 100 $\mu$ rad (eq. 100m on ground).

### 4.7.2 IIS / AVHRR co-registration

The IIS/AVHRR co-registration is permanently estimated by the L1 processing chain.

Note that AVHRR channels 4 and 5 are within the IIS spectral filter. The spatial resolution of the IIS (0,7km) is close to AVHRR (1km).

The IIS/AVHRR offset guess in the ground segment configuration is used when the algorithm of correlation between IIS and AVHRR does not converge (typically over homogeneous scenes).

The following figures show a comparison of IIS-AVHRR offsets (GlaOffsetIISAvhrr) mean profiles.

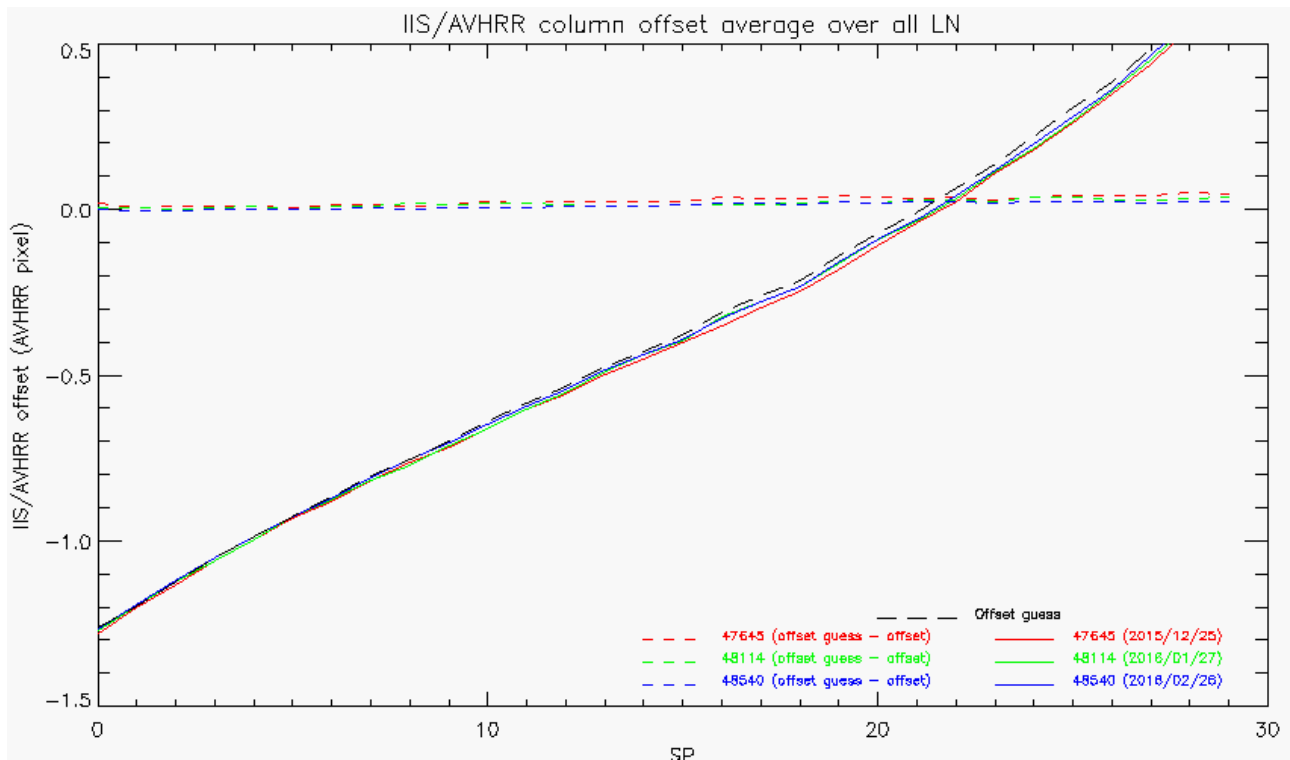


Figure 36 : Column offset (black) guess vs. column offset averaged over all lines (LN) as a function of the scan position (SP=SN), and orbit number



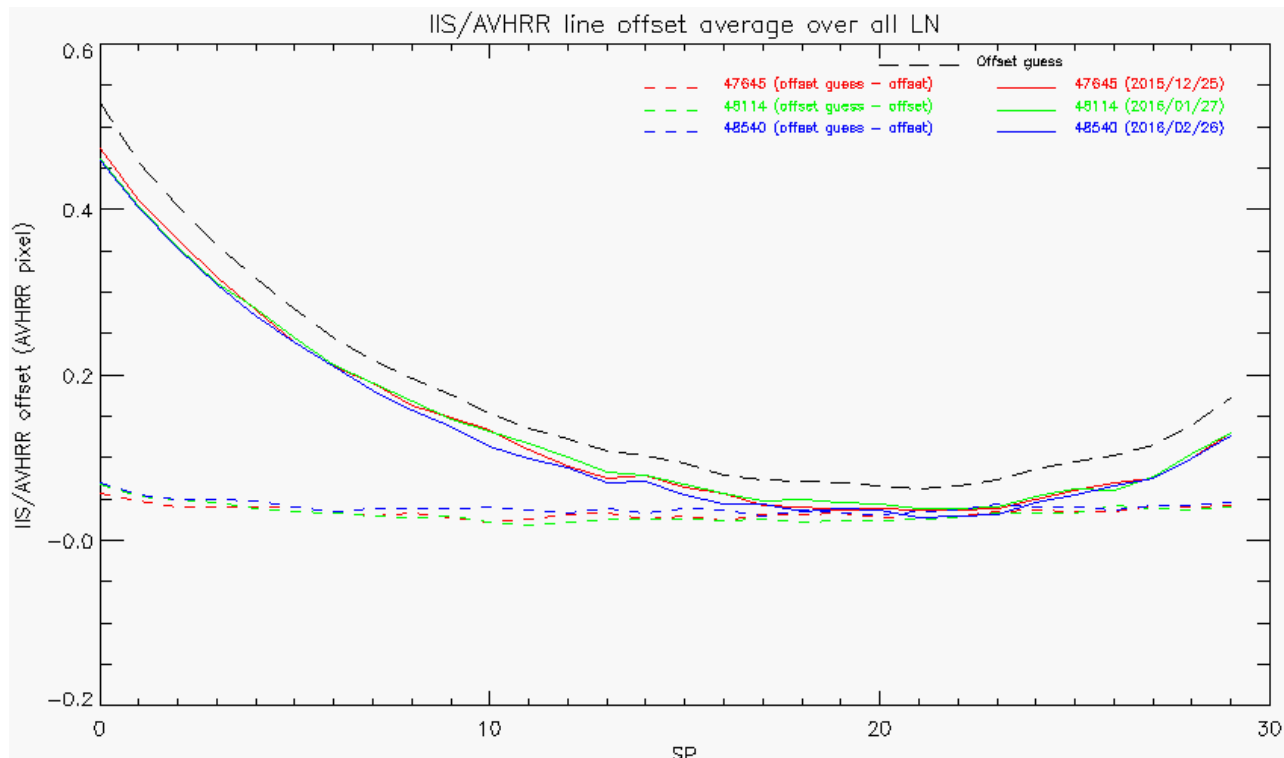


Figure 37 : Line offset guess (black) vs. line offset averaged over all lines (LN) as a function of the scan position (SP=SN), and the orbit number

For both across track and along track, the residuals between measured and IIS/AVHRR offset guess in the ground segment configuration are lower than 0.1 AVHRR pixel for all viewing angles, that is equivalent to 100m on ground.

The values are stable.



### 4.7.3 Conclusion

The positions of IASI pixel are considered stable and well within specification.

IIS-sounder co-registration is stable at about 100 $\mu$ rad which is equivalent to 100m on ground (specification : < 0.8 mrad).

IIS-AVHRR offset is lower than two pixels and stable over time: less than 0.1 AVHRR pixels over three months (specification: < 0.3 AVHRR pixel).

IASI pixel centre location accuracy in AVHRR raster is around 200m. The geolocation of IASI pixels are thus considered stable and well within specification (5 km).

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## 4.8 IASI-A INTER-CALIBRATION WITH CRIS AND AIRS

The objective of the radiometric inter-calibration between IASI-A and CRIS and AIRS is to perform an external monitoring of the IASI performances and to ensure the consistency of the hyperspectral TIR sensors. A radiometric and spectral inter-calibration of IASI-B versus IASI-A is given in the IASI-B quarterly performance report. We aim here at checking the performance of the radiometric absolute calibration accuracy of 0.5K per IASI channel.

The inter-calibration is performed with the SIC software. The methodology is described in the SIC algorithm description [DA.2]. Roughly, the methodology is based on the construction of a database in which each data is the difference IASI-A - CRIS or IASI-A - AIRS over a common observation made by both sounders. “Common” means same place, same time and same viewing conditions so that the only difference is not geophysical but due to a calibration bias. Statistics over this database emphasize the radiometric biases. For IASI and CRIS or AIRS, the database is built with Simultaneous Nadir Overpasses, occurring at high latitudes. Each scene is a regional averaging of the soundings, with a spectral reduction in broad pseudo-bands.

### 4.8.1 IASI-A inter-calibration with CRIS

Figure 38 shows the mean and standard deviation of IASI-A - CRIS for one year of data. We see that IASI-A and CRIS are very well cross-calibrated, with biases lower or equal to 0.1K in absolute value (except for one pseudo-band).

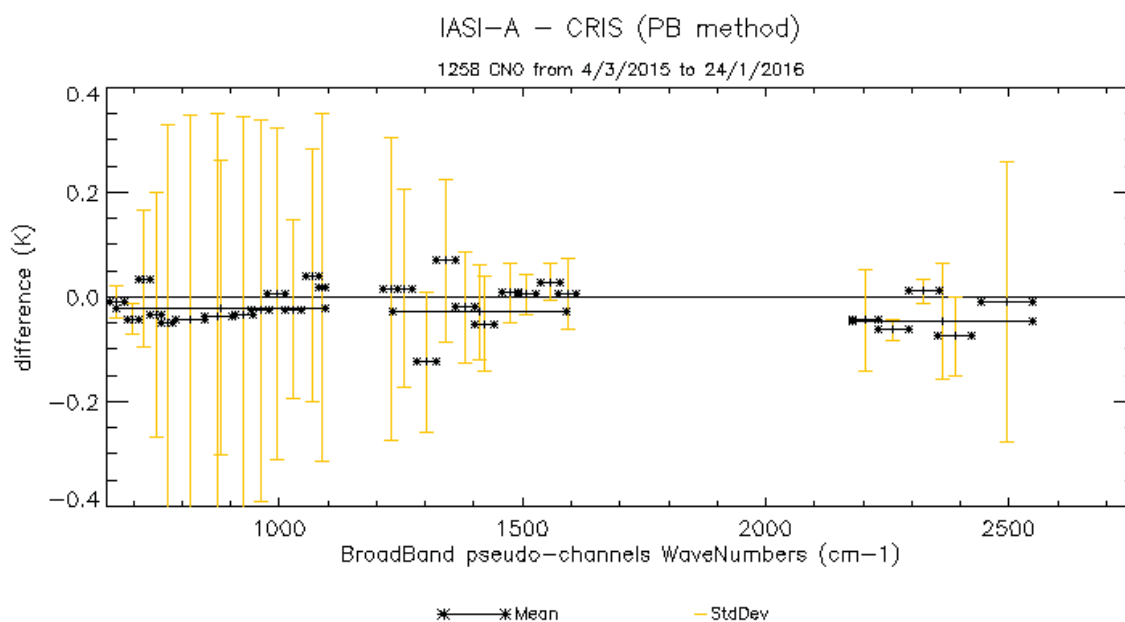


Figure 38 : Mean (black curve) and standard deviation (yellow) of the difference in brightness temperature IASI-A – CRIS

Figure 39 shows the temporal evolution of the three broadest pseudo-channels, corresponding approximately to the IASI bands 1, 2 and 3 (more precisely, to their overlapping with the CRIS coverage). We see that the difference between IASI-A and CRIS is stable with time, with slight variations in IASI B1. Note that some large temporal gaps may be encountered for IASI-A / CRIS, due to the orbital configuration and the tolerance in simultaneity, making some monthly means meaningless.

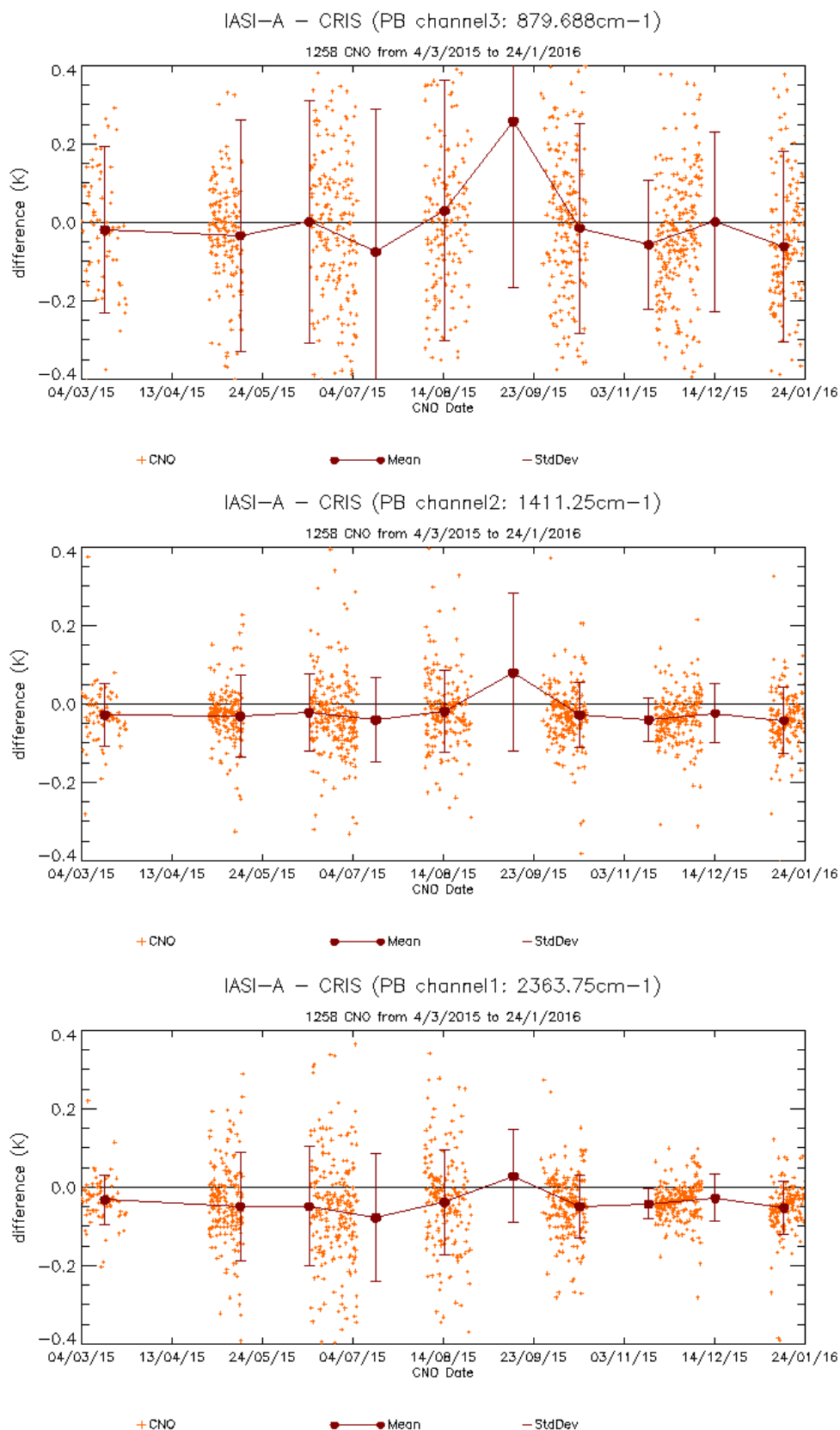




Figure 39 : Temporal evolution of the differences IASI-A – CRIS in brightness temperature over the approximate three IASI bands, with monthly means

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#### 4.8.2 IASI-A inter-calibration with AIRS

Figure 40 shows the mean and standard deviation of IASI-A - AIRS for one year of data. We see that IASI-A and AIRS are very well cross-calibrated, with biases lower or equal to 0.1K in absolute value (except for one pseudo-band).

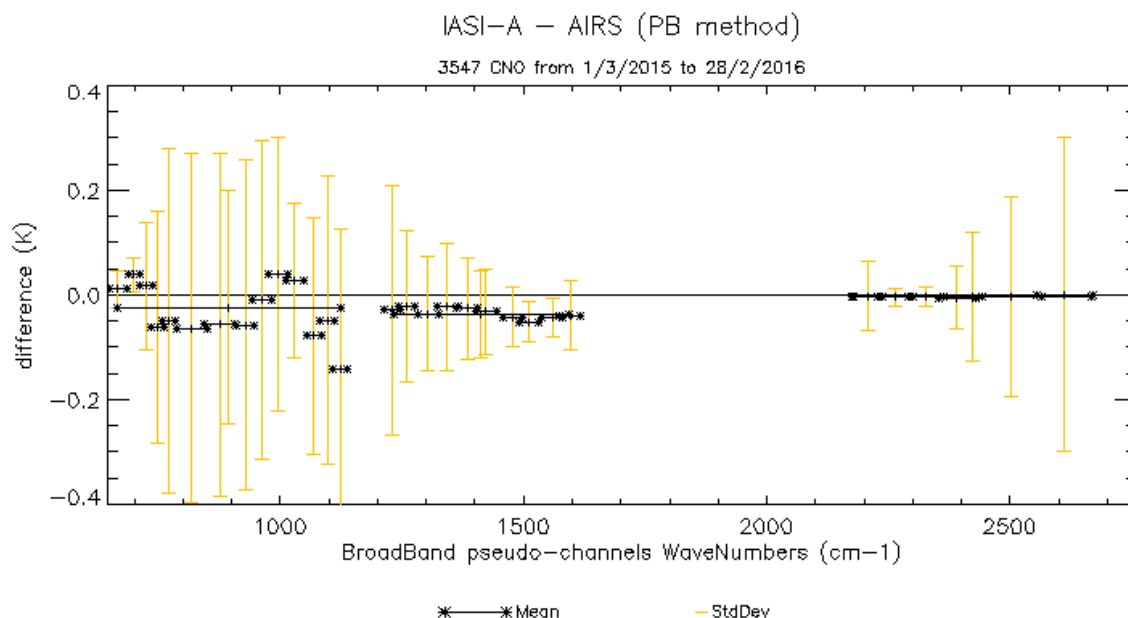


Figure 40 : Mean (black curve) and standard deviation (yellow) of the difference in brightness temperature IASI-A - AIRS.

Figure 41 shows the temporal evolution of the three broadest pseudo-channels, corresponding approximately to the IASI bands 1, 2 and 3 (more precisely, to their overlapping with the AIRS coverage). We see that the difference between IASI-A and AIRS is very stable with time, even more than IASI-A - CRIS. For IASI-A / AIRS no large temporal gaps are observed.

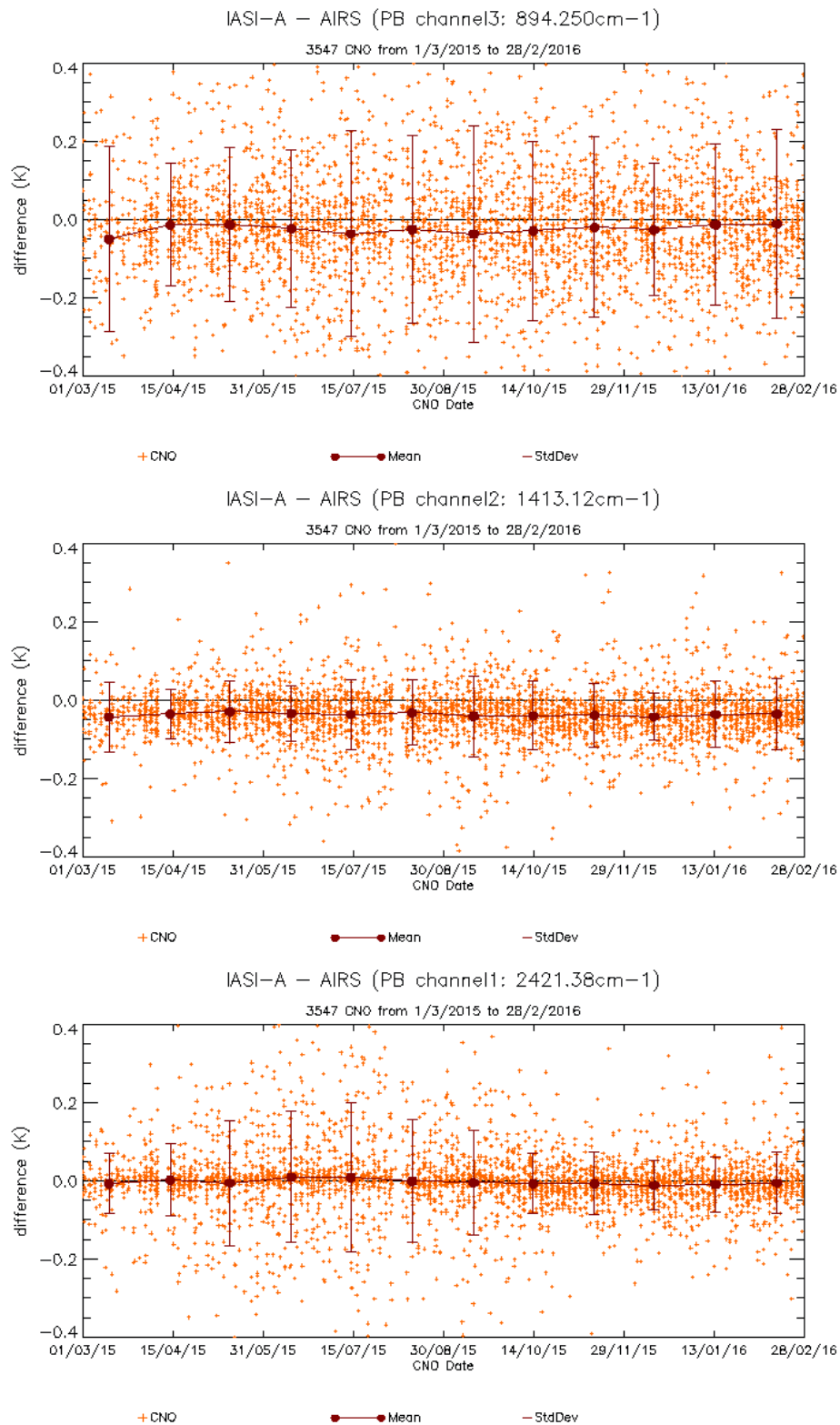




Figure 41 : Temporal evolution of the differences IASI-A – AIRS in brightness temperature over the approximate three IASI bands, with monthly means

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## **4.9 IIS RADIOMETRIC PERFORMANCES**



The main task of IIS is to insure a good relative positioning of IASI sounder pixels with respect to AVHRR. Its performances are studied on one monthly external calibration in three and the monitoring is performed one time a year for REVEX.

### **4.9.1 IIS Radiometric Noise Monitoring**

Refer to REVEX, paragraph 6.9.1.

### **4.9.2 IIS Radiometric Calibration Monitoring**

Refer to REVEX, paragraph 6.9.2.

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## 5 IASI TEC SOFTWARE AND INTERFACES

### 5.1 *IASI TEC EVOLUTION*

A new version “8.6” of the software “TEC” was installed

Table 19 lists previous software evolutions.

IASI TEC software version	implementation	Comments
8.1	6 October 2011	Automatic downloads of L0 products from EUMETSAT FTP
8.2	12 April 2012	New version of product browser (handling IASI L0, L1C products and board configuration).
8.3	22 August 2012	Regularization version before IASI-B CAL/VAL CCAT replaced by CBST in TEC's logs
8.4	19 December 2013	New parameter SP_NV in SLT files Integration of board configuration generation tool (UTOPIE) Integration of LBR products management tool
8.5	19 November 2014	Monitoring of offset voltages from equalization images
8.6	1 February 2016	Detection chain gain and offset optimization Spectral Data Base initialization

*Table 19: IASI TEC at CNES Toulouse*



### 5.2 *SIC EVOLUTION*

No evolution within the period.

This Table lists the recent evolutions of the software:

SIC software version	implementation	Comments
3.2	04/2014	Add of spectral inter-calibration, synthesis tool processed per days, new methods for CNO pre-selection, add of preselection for IASI/AIRS and IASI/CRIS, intercalibration with convolution method for IASI/CRIS
3.3	04/2015	New functionalities for data transfer. Improvement of CNO prediction and pre-selection. New parametrization for the synthesis tool. Automation of the software installation and deployment, IDL upgrade, correction of NCR

*Table 20: SIC at CNES Toulouse*

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### 5.3 EUMETCAST INTERFACE

EUMETCast dissemination is used for Near Real Time data reception by IASI TEC at CNES, Toulouse. Each orbit, L1 ENG, L1 VER, and AVHRR 1B products are received under continuous series of 3 minutes PDU. Full dumps are reconstructed by the EUMETCAST terminal and pushed to a IASI TEC server. Since August 2012, NPP/CrIS PDU are also received to perform inter-comparison with IASI.

In case of failure of the prime EUMETCAST station, products remain available several days on a redundant station.

The behaviour of the EUMETCAST reception is nominal.

The following table lists the recent modifications in the EUMETCAST configuration:

Date	EUMETCAST configuration
29/03/2011	End of IASI L0 dissemination via EUMETCAST
03/08/2011	Hardware and software upgrade of the prime station
04/12/2011	Hardware and software upgrade of the back-up station
13/07/2012	Software patch to correct an anomaly concerning AVHRR files (reception of 0 byte files from EUMETCAST)
24/08/2012	Modification of EUMETCAST configuration to receive NPP/CrIS data
03/2013	“PARALLEL_RECONSTRUCTIONS” set to 3 to avoid missing PDU problems
09/2013	“RECONSTRUCTION TIME-OUT” set to 90 to avoid missing PDU problems
09/12/2014	Antenna repointing and update of reception parameters
21/04/2015	New SR1 router
10/08/2015	Replacement of Back-up station HDD



*Table 21: EUMETCAST configuration at CNES Toulouse*

### 5.4 FTP INTERFACE

Since March 29<sup>th</sup> of 2011, IASI L0 full dumps are available in Near Real Time on a EUMETSAT FTP server. The IASI TEC software automatically downloads products from the EUMETSAT FTP server.

The reception of L0 products at IASI TEC is nominal.



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## 6 **CONCLUSION AND OPERATIONS FORESEEN**

Please visit <https://iasi.cnes.fr> to get IASI news.

### 6.1 ***SUMMARY***

The IASI FM2 instrument is fully operational.

IASI Instrument Management System (IMS) switched to redundant side since 13 April 2015.

The main events are :

- Moon avoidance on 29-30 December 2015
- Heater Refuse on 8 January 2016
- IIS equalisation on 14 January 2016
- IP Manoeuvre on 10 February 2016 (collision avoidance)

### 6.2 ***OPERATIONS FORESEEN***

- IP Manoeuvre on 8 June 2016
- Update of scan mirror reflectivity coefficients end 2016

*End of document*